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# Birth and closure of the Kallipetra Basin: Late Cretaceous reworking of the Jurassic Pelagonian–Axios/Vardar contact (northern Greece)

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**Abstract.** Some 20 Myr after the Late Jurassic to Early Cretaceous obduction and collision at the eastern margin of Adria, the eroded Pelagonia (Adria)-Axios/Vardar (oceanic complex) contact collapsed, forming the Kallipetra Basin, described around the Aliakmon River near Veroia (northern Greece). Clastic and carbonate marine sediments deposited from the early Cenomanian to the end of the Turonian, with abundant olistoliths and slope failures at the base due to active normal faults. The middle part of the series is characterized by red and green pelagic limestones, with a minimal contribution of terrigenous debris. Rudist mounds in the upper part of the basin started forming on the southwestern slope, and their growth competed with a flux of ophiolitic debris, documenting the new fault scarps affecting the Vardar oceanic complex (VOC). Eventually, the basin was closed by overthrusting of the VOC towards the northeast and was buried and heated up to  $\sim$  180 °C. A strong reverse geothermal gradient with temperatures increasing up-section to near 300 °C is recorded beneath the VOC by illite crystallinity and by the crystallization of chlorite during deformation. This syntectonic heat partially reset the zircon fission track ages bracketing the timing of closure just after the deposition of the ophiolitic debris in the Turonian. This study documents the reworking of the Pelagonian-Axios/Vardar contact, with Cenomanian extension and basin widening followed by Turonian compression and basin inversion. Thrusting occurred earlier than previously reported in the literature for the eastern Adria and shows a vergence toward the northeast, at odds with the regional southwest vergence of the whole margin but in accordance to some reports about 50 km north.

#### 1 Introduction

The Hellenides are an integral segment of the main Alpine– Himalayan orogenic belt (Fig. 1). They have recorded polyphase Alpine deformation since the Middle Jurassic, when they were involved in the obduction of imbricate oceanic units over the eastern Apulian margin (e.g., Bernoulli and Laubscher, 1972; Zimmerman and Ross, 1976; Schmid et al., 2020). In the Internal Hellenides (Fig. 1), continuous convergence led to the collision of continental promontories with Eurasia in the Late Jurassic-Early Cretaceous that built a metamorphic crustal-scale orogenic wedge involving the Pelagonian zone and Rhodope (Ricou et al., 1986; Burg et al., 1996; Schenker et al., 2014). In the Late Cretaceous, the metamorphic thrust sheets of the Pelagonian zone were exhumed to shallow depths. This is testified by a cooling below ca. 240 °C from 83 Ma onwards along the northern margin of the Pelagonian zone (Most, 2003; Schenker, 2013), from 54 Ma to the south (Lips et al., 1998, 1999; Coutand et al., 2014), and by the deposition of metamorphic Pelagonian detritus in a Late Cretaceous basin (Kossmat, 1924; Schenker

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et al., 2015), subsequently referred to as the Kallipetra Basin in this study. During the Late Cretaceous–Eocene, thrusting resumed in the Internal Hellenides (Godfriaux et al., 1988; Schermer, 1993) and progressively migrated southward to the External Hellenides (e.g., Aubouin, 1973). Finally, in the Oligocene–Miocene, the Pelagonian zone was dissected by diachronous normal faults (Schermer et al., 1990; Lacassin et al., 2007; Coutand et al., 2014; Schenker et al., 2014) within a southward extensional deformation front that affected most of the Hellenides (e.g., Papanikalaou and Royden, 2007).

In the Pelagonian zone and adjacent units, the record of this orogenic system in the time interval between collision in the Early Cretaceous and resumed thrusting in the Late Cretaceous-Early Cenozoic remains sparse, and the discontinuous, sometimes contrasting, large-scale interpretations stem from the difficulties in establishing a coherent tectonic history across the Rhodope and the Pelagonian zone (Fig. 2). To elucidate part of these controversies, this study investigates the small Upper Cretaceous Kallipetra Basin that formed on both oceanic and continental units along the eastern Pelagonian margin (Fig. 1) and was overthrusted by serpentines of a Jurassic oceanic floor (the ophiolitic fragments lying on the Pelagonian zone named Axios/Vardar/Almopias zone by Schenker et al., 2015). According to the scenarios proposed in Fig. 2, the Kallipetra Basin may have formed (i) within a long-lived Jurassic-Cretaceous passive margin characterized by the income of flysch from the approaching Rhodopian trench to the north-northeast (e.g., Papanikolaou, 1989; Ricou et al., 1998; Papanikolaou, 2009); (ii) during an extensional tectonic event in between the Pindos obducting from the south southwest and the Axios/Vardar/Almopias zone subducting to the north–northeast (Sharp and Robertson, 2006); (iii) over the obducted Axios/Vardar/Almopias zone (Froitzheim et al., 2014); or (iv) within a collisional wedge that incorporated the obducted Axios/Vardar/Almopias zone (Schenker et al., 2014).

The Kallipetra Basin collected coarse detritus – metamorphic and ophiolitic rocks from its shoulders – followed by the deformation of its deposits, by thermal conditions that locally partially or totally reset the cooling ages, and by cooling during the Late Cretaceous (Schenker et al., 2015). However, the stratigraphic evolution and the depositional age of this basin are so far only partially constrained. Moreover, it remains unclear how thermal conditions (temperatures > ca. 240 °C) and deformation in this basin relate to the apparent tectonic quiescence associated with extensive Late Cretaceous cooling recorded elsewhere in the Pelagonian zone (Schenker et al., 2015) and to the diachronous and complex tectonic evolution of the Hellenides. Finally, the timing of opening and sealing of the basin and the tectonic environment of deposition can provide fundamental constraints to unravel the Late Cretaceous interval of the long history of accretion, subduction, arc magmatism, and large-scale extension in the Hellenic subduction system (e.g., Jolivet and Brun, 2010; Ring et al., 2010; Burg, 2012; and references therein), therefore allowing us to better define the geodynamic models proposed so far.

This study uses conventional geological mapping techniques, stratigraphic analysis, illite crystallinity, and lowtemperature thermochronology to obtain new constraints on the tectonic evolution of the eastern margin of the Pelagonian zone and to unravel the Late Cretaceous detrital record. Our data indicate that the Upper Cretaceous Kallipetra Basin was relatively shallow and tectonically active as testified by the presence of olistoliths, large gravitational features such as rotational growth faults and slumping, and early diagenetic deformation. Rudist bioherms were accumulated on the slopes of the basin with flank deposits dipping into the basin. The bioherms were terminated through environmental restriction or burial due to increased serpentinite sediment input from the south-southwest-eroding oceanic complex. Moreover, based on illite and petrographic data, we find an inverted, high, non-linear geothermal gradient related to a heating event, which likely occurred during the overriding of the Vardar oceanic complex (VOC) in the Late Cretaceous.

### 2 Background

## 2.1 Large-scale tectonic setting

Following the Variscan Orogeny and Permian strike-slip and extension (Schenker et al., 2018), the Permian-Triassic rifting led to the creation of the Tethys and its seaways, namely the Pindos, Vardar/Maliac, and Meliata basins, which continued to open during the Triassic to Early Jurassic (e.g., Bernoulli and Laubscher, 1972; Schmid et al., 2008; Papanikolaou, 2009). The convergent motion between Eurasia and Adria led to a northward intra-oceanic subduction in the Vardar in the Early-Middle Jurassic that saw the production of magmatic arcs to the north (Dimitrijevic, 1982; Bortolotti et al., 1996; Burg, 2012). In the Late Jurassic, there was southwestward obduction of the Tethys ophiolite from the Vardar Ocean onto the passive continental margin of the Pelagonian zone to the south (Bernoulli and Laubscher, 1972; Dimo-Lahitte et al., 2001). Jurassic-to-Lower-Cretaceous sediments were imbricated during the accretion of the ophiolitic units (e.g., Robertson and Dixon, 1984; Bortolotti et al., 2005; the complex named Axios/Vardar/Almopias zone in Schenker et al., 2015). Continuous crustal shortening caused the accretion of Rhodope by the latest Jurassic-Early Cretaceous and of the Pelagonian zone by the Early Cretaceous (Figs. 1 and 2; Burg et al., 1996; Ricou et al., 1998; Schenker et al., 2014; Moulas et al., 2017). The buried Pelagonian basement experienced regional amphibolitic-facies metamorphism to the north (U-Pb zircon ages from the leucosomes of migmatites at 130–117 Ma; Schenker et al., 2015, 2018) and an upper greenschist- to blueschist-facies metamorphism to the south (Ar-Ar ages on

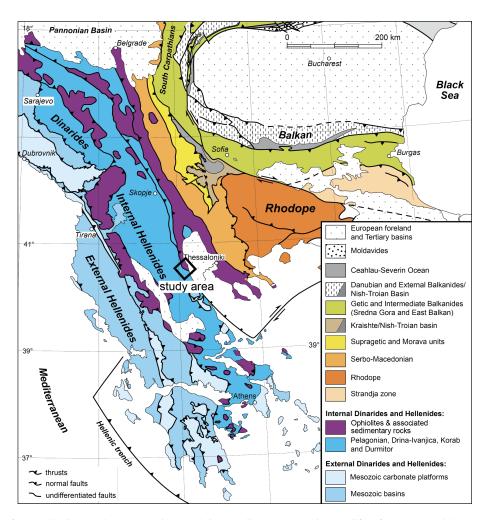


Figure 1. Location of the Hellenides and study area in the Alpine Mediterranean chain. Modified from Burg (2012).

muscovite at 100-85 Ma; Schermer et al., 1990; Lips et al., 1998).

The tectono-sedimentary history during the Early Cretaceous is highly debated. Sharp and Robertson (2006) suggest that the Pelagonian zone and its emplaced ophiolitic rocks underwent extensional exhumation already during the Late Jurassic. Rather, Early Cretaceous tectonic activity has been recorded near Edhessa (Fig. 1) in the Axios/Vardar/Almopias zone, along with late Aptian–Early Albian transgression on both the Pelagonian platform and the Axios/Vardar/Almopias zone (Mercier, 1968; Mercier and Vergely, 2002). To the south, the Aptian–Albian time was characterized by a sedimentary hiatus (~120–100 Ma) over lower Aptian deformed flysch and bauxitic laterites instead (Nirta et al., 2015, 2018), suggesting a growing topography in the frontal part of the orogenic wedge.

Thereafter, transgressive Cenomanian-to-lower Campanian limestones and deep-water Paleocene turbidites unconformably overlay the eroded Pelagonian and Axios/Vardar/Almopias imbricated units (Mercier, 1968;

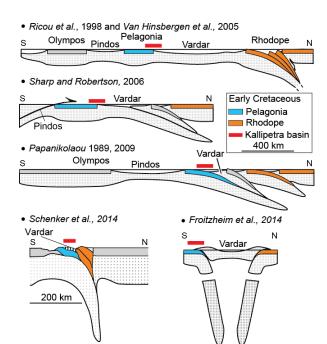
Mercier and Vergely, 2002; Papanikolaou, 2009) attesting deepening below sea level of the Rhodope-Pelagonian crustal-scale orogenic wedge. Moreover, during the Late Cretaceous to Eocene and locally since the Campanian, the imbrication of the Axios/Vardar/Almopias units resumed at several locations in relation to thrusting with vergence to the NE and to the SW. This has been documented in the central-eastern Vardar near the study area (Paikon Window; Godfriaux and Ricou, 1991; Bonneau et al., 1994; Brown and Robertson, 2003; Katrivanos et al., 2013), in the northwestern Vardar (Grubić et al., 2009; Ustaszewski et al., 2009), in the northeastern Pelagonian zone (Kilias et al., 2010), in the southern Pelagonian zone (Baumgartner, 1985), and in the Pindos zone (e.g., Aubouin, 1959, 1973; Papanikolaou, 1997). Continuous convergence up to the Neogene progressively deformed the continental margin of the Adriatic plate into southwest-verging fold and thrust sheets (Fig. 1; Channell and Hovarth, 1976). Final exhumation of the stacked crustal and oceanic slices occurred through extensional metamorphic domes between the Eocene in the north and late Neogene in the south (e.g., Lister et al., 1984; Dinter and Royden, 1993; Gautier et al., 1993, 1999; Brun and Sokoutis, 2007; Jolivet and Brun, 2010; Burg, 2012).

# 2.2 Main geologic features of the eastern Pelagonian margin

The Pelagonian basement consists of deformed (i) orthogneisses crosscut by leucogneisses and leucogranites; (ii) mafic amphibolite bodies; and (iii) interlayered marbles (Schenker, 2013, and references therein). Cooling of the Pelagonian core complex carapace rocks may have started at or after collisional doming at 118±4 Ma (U–Pb metamorphic zircon ages; Schenker et al., 2018). <sup>40</sup>Ar/<sup>39</sup>Ar white mica ages from the Pelagonian gneisses show a younging toward the dome core from 111–100 to 80–64 Ma that witness the slow exhumation and cooling of the deeper units of the basement (Schenker, 2013).

The Axios/Vardar/Almopias unit includes a mélange zone made of tectonically superimposed marbles, serpentinites (ophicalcites), flysch-phyllitic series, volcanoclastic sediments, amphibolites, and carbonatic sequences imbricated southwestward during the Late Jurassic obduction over the Pelagonian zone (e.g., Smith et al., 1975; Ricou and Godfriaux, 1995; Sharp and Robertson, 2006; Ferriere et al., 2016). In the study area, the Axios/Vardar/Almopias unit is represented by serpentinites that are referred to as the VOC, which consists of at least five lithologies: (i) ophicalcites; (ii) dark massive fractured to brecciated serpentinites; (iii) sedimentary serpentinite breccias; (iv) sedimentary serpentinite breccia with platform carbonate clasts; and (v) foliated serpentinites and limestones. Ferromanganese-rich chert nodules within the VOC, dated further to the south at approximately 175 Ma by Chiari et al. (2013), attest the involvement of this Jurassic oceanic floor in the intra-oceanic Tethys subduction and subsequent obduction.

On the eastern margin of the Pelagonian zone, relatively thick packages of Upper Cretaceous carbonate and siliciclastic sediments with both a Pelagonian and ophiolitic provenance unconformably cover the VOC and the Pelagonian basement (Sharp and Robertson, 2006; Papanikolaou, 2009; Schenker, 2013; Schenker et al., 2015). In the study area, the sediments belong to a sedimentary basin that here is referred to as the Kallipetra Basin (Fig. 3). It formed as an elongated NNW-SSE-oriented belt overlying the VOC and the Pelagonian continent. In this basin, the presence of reworked Lower Cretaceous orbitolinids, Globotruncana sp., and Turonian Helvetoglobotruncana helvetica indicates deposition during the Cretaceous (Schenker et al., 2015). Based on these depositional ages, a zircon fission track (ZFT) age of 67 Ma from an orthogneiss boulder (sample 10-128) at the top of the Kallipetra Basin was previously interpreted as indicating a very short lag time between cooling of the dome and deposition in the basin (Schenker et al., 2015). Two more



**Figure 2.** Geodynamic interpretations of the Hellenides in the Early Cretaceous according to different authors. Note that there is no consensus on the Early Cretaceous geodynamic framework at the onset of the Kallipetra Basin between the Pelagonian zone and the Vardar domain.

samples (10-029 and 10-130) from the Kallipetra Basin were interpreted as possibly partially to non-annealed (Schenker et al., 2015).

From the early Oligocene, an overall southwestward tectonic denudation from shallow depths is documented in the Kallipetra Basin by AFT ages of 32.7 Ma (sample 10-128) and in the Pelagonian basement by ZFT ages of 24–20.7 Ma and AFT ages between 22.9 and 16.1 Ma (Schenker et al., 2015).

#### 3 Methods

## 3.1 Geologic mapping and stratigraphy

Geological mapping and structural analysis were conducted to reconstruct the geometry of the basin and the ductile and brittle deformation that affected the Kallipetra Basin and the overlying VOC. The paleogeography, depositional environment, and age of the sedimentary sequences were determined based on stratigraphic logging, optical microscopy, and biostratigraphy. Planktonic foraminifera and nannoplankton were used to establish ages of the sedimentary succession. Simple smear slides were produced using standard techniques to retain the nannofossil assemblages and original sediment composition. Quantitative analyses were carried out using a polarizing light microscope at a magnification of 1250×.

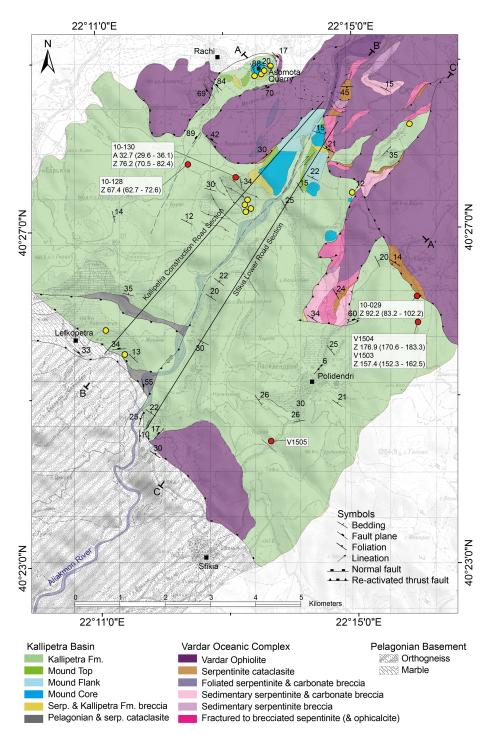


Figure 3. Geological map of the study area. Yellow circles indicate locations of illite samples; red circles indicate location of ZFT samples and their respective ages.

#### 3.2 Illite crystallinity

The Kübler index of illite crystallinity is a method used to determine the diagenetic grade in metapelitic sequences by measuring the changes in shape of the first dioctahedral illite–muscovite basal reflection at a 10 Å X-ray diffrac-

tion (XRD) spacing (Kübler and Jaboyedoff, 2000). To analyze illite crystallinity, bulk-rock mineralogy was obtained through the conventional powder XRD method using the ARL Thermo X'tra powder diffractometer at the University of Lausanne. Samples were then de-carbonated, followed by the extraction of  $< 2 \, \mu m$  clay fraction and 2–16  $\mu m$  fraction

that were used for further analysis. Oriented samples were prepared by sedimentation on a glass slide from the suspended fraction. Samples were first air-dried (AD), and then treated with ethylene glycol (EG) to recognize any overlapping effect of smectite peaks. XRD diffractograms were performed on both the AD and EG treatments. The full width at half-maximum height (FWHM) of the illite 10 Å XRD peak that is measured on both AD and EG clay samples (< 2 µm size fraction) gives the Kübler index (KI) (Kübler and Jaboyedoff, 2000). The KI is expressed as  $\Delta^{\circ}2\theta$  CuK $\alpha$ . The air-dried KI value is used for the determination of low-grade metamorphic zones and approximate temperatures. It should be noted that the KI does not serve as a precise geothermometer but provides a qualitative indicator of stages that phyllosilicates may have reached through metastable mineral reactions (Merriman and Peacor, 1998; Abad, 2007).

Significant asymmetrical peak broadening, caused by a tail in the  $10\,\text{Å}$  peak and produced by the presence of smectite and expandable mixed layers, is reduced following EG treatment (Abad, 2007). These peaks may indicate the presence of detrital illite, which gradually decreases with burial and essentially disappears in the anchizone (Kübler and Jaboyedoff, 2000). The decrease in KI values with increasing metamorphic conditions and temperatures is a consequence of the increase in the number of layers and the disappearance of expanding layers. The Neuchâtel IC scale was calibrated with the Lausanne diffractometer and therefore produced anchizone limits of 0.18 and 0.36°  $\Delta 2\theta$  CuK $\alpha$ , which we use in this study (Jaboyedoff et al., 2000).

#### 3.3 Zircon fission track dating

Two of five collected samples provided enough zircons to date using ZFT analysis: V1503 and V1504. These samples integrate our previous samples (10-128, 10-129, 10-130; Schenker et al., 2015). The new samples were collected with the aim of revealing the full age distribution, which in our previous samples was limited by the low number of available zircons. To this goal, the new samples were > 5 kg each. All the samples consist of arenites and conglomerates. Zircons were separated from the whole rock by initial SELFRAG fragmentation, followed by density-based liquid separation using a Wilfley water table and heavy-liquid separation. The heavy fraction was passed through the Frantz magnetic separator stepwise to remove magnetic minerals from the zircons. Zircons were embedded in PFA (perfluoroalkoxy) Teflon and the prepared mounts were polished to expose the smooth internal zircon surfaces. The polished mounts were etched using a eutectic mix of NaOH and KOH to preferentially damage the fission tracks, enabling them to be fully revealed for optical analysis. To reveal the whole age distribution, we prepared up to four mounts per sample that we etched at very short time steps of 3.5 h. Fully etched zircons were first recognized after 10.5 h and then we etched the remaining mounts to 14 and 17.5 h.

#### 4 Results

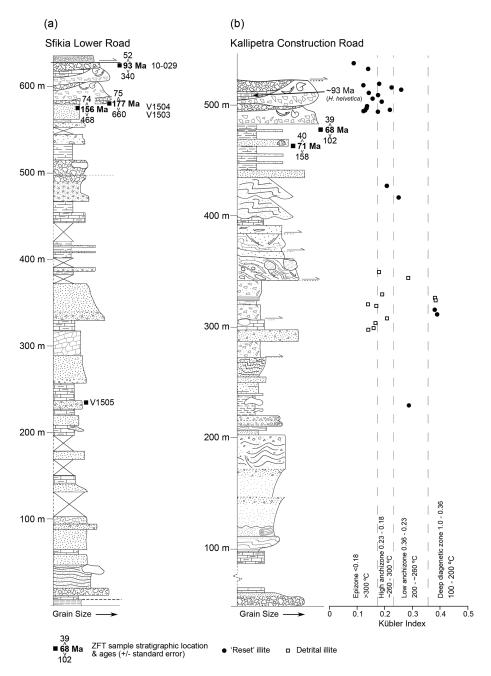
#### 4.1 The Kallipetra Formation: facies and boundaries

The study area is divided into three units: (1) the Pelagonian basement; (2) a stratigraphic unit that we name the Kallipetra Formation, described here for the first time; and (3) the VOC. The Kallipetra Formation is the focus of this study and consists of several lithofacies that collectively characterize a sedimentary basin (Fig. 3). Most field data were collected along two composite stratigraphic sections (the Kallipetra and Sfikia sections, Fig. 4).

The base of the basin is exposed close to the contact with the Pelagonian basement. Locally, the latter consists of a thick package of white, foliated cataclasite containing both serpentinite and gneiss fragments (Fig. 3). Directly overlying the cataclasite is a very dark massive fractured to brecciated serpentinite, followed by pebbly sandstones and wellbedded dark gray limestones. Elsewhere, the base of the basin is characterized by a thick package of serpentinite-rich conglomerates, breccias, and minor amounts of dark gray limestone (Fig. 4a). The conglomerate is clast-supported and poorly sorted, with a dominance of sub-rounded to rounded clasts greater than 15 cm. The conglomerate is composed of dark green to black serpentinite ( $\sim 90\%$ ), dark gray limestone, marble, and orthogneiss clasts, and a fine-grained serpentinite matrix. Thickly bedded, poorly sorted calc arenites stratigraphically overlie the serpentinite conglomerate. These are openly folded on the meter scale and bedding is deformed around large olistoliths of dark gray veined marble and serpentinite. The occurrence of olistoliths decreases significantly up-section (Fig. 4). In the southeast, the basal contact is sharp and consists mostly of marls, shales, and subordinate calc arenites of the same kind as those observed in the central part of the basin, which are described below (Fig. 4b).

Calc arenites are observed throughout the basin, typically at intermediate stratigraphic levels (Fig. 4b). The arenites range from fine- to coarse-grained, are medium to thickly bedded, and often display slumping folds. Locally, these folds and the synsedimentary gravity faults show a top-tothe-NE vergence. The quartz content varies with location, with the highest proportion of quartz being in the northwest region of the study area. Locally, the calc arenites consist of medium- to coarse-grained poorly sorted pebbly sandstones with 1-6 cm sized clasts of red arenite and red-pink carbonate. Very distinctive thinly bedded and laminated red and green marly limestones occur at intermediate-to-high stratigraphic levels (Fig. 4b). The red layers range from 2-5 cm thick, and green layers typically range from 0.5–2 cm thick. These deposits represent the deepest pelagic facies of the basin.

Towards the top of the basin, massively bedded conglomerates and breccias are often interbedded with the calc arenites and pebbly sandstones and consist of limestone, bioclastic limestone, arenite, marl, serpentinite, mudstone, and cal-



**Figure 4.** (a) Stratigraphic column taken along the lower road leading to Sfikia, located directly south of the Aliakmon River, and stratigraphic locations and ages of ZFT samples. (b) Stratigraphic column taken along the Kallipetra Monastery construction road, located north of the Aliakmon River, the ages and stratigraphic locations of ZFT samples, and illite crystallinity samples. Illite crystallinity samples are plotted against Kübler index and diagenetic zone.

careous mudstone as rounded to sub-rounded clasts in a calcareous matrix.

Lateral variations in facies occur frequently, the most evident being the changes observed from the northwestern to the central and southeastern portions of the mapping area. In the northwest (Fig. 4b), the stratigraphy is dominated by coarse to pebbly sandstones, breccias, and conglomerates, whereas shaley limestones, marls, and mudstones prevail in

the southeast (Fig. 4a). In the northwest, lithic fragments of quartz, gneiss, and marble are major components of the coarse sediments, with quartz content ranging from 45 % at the base to 90 % up-section (310 m, Fig. 4b), where serpentinite forms a minor component. In addition, olistoliths and evidence of slumping are frequent at high stratigraphic levels in the northwestern sector (Fig. 4b). This differs greatly from the southeastern sector (Fig. 4a), whereby slumped calc

arenites with olistoliths appear only at the base of the section and the average quartz content is lower.

The top of the Kallipetra Basin is marked by the occurrence of rudist mounds, five of which, some tens of meters thick, were identified in the study area. The mounds produce prominent cliffs and dome-like structures in the topography. Each mound can be separated into four different facies associations (Fig. 5): (i) the serpentinite and Kallipetra carbonate breccia (SKB); (ii) the mound core; (iii) mound flank; and (iv) the mound top.

The SKB is a sub-angular, moderately sorted, clast-supported breccia that is poorly bedded and massive. Clasts comprise serpentinite, dark gray limestone, rudist-rich microsparite, pink micrite, and minor lithic fragments like quartz, feldspar, and some dark pyroxenes. The rudist microsparite and pink micrite clasts are identical to the mound core. The matrix is composed of a fine- to medium-grained calcareous arenite. Orbitolinids were discovered in a clast by Schenker (2013). The SKB is usually found on the southern side or stratigraphically below the mound.

The mound core is characterized by light gray to pink, massively bedded micrite and microsparite, in which float abundant whole rudists. *Hippurites* and radiolitid rudists are present along with encrusting sponges and echinoderm fragments. Rudists are scattered throughout the mound and seem to have no preferred orientation. Vertical calcite veins and en échelon veins are frequently observed at the margins of the mound core.

The mound flank is a heterogeneous lithology that varies with distance from the mound core and location within the basin. In general, a moderately sorted, clast-supported breccia containing large, angular clasts of rudists, red limestone, greenish marls, micrite, and minor serpentinite clasts occurs closest to and on the northern side of the mound core. The number of clasts decreases into a matrix-supported breccia with a marly, green-colored matrix. The serpentinite content gradually increases up-section, and gravel sandstones contain > 60 % serpentinite in addition to red microsparite clasts and rudists from the mound core. A sharp sub-vertical boundary often separates the mound flank and the mound core. The mound flank facies differ slightly throughout the area depending on the location of the mound core. The flank of the northernmost and youngest mound is first characterized by a massive clast-supported breccia consisting of micrite, rudists, sponges, and echinoderm fragments, dissected by neptunian dikes, with onlapping red pelagic marls. Differential compaction structures (load casts and fluid escape features) can be observed in the pelagic sediments where a stratigraphically higher mound core overlies them. Secondly, the clast-supported breccia passes rapidly into a  $\sim$  34 m thick sequence of pelagic marly limestones only seen on top of the northernmost mound. The proportion of marls within the mound flank gradually increases from the southerly mounds to the northernmost mound.

Stacking of serpentinite-rich breccias always occurs on the southern slope of the rudist mounds. Flank deposits, either marls or a succession of sandstones and breccias, dip away from the mound core always on the northern mound side.

The mound top, where observed fully, is approximately 6 m thick and stratigraphically overlies the mound core (Fig. 5). It generally consists of several meters of very poorly sorted, angular to sub-angular gravel of serpentinite and quartz within a white calcareous matrix. A thin layer of rudist-rich, elongated carbonate clasts overlies the gravel. There is a gradual transition into a clast-supported conglomerate with a reddish calcareous matrix, plus arenite and minor serpentinite clasts. The clasts of this conglomerate are very deformed, where the VOC tectonically overlies them. The full stratigraphy of the mound top was only observed at one, the southernmost, rudist mound (Fig. 5).

#### 4.2 Biostratigraphic data

Although significant amounts of sample were collected for biostratigraphic analysis, nearly all of them were barren, or included dissolved, silicified, or recrystallized nannoplankton and foraminifera, making most species indistinguishable. Table 1 summarizes the recognizable planktonic foraminifera that were only found near the northernmost mound (Asomata Ouarry).

The orbitolinid found in sample M2-TS3, Mesorbitolina pervia (Annie Arnaud-Vanneau, personal communication 2020), has an older stratigraphic distribution than most of the ages displayed in Table 1, with the minimum age being in the basal late Aptian. However, the fossils are deformed and found in a clast together with other indicators of shallowwater platform conditions, suggesting that they have been reworked and are supplied from elsewhere (Fig. S1 in the Supplement). Indeed, Schenker (2013) discovered Lower Cretaceous orbitolinids in the VOC, located very close to the tectonic contact with the Kallipetra Basin. Late-Jurassic to Lower Cretaceous limestones directly overlying the Pelagonian basement and the dismembered, eroded ophiolites are the probable source of these fossils (e.g., Brown and Robertson, 2004; Sharp and Robertson, 2006). Therefore, this sample is excluded from discussions about the depositional age of the Kallipetra Basin.

Species abundance and totals of calcareous nannofossil were semi-quantitatively evaluated as F (frequent) and R (rare). In the studied section (M2), the major calcareous nannofossil events in stratigraphic order are as follows: the first occurrence of *Quadrum gartneri*, *Eprolithus octopetalus*, and *Eprolithus eptapetalus* (sample M2/2, 50 cm from the bottom of the section); the first occurrence of *Eiffellithus eximius*; and first and last occurrence of *Kamptnerius magnificus* (sample M2/6B, 2 m from the bottom of the section) (Fig. S2).

Q. gartneri, E. octopetalus, and E. eptapetalus can be correlated with the UC7 zone in the Turonian stage, giving the

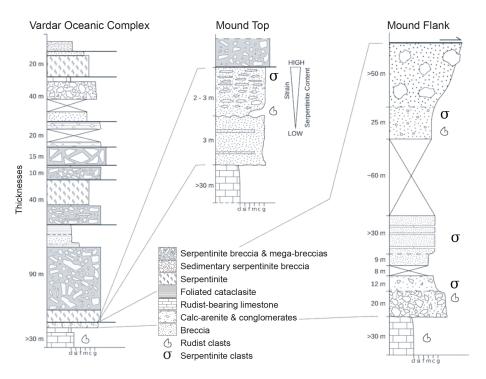


Figure 5. Stratigraphic sections of the Vardar ophiolitic complex, mound top, and mound flank.

Table 1. Table of observed foraminifera.

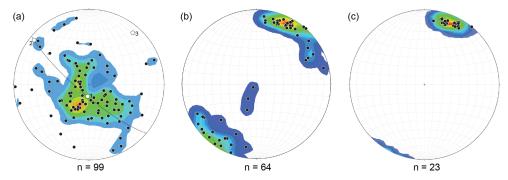
Sample	Location	Foraminifera	Stratigraphic distribution	Age
M2-TS1 M2-TS1 M2-TS3 M2-TS2 M2-TS3	40°28′53″ N, 22°13′38″ E 40°28′53″ N, 22°13′38″ E 40°28′53″ N, 22°13′39″ E 40°28′55″ N, 22°13′42″ E 40°28′55″ N, 22°13′42″ E	Helvetoglobotruncana helvetica Dicarinella hagni (?) Whiteinella sp. Whiteinella (inornata?) Mesorbitolina pervia	Lower-middle Turonian Lower Turonian-Coniacian Upper Cenomanian-Campanian Cenomanian-Turonian boundary-Santonian Mid-upper Aptian	93.5–92.7 Ma 93.52–86.71 Ma 100.5–72.05 Ma 94.03–84.19 Ma

base of the M2 section a minimum age of 93.6 Ma (Burnett et al., 1998). *E. eximius* and *K. magnificus* can be correlated with the base of the UC8 zone in the Turonian stage.

### 4.3 Post-sedimentary structural data

In the marls and marly limestones of the Kallipetra Basin, the foliation is mostly parallel to the bedding and defined by flat and elongated quartz clasts and clay minerals. Bedding and foliation dip at a very low angle either to the NW or to the SE due to bending around an axis plunging shallowly to the NE (Fig. 6a, b; Table S3 in the Supplement). Stretching lineations are observed mostly on lamination surfaces in fine-grained marls, limestones, and mudstones, and they are formed by the alignment of elongated clay minerals. Mineral lineations occur mainly in the Pelagonian basement where the long axis of elongated quartz and feldspar crystals are aligned. In all the lithologies the lineations strike NE–SW at low dip angles (< 20°; Fig. 6b). Stretching lineations along with asymmetric interlayered boudinaged beds indicates a top-to-the-NE shear sense within the basin (Fig. 7a).

The top of the Kallipetra Basin is tectonically covered by the VOC. In the profiles A-A' and B-B' (subparallel to the NE–SW lineations, Fig. 8), the contact appears preferentially flat and dips with shallow angle ( $< 15^{\circ}$ ) to NE. Along the profile C-C' (orthogonal to the lineations, Fig. 8), the contact is bent over the mound cores and flanks, forming an undulate surface. To the NE, the Kallipetra sediments overlie the VOC-forming tectonic duplexes (Profile B-B', Fig. 8). The shear zone in the footwall of the VOC is characterized by 2 to 6 m thick foliated cataclasites and by a strain gradient visible through the increase in the intensity of the foliation. The cataclasite is usually white in color, clay-rich, and often features floating carbonate and/or serpentinite blocks in the matrix. Conglomerates below the contact between the VOC and rudist mounds show a 3 m thick strain gradient from almost non-deformed clasts at the bottom to cigar-shaped and highly elongated clasts at the contact (Fig. 7b; Profile B-B', Fig. 8). The orientation of the long axis of the cigars is subparallel to the stretching lineations observed throughout the study area suggesting that the deformation during the tec-



**Figure 6.** Lower hemisphere stereoplots of: (a) foliation poles whereby measurements were taken from marls and limestones; (b) mineral and stretching lineation measurements of the mapped area; and (c) stretching lineations of the strained conglomerate at grid reference  $40^{\circ}27'23''$  N  $22^{\circ}15'00''$  E. These measurements do not include foliations observed in foliated cataclasites.

tonic emplacement of the VOC was penetrative within the basin. Shear bands, stepover structures, and sigma clasts in the cataclasite indicate a top-to-the-NE tectonic movement (Fig. 7c) synthetic to the intra-basin shearing. However, new growth of syntectonic chlorite (Fig. 9) along the main contact shows that shearing below the VOC occurred at higher thermal conditions with respect to the intra-basin deformation.

Late normal faults trending W–E to NW–SE and two transtensional to strike-slip faults crosscut the Pelagonian basement, the Kallipetra Basin, and the VOC (Fig. 3). Most steep normal faults plunge to the NE (Fig. 8). Low-angle fault zones are observed within the VOC, dipping approximately 35° towards the NE with a normal top-to-NE shear sense. This later extensional phase may have locally reactivated the major tectonic contact. The dextral strike-slip component of the fault along the valley of the Aliakmon River reached approximately 50 m in the south and just a few meters in the north. Another transtensional fault causes a small normal displacement of approximately 50 m that uplifts the northernmost mound core.

## 4.4 Illite crystallinity data

Thirty-seven samples for illite crystallinity analysis were taken up-section in the north part of the Kallipetra Basin (Kallipetra section). Four samples were unsuitable for illite crystallinity analysis as the  $< 2 \, \mu m$  portions contained no illite. The remaining samples have a KI ranging from 0.091 to 0.39 (Fig. 4b; Table 2).

Stratigraphically higher samples have a KI ranging from 0.09 to 0.25, and stratigraphically lower samples have a KI of 0.39. The KI appears to increase down-section for samples containing only non-detrital illite. The sample with the lowest KI of 0.091 is characterized by an XRD pattern that reveals the presence of chlorite. The sample with a KI of 0.14 contains paragonite which indicates epizone conditions.

The samples containing detrital illite are limited to stratigraphic heights between 300 and 350 ms and show a large range of KI between 0.14 to 0.383. Non-detrital illite, on the contrary, is mostly confined to stratigraphic heights above 400 m. Using the anchizone limits as calibrated for our lab (see Sect. 3.2; Jaboyedoff et al., 2000) and given the fact that the effects of detrital micas disappear in the anchizone ( $\sim$  200–300 °C), the results indicate that the Kallipetra sediments experienced higher temperatures (lower KI values) closest to the tectonic contact with the VOC. This is supported by the presence of paragonite at the top of the section. The KI values subsequently increase away from the contact, indicative of an inverse geothermal gradient from > 300 to 100–200 °C within  $\sim$  165 m (Fig. 4b).

## 4.5 Zircon fission track

We collected our samples along a down-section direction within the Kallipetra Basin: the only ones that produced enough countable zircons are from close to the contact with the VOC. Results are reported in Table 3. The two successful samples are from the same location but from two different layers: a sandstone and a conglomerate. Both rocks are sheared and contain newly formed chlorite (Fig. 9). In sample V1504, 61 grains could be counted on the 10.5 and 17.5 h etch. In sample V1503, 79 grains could be counted on four mounts with the three different etch times (10.5, 14, and 17.5 h). Both samples consist of multiple age populations as attested by the  $\chi^2$  test that gives a probability of 0% (Galbraith, 1981). V1504 has grain ages in the range from 75 to 660 Ma with a central age of  $156 \pm 10$  Ma and V1503 from 74 to 468 Ma with a central age of  $177 \pm 13$  Ma (Fig. 10; Table S4). At least two to three age populations can be identified using the software DensityPlotter (Vermeesch, 2012). The age distribution of sample V1504 has two major peaks: one centered at  $150 \pm 6$  Ma contains 84 % of the grains; the other at  $433 \pm 68$  Ma is formed by 16 % of the grains. The main younger peak might represent the sum of two populations at  $128 \pm 11$  Ma and at  $183 \pm 19$  Ma. The age distribu-



**Figure 7.** (a) Asymmetric boudin of a limestone showing top-to-the-NE shear sense. (b) Cigar-shaped clasts in a polymictic conglomerate of the mound top within the shear zone below the VOC (with sketch of the uniaxial ellipsoid mimicking the shape of the clasts). The stretch axis (x) of the prolate finite strain is parallel to the regional mineral and stretching lineation. (c) Tectonic contact between the Kallipetra Basin and the VOC with a top-to-the-NE shearing.

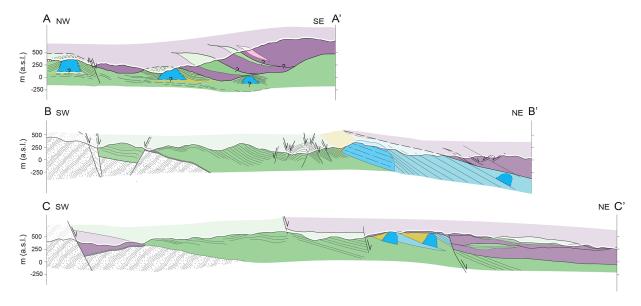


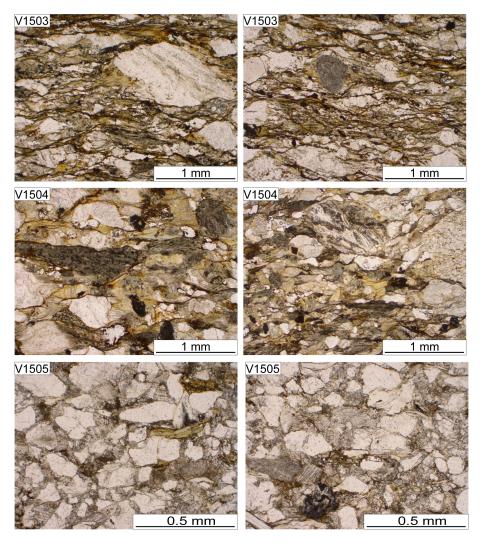
Figure 8. Geologic cross sections of the mapped area, colors corresponding to those on the geological map (Fig. 3).

tion of sample V1503 has a major peak with a pronounced shoulder and a long tail towards older ages. The central peak represents the largest age population that consists of 70% of the grain and that has an age of 158  $\pm$  14 Ma. The shoulder represents a minor population centered at 105  $\pm$  14 Ma with 19% of the grains. A third population might be located along the tail of the distribution at 252  $\pm$  50 Ma.

### 5 Discussion

## 5.1 Onset and evolution of the Kallipetra Basin

The orthogneiss- and serpentinite-rich composition of the previously described basal cataclasite suggests that it was formed prior to or at the same time as the formation of the Kallipetra Basin and mainly at the expense of the Pelagonian



**Figure 9.** (a) Sample V1503, newly formed chlorite; (b) sample V1503, newly formed chlorite; (c) sample V1504, newly formed chlorite; (d) sample V1504, newly formed chlorite; (e) sample V1505, detrital chlorite; (f) sample V1505, detrital chlorite.

basement and of the VOC. These normal faults crosscut duplicates of Pelagonian mylonitic marble and must be younger than ca. 120 Ma (Schenker et al., 2014). Normal faulting during or following the exhumation and doming of the Pelagonian zone from the late Early Cretaceous (Schenker et al., 2014; Schenker et al., 2015) probably contributed to subsiding the deformed wedge below sea level to create the basin. As discussed in Sect. 2.1, the onset of the Cretaceous basins along the Pelagonian zone is diachronous and remains enigmatic. However, some evidence may suggest that the Albian-Aptian topographic response to the Early Cretaceous continental accretion was uneven along strike of the Pelagonian zone, likely due to northward-decreasing shortening to a wider Rhodope to the south, thinning out to the north (Fig. 1). Indeed, to the south, non-metamorphosed Pelagonian sediments showing a sedimentary Aptian–Albian hiatus ( $\sim 120$ – 100 Ma) over lower Aptian deformed flysch and bauxitic laterites testify to a growing topography in the frontal part of the orogenic wedge (Nirta et al., 2015, 2018). To the north, where the Rhodope starts to thin out in map view (Fig. 1), the Pelagonian zone and western margin of the Vardar Zone were transgressed during Aptian–Albian times by marine sediments with an eastward deepening that evolved to carbonate–clastic successions (Mercier, 1968; Mercier et al., 1987; Brunn, 1982; Sharp and Robertson, 2006).

Serpentinite-rich conglomerates represent the first sediments deposited within the Kallipetra Basin through subaerial erosion of the VOC, which created an uneven topography. North of the study area, conglomerates containing serpentinite pebbles and Pelagonian marbles of Albian age are observed on the eastern Pelagonian zone border, demonstrating prior deep erosion of the obducted ophiolitic sheet likely of the Pelagonian Zone (Mercier and Vergely, 2002). The occurrence of conglomerates followed by a succession of calc

Table 2. Illite crystallinity data.

SP IL1         540         0.091         0.109         Epizone         310           CG IL1         535         0.141         0.122         Epizone         300           M3/2         527         no illite         310           M3/1         525         no illite         310           M2/16         520         0.181         Low epizone         295           M2 IL3         519         0.126         0.11         Epizone         305           M2/13         517         0.225         High anchizone         280           M2/13         517         0.225         High anchizone         230           M2/10         512         0.145         Epizone         300           M2/1         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285
M3/2         527         no illite         310           M3/1         525         no illite         310           M2/16         520         0.181         Low epizone         295           M2 IL3         519         0.126         0.11         Epizone         305           M2/13         517         0.225         High anchizone         280           M2/13         517         0.225         High anchizone         230           M2/13         517         0.225         High anchizone         230           M2/10         512         0.145         Epizone         300           M2/10         512         0.145         Epizone         290           M2/1         510         0.175         Epizone         290           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1/1<
M3/1         525         no illite         310           M2/16         520         0.181         Low epizone         295           M2 IL3         519         0.126         0.11         Epizone         305           M2/13         517         0.225         High anchizone         280           M2/13         517         0.258         0.131         Low anchizone         230           M2/10         512         0.145         Epizone         300           M2/10         512         0.145         Epizone         290           M2/7         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1/1         495         0.176         Epizone         290
M2/16         520         0.181         Low epizone         295           M2 IL3         519         0.126         0.11         Epizone         305           M2/13         517         0.225         High anchizone         280           M2/12         515         0.258         0.131         Low anchizone         230           M2/10         512         0.145         Epizone         300           M2/10         512         0.145         Epizone         290           M2/1         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizo
M2 IL3         519         0.126         0.11         Epizone         305           M2/13         517         0.225         High anchizone         280           M2 IL2         515         0.258         0.131         Low anchizone         230           M2/10         512         0.145         Epizone         300           M2/1         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1/1         496         0.127         0.116         Epizone         290           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         275           CRS1/2         350         0.18
M2/13         517         0.225         High anchizone         280           M2 IL2         515         0.258         0.131         Low anchizone         230           M2/10         512         0.145         Epizone         300           M2/7         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRS1/2         350         0.18         High anchizone         275           CRS1/1         345         0.286 <td< td=""></td<>
M2 IL2         515         0.258         0.131         Low anchizone         230           M2/10         512         0.145         Epizone         300           M2/7         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRS1/2         350         0.18         High anchizone         275           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191 <td< td=""></td<>
M2/10         512         0.145         Epizone         300           M2/7         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone
M2/7         510         0.175         Epizone         290           M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188
M2/4         507         0.157         Epizone         295           M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
M2/1         504         0.192         High anchizone         290           M1/7         500         0.137         Epizone         300           M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         290           CRN1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
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M1/5         498         0.131         Epizone         300           M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
M1/3         497         0.22         High anchizone         285           M2 IL1         496.5         0.131         0.122         Epizone         300           M1 IL1         496         0.127         0.116         Epizone         300           M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
M2 IL1       496.5       0.131       0.122       Epizone       300         M1 IL1       496       0.127       0.116       Epizone       300         M1/1       495       0.176       Epizone       290         CRN1/1       428       0.209       High anchizone       285         CRN1/3       418       0.25       Low anchizone       275         CRS1/2       350       0.18       High anchizone       290         CRS1/1       345       0.286       Low anchizone       230         CRS1/3       330       0.191       High anchizone       290         KM2 IL3       327       0.383       0.188       Detrital       250
M1 IL1       496       0.127       0.116       Epizone       300         M1/1       495       0.176       Epizone       290         CRN1/1       428       0.209       High anchizone       285         CRN1/3       418       0.25       Low anchizone       275         CRS1/2       350       0.18       High anchizone       290         CRS1/1       345       0.286       Low anchizone       230         CRS1/3       330       0.191       High anchizone       290         KM2 IL3       327       0.383       0.188       Detrital       250
M1/1         495         0.176         Epizone         290           CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
CRN1/1         428         0.209         High anchizone         285           CRN1/3         418         0.25         Low anchizone         275           CRS1/2         350         0.18         High anchizone         290           CRS1/1         345         0.286         Low anchizone         230           CRS1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
CRN1/3       418       0.25       Low anchizone       275         CRS1/2       350       0.18       High anchizone       290         CRS1/1       345       0.286       Low anchizone       230         CRS1/3       330       0.191       High anchizone       290         KM2 IL3       327       0.383       0.188       Detrital       250
CRS1/2       350       0.18       High anchizone       290         CRS1/1       345       0.286       Low anchizone       230         CRS1/3       330       0.191       High anchizone       290         KM2 IL3       327       0.383       0.188       Detrital       250
CRS1/1       345       0.286       Low anchizone       230         CRS1/3       330       0.191       High anchizone       290         KM2 IL3       327       0.383       0.188       Detrital       250
CRS 1/3         330         0.191         High anchizone         290           KM2 IL3         327         0.383         0.188         Detrital         250
KM2 IL3 327 0.383 0.188 Detrital 250
KM2 II.2 325 0.383 0.224 Detrital 250
11.12 122 0.303 0.21 Delitar 230
KM1/2 321 0.14 Smectite 200
KM1/1 320 0.168 200
KM1 IL1 316 0.383 0.353 Deep diagenetic zone 160–200
KM2 IL1 312 0.388 0.164 Deep diagenetic zone 160–200
KM2/8 308 0.209 Diagenetic 160–201
KM2/6 304 0.166 Diagenetic 160–202
KM2/4 300 0.16 Diagenetic 160–203
KM2/2 298 0.14 Diagenetic 160–204
CRS1/4 230 0.286 Low anchizone 230

arenites at the base of the basin indicates shallow marine depths. Marble olistoliths and slumping at the base of the Sfikia section (Fig. 4a) indicate instability during the first phases of basin formation and the presence of a proximal steep slope in which gravitational instability drove slumping. In the northwestern part of our study area, the presence of orthogneiss blocks, the dominance of quartz, feldspar, gneiss, and marble lithics in the sediments, and the lack of such components in the southeast suggest an intra-basinal high, emergent land or continent existed northwest of the Kallipetra Basin, where the Pelagonian basement was exposed. In the southeastern part of the study area, the dominance of silty limestones, marls, and lime mudstones and the rare presence of olistoliths indicate that there was a deepening of the basin towards the southeast. The mid part of the Kallipetra Formation is devoid of serpentinite coarse detritus, suggesting the initial fault scarps were smoothed by sediments. This expansion of the basin toward the southern slope is documented by the transgression of Kallipetra deposits onto the VOC, recorded in the study zone. It is worth noting that this basin widening corresponds to the deepest facies in the basin, likely correlating the Cenomanian–Turonian eustatic sea level high (e.g., Haq, 2014). Mercier (1968) and Sharp and Robertson (2006) also record marine transgression and eastward deepening of mixed carbonate–clastic successions of the combined Pelagonian and western Almopias zones.

# 5.2 The rudist mounds: facies and evolution at the slope of the Kallipetra Basin

Rudists constituted more than 60 % of reef frames during the Aptian and Albian and became the most dominant frame-building organism in the Late Cretaceous (Scott, 1988; Voigt et al., 1999). Widespread tectonic extension combined with eustatic continental flooding occurring around the Cretaceous Tethyan Ocean allowed the growth of broad carbon-

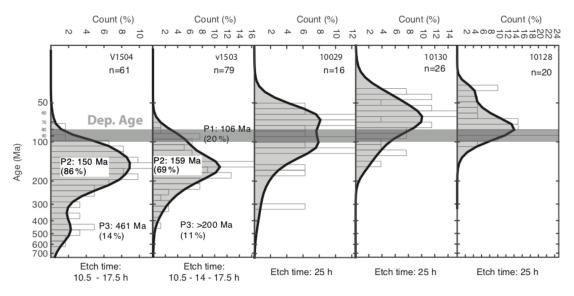


Figure 10. Zircon fission track ages of samples taken in the mapped area.

ate platform complexes, on slopes from a few up to  $40^{\circ}$  (Gili et al., 1995). Previously described carbonate mud mounds and rudist biostromes have some similarities to the rudist mounds observed in the study area (e.g., Camoin, 1995; Negra et al., 1995; Sanders, 1998; Sanders and Pons, 1999; Sanders and Höfling, 2000). It has been suggested that bioerosion processes leading to pervasive micritization of invertebrates may result from endolithic microorganism activity, accounting for part of the lime muds (Camoin, 1995). Alternatively, Camoin (1995) also suggested that the lime muds are formed through in situ precipitation promoted and induced by microbial activity and/or the decay of microbial communities. The latter is the most likely option regarding these mounds, with the dense micrite deposited as leiolite (sensu Riding, 2000) in a microbe-rich upper slope. Microbial mud mounds were common in the Late Cretaceous of the western Tethys, especially on the shelf or ramp rimming the Adria microplate (e.g., Picotti et al., 2019). The distinct dome shape of the mounds built mainly from lime mud, suggests that the mounds themselves were sites of increased carbonate productivity. Furthermore, the upward growth of rudists is said to be an adaptation to environments with positive net sedimentation rates (Gili et al., 1995), which may be the case for the Kallipetra Basin. The sub-vertical and sharp nature of some of the contacts between the mound cores and the flank deposits and the presence of breccia bodies suggest early diagenetic consolidation allowing stability of the steep mound slopes. By combining the observations of breccia bodies stacking up against the southern flanks of mounds and the presence of stratigraphically underlying slumping and massflow deposits, the rudist mounds were built on a slope environment. Open shelf or shallow-water platforms, conditions suggested by Scott (1988) and Camoin (1995), are unlikely for the Kallipetra Basin due to its dynamic and tectonically

active history. The termination of the rudist mounds, found in the upper part of the stratigraphy of the Kallipetra Basin, occurred by environmental restriction due to gradually increased sediment input from the approaching ophiolite talus (Sanders and Pons, 1999; Sanders and Höfling, 2000).

The mound flanks consist of a succession of sandstones, breccias, and occasionally marls. Our observation of polymictic breccias on the south–southwestern mound slopes bears an important paleogeographic meaning. In this case, the presence of serpentinite clasts indicates the breccias were not formed solely from the erosion and collapse of mound flanks, but rather they were derived from an ophiolitic source upslope from the mounds, possibly associated with new fault scarps in the south-southwestern slope of the Kallipetra Basin. On the northern flanks of the mounds, the sediments display a shallower dip and are interfingered with the mound talus breccias. The youngest and northernmost mound at Asomata Quarry displays pelagic marls and limestones at the northwestern flank, suggesting deeper bottom conditions to the N and NE. On the other hand, the absence of serpentinite detritus in the mound flanks other than the southern ones documents a shadow effect of the mounds with respect to the south–southwestern provenance of the serpentinite clasts. This evidence corroborates the presence of a slope dipping to the north-northeast. North of the study area, Cenomanianaged hippuritid-bearing rudist mounds have been observed, where Sharp and Robertson (2006) suggest they also developed on an east-facing ramp. These authors also observe younger Campanian–Maastrichtian rudist biostromes that developed on an east-facing ramp-shelf margin in the Pelagonian and western Almopias zones. However, as discussed in the following sections, these must have formed subsequent to Kallipetra Basin closure, likely in basinal areas not involved in Turonian compression. The increasing ser-

Tircon fission-track data. Variable amounts of zircons were analyzed on multiple mounts for each sample that were etched for different times. As a fluence monitor, a glass standard CN1 with a U concentration of 39.8 ppm was used. Central ages were calculated using a  $\xi$  calibration value of 145.39 $\pm$ 7.04. nD and  $\rho_D$ : number and density of induced tracks from the fluence monitor.  $n_s$  and  $\rho_s$ : number and density of spontaneous tracks in the zircons.  $n_i$  and  $\rho_i$ : number and density of spontaneous tracks from the zircons.  $P_{\chi^2}$ :  $\chi^2$  probability.

ı	1	1
σ1 Ma	10.07	13.22
Central age Ma	155.99	34 176.65 13.22
Age dispersion %	24	34
$P_{\chi^2}$	0	0
$\begin{array}{ccc} \rho_{\rm i} & P_{\chi^2} \\ 10^6  {\rm tracks} & \% & {\rm di} \\ {\rm cm}^{-2} & & & \\ \end{array}$	2.939	2.844
n <sub>i</sub> tracks	1353	1083
$\frac{\rho_{\rm S}}{{ m cm}^{-2}}$	12.226 1353	13.396 1083
n <sub>s</sub> tracks	5628	5101
$\rho_{\rm d}$ $10^5  {\rm tracks}$ ${\rm cm}^{-2}$	5.29 5.279 5.267 5.256	5.233
n <sub>d</sub> tracks	6594 6579 6565 6551	6523 6537
No. grains	10 19 27 23	18
Etch time h	17.5 14 10.5 10.5	17.5
Mount	a c c	в
N Elevation m m	800	800
Z E	4476334.86	4476334.86
ш ш	V1503 34T 607451.37 4476334.86	V1504 34T 607415.37 4476334.86
MITU	34T	34T
Sample UTM ID	V1503	V1504

pentinite content in the sandstones and breccias up-section suggests that the ophiolitic source was moving closer to the mound structure and gradually providing material to the slope. The positioning of the flank deposits and the northeastward directed stacking pattern of the two or three youngest mounds, with the highest – and therefore youngest – mound being at the Asomata Quarry in the northeast of the study area, suggest that in the upper part of the Kallipetra stratigraphy, there was a movement of the ophiolite (VOC) from SW to NE providing at first the slope for the growth of the mounds and then the burial for them (Figs. 12 and 13).

### 5.3 Stratigraphy and age of the Kallipetra Formation

The Kallipetra Formation was deposited on top of the eroded Pelagonian continent and obducted ophiolite following the collision-related burial and cooling or exhumation of the Pelagonian zone at  $\sim 116$  Ma (Aptian; Schenker et al., 2014).

The red and green limestones can be loosely correlated across the Kallipetra Basin and they first occur at approximately  $\sim 250\,\mathrm{m}$  from the base of the Kallipetra construction road section and 400 m from the base of the Sfikia section (Fig. 4a). These facies indicate deposition in a deep, calm pelagic environment. The distinct red and green alternations are typically attributed to bottom-water redox cycles (e.g., Luciani and Cobianchi, 1999), developed around the upper Cenomanian OAE2 (Luciani and Cobianchi, 1999; Negri et al., 2003; Mort et al., 2007). This event belongs to the global Cenomanian–Turonian sea level high (Haq, 2014), therefore explaining the relative absence of clastic input in this interval, which could represent the deepest stage of the basin development. Indeed, this should be the timing of the onlap of the Kallipetra Basin toward the southern ophiolitic slope.

Rudist mounds and breccias are lacking in the Kallipetra sediments found on top of the VOC (Figs. 3 and 8), whereas fine sediments dominate. The fine material suggests that the tip of the VOC was under sea level, with a transgressive trend and widening of the basin, allowing the onlap of fine material over the VOC slopes at the same time as red and green marl deposition. During this time, the source area for sediments is moving away. This Kallipetra material overlying the VOC is somewhat separated from the main Kallipetra Basin sediments, possibly through a structural high or as perched basins (Fig. 13). Alternatively, these deposits represent one flank of the basin that was subsequently tectonically emplaced over the basin, suggesting this movement was just a few kilometers and a local event.

Helvetoglobotruncana helvetica indicates that the marls overlying the mound at the top of the Kallipetra stratigraphy are lower Turonian. This agrees with the nannoplankton that we found in the overlying section that are early to middle Turonian. The youngest proven depositional age of the Kallipetra Basin is  $\sim$  92 Ma, but the top 35 m of hemipelagic marls are barren; therefore we cannot exclude a late Turo-

nian or even Coniacian age for the very top of the Kallipetra

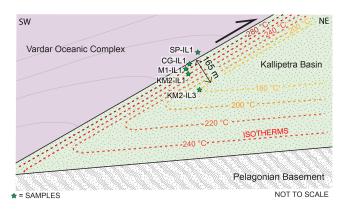
The stratigraphic thickness between the red and green marls, and the marls adjacent to the mound core containing the Helvetoglobotruncana helvetica is 200 m. By using an age of 93.9 Ma (Cenomanian-Turonian boundary) for the red and green marls (Cohen et al., 2013) and the youngest age of Helvetoglobotruncana helvetica - 91.3 Ma (BouDagher-Fadel and Price, 2019) – for the marls at the top of the section, the average sedimentation rate of the basin infill is  $0.08 \,\mathrm{mm}\,\mathrm{yr}^{-1}$ , a value compatible with the recorded mixture of pelagic and terrigenous sediments. Assuming a constant sedimentation rate during infilling of the basin, the first sediments deposited at the bottom of the basin are lower Cenomanian. This would suggest approximately 20 Ma of erosion and/or subsidence between the final stage of the collisional doming of the Pelagonian basement and the subsidence and deposition of the first Kallipetra sediments. This 20 Ma time interval agrees with the Aptian-Albian sedimentary hiatus ( $\sim$  120–100 Ma) described further south in the Pelagonian zone by Nirta et al. (2015, 2018).

# 5.4 Zircon fission-track age distribution and thermal overprint

Our new ZFT samples come from the top of the Kallipetra Basin where the depositional age should not be older than 92 Ma and therefore should be younger than the ZFT central ages of our samples, which range between 156 and 177 Ma (Fig. 10). However, both ZFT samples contain a few young grains with ages overlapping with the depositional age of the Kallipetra Formation (Fig. 10). They are located a few hundred meters to the south of a previous sample, 10-029, that is adjacent to the contact with the VOC (Schenker et al., 2015; Fig. 10). The age range of this sample is from 52 to 340 Ma, and it consists of 16 grains that define only one age population centered at  $92 \pm 9$  Ma, in overlap with the depositional age. Thus, all these samples can be interpreted as partially to non-annealed, and the sample closest to the contact with the VOC has the youngest age. They are all from clastic sediments that contain newly formed chlorite. No illite crystallinity data are available as the rock type of the ZFT samples does not allow the illite method to be applied. However, we observed a NE-SW metamorphic gradient along the section where we collected the dated samples together with others that unfortunately provided no zircons. This gradient is indicated by the fact that the sandstones of V1503, V1504, and 10-029 contain newly formed chlorite, whereas the sandstones (V1505) located 4 km to the SW towards Sfikia show only detrital minerals (Figs. 4 and 9). The presence of newly formed chlorite in samples V1503, V1504, and 10-029 suggests temperature conditions > 250 °C, which could be within the partial annealing zone (PAZ) for natural zircons bearing radiation damage (Reiners and Brandon, 2006). However, the temperature range of the PAZ depends not only on the degree of radiation damage of the zircons but also on the rate of heating and cooling such that during a short-lived heating event, followed by rapid cooling, higher temperatures are needed to obtain fully reset ages. The fission-track kinetic parameters in natural zircons are constrained only based on exposed fossil annealing zones (Brandon et al., 1998) such that modeling their time—temperature history would not give any deeper insight into the conditions that could have produced the observed age distribution.

Two more samples were previously dated along the Kallipetra section where we collected our new illite crystallinity data (Fig. 10; Schenker et al., 2015). There, 20 grains from the sample at the top (10-128) of the section define an age range between 39 and 102 Ma and a central age of  $67 \pm 4 \,\mathrm{Ma}$ ; 26 grains from the lower sample (10-130) have ages from 40 to 158 Ma and centered at  $72 \pm 5$  Ma. These samples come from the top of the Kallipetra section, and they were previously interpreted as non-annealed. However, based on the revised depositional age of the Kallipetra Basin documented here, the central ages of these samples result in ages younger than the depositional age, but their age ranges partly overlap with the depositional age. Thus, these samples can be defined as partially to fully annealed and they are younger than the samples 10-029, V1503, and V1504. The illite crystallinity data indicate temperatures up to  $\geq 300\,^{\circ}\text{C}$  towards the top of the Kallipetra section. The top ZFT sample 10-128 is from a higher stratigraphic location than the illite samples, whereas sample 10-130 comes from the same location as the uppermost illite samples. Thus, the ZFT samples along the Kallipetra section should have been subject to  $T \ge 300 \,^{\circ}\text{C}$ , but we cannot say if and how much higher these temperatures could have been relative to the other samples.

The different ZFT age ranges and central ages hint at highly variable degrees of annealing. Our petrographic and illite crystallinity data constrain a strong, inverse, vertical (up-section) thermal gradient, but they cannot discriminate between possible lateral gradients across the basin. However, they indicate that the Kallipetra Basin has been subject to temperatures that locally could have totally or partially annealed our samples. Whether these gradients are reflected by the ZFT central ages or grain-age distributions must be carefully weighed against other factors that could also affect our results. In fact, we processed the new and the old samples purposefully in different ways because, while processing the previous set of samples, we realized that the low number of available zircons limited the applicable etch procedure, which was not optimal to reveal the full age spectra of our samples. However, even though at the time we opted for an etch procedure aimed at maximizing the young grain ages, our results indicated that the annealing degree of our samples might have been incomplete. With the new samples, we aimed at verifying the degree of annealing by maximizing the zircon yield that allowed applying a multiple etch procedure. This in turn revealed that in fact there are wide age distributions in the new samples, which include non-reset ages, and



**Figure 11.** Schematic diagram showing the inverse geothermal gradient at the contact between the VOC and Kallipetra Basin.

this confirmed our previous observations on a partial degree of annealing. Unfortunately, our new data do not answer all the questions concerning the ZFT ages in the study area but highlight a complex thermal and annealing record.

# 5.5 The inverted geothermal gradient in the Kallipetra Basin

The KI data constrain an inverse geothermal gradient at the top of the Kallipetra Basin from  $> 300\,^{\circ}\text{C}$  at the tectonic contact to  $100\text{--}200\,^{\circ}\text{C} \sim 165\,\text{m}$  below the overridden VOC (Figs. 4 and 11). Stratigraphically below this zone, the sediments reached only deep diagenetic conditions. The newly formed syntectonic chlorites in the top sediments at the base of the VOC further testify to high ( $> 200\,^{\circ}\text{C}$ ; Beaufort et al., 2015) and inverse temperatures that peaked at the time of deformation.

The illitization reaction (i.e., the conversion of smectiterich I-S into illite-rich I-S) is also dependent on the availability of K<sup>+</sup> ions, sometimes requiring enhanced K<sup>+</sup>-rich fluid circulation (Dellisanti et al., 2008). However, the agreement of temperatures between the chlorite-in reaction and the KI values suggests that the K<sup>+</sup> ions were available during the increase in the metamorphic conditions and that other mechanisms influencing the crystallinity of illite such as shearrelated recrystallization (Merriman and Peacor, 1998; Árkai et al., 2002) were less important in controlling the KI values. Hence, it is likely that the KI values represent metamorphic temperatures recording an inverse geotherm. Sedimentary strata within thrust belts are known to sometimes experience transient thermal histories, and "sawtooth" geotherms with inverse metamorphic fronts from the base of the hangingwall into the footwall have been recognized in a series of thrust systems (Graham and England, 1976; Furlong and Edman, 1989).

Overall, our data document an inverse thermal gradient of the Kallipetra Basin, pointing to a syntectonic heating event that produced a transient, inverse, non-linear, and disturbed geotherm (Fig. 11). The sedimentary history suggests that the closure of the Kallipetra Basin by the VOC occurred just after the deposition of the ophiolitic debris that buried the rudist mounds, when the sediments were porous, permeable, and saturated. Although the ultimate sources of this heat have not been established, the non-reset to partially reset FT ages testify that this syntectonic heating event formed in the Late Cretaceous, during the closure of the basin in the Turonian. Cooling slightly postdates the deformation as the youngest ZFT population is older than the Turonian closure.

# 5.6 Sealing of the Kallipetra Basin and large-scale implications

The stacking pattern of the rudist mounds documents closure of the Kallipetra Basin through a NE-facing slope and serpentinite detritus supplied from the SW. This pattern cannot be explained by the activity of a normal fault (Fig. 12-2a), since, in this case, the stacking of the mounds should have been to the SW, following the widening of the basin. Therefore, only a NE-verging thrust of the VOC found at the southwestern margin of the basin can explain the observed rudist mound stacking pattern, as well as the serpentinite breccias and the northeast mound shadow (Figs. 12-2b and 13). Kinematic indicators, such as shear bands, stepover structures, and sigma clasts along the upper tectonic contact of the Kallipetra Basin, indicate a top-to-the-NE tectonic movement. If the NE-dipping contact of the VOC over the Kallipetra Basin is regarded as the pristine fault orientation, the sealing may be interpreted as a Turonian or younger normal fault (Fig. 12-2a.i). However, this interpretation fails to explain the observed cutoff angles, since the tectonic contact should have cut the basinal deposits down-section, and this is not the case (Figs. 8 and 12-2a.ii). The cutoff of the Kallipetra deposits is compatible only with a NE-directed thrust (Fig. 12-2b.i). Post-Turonian tectonics are considered responsible for the northeastward block rotation and the normal faulting (Fig. 12-2b.ii).

The partially to fully annealed ZFT ages combined with illite crystallinity and crystallization of syntectonic chlorite, indicating high temperatures at the tectonic contact, but only deep diagenetic conditions below, suggest that the tectonic movement occurred in the Turonian when sediments were not fully compacted and still permeable. Therefore, the Kallipetra Basin was sealed in the Turonian by SW to NE tectonic transport of the VOC (Figs. 12 and 13). Northeast-directed thrusting is documented some 50 km north of our study area in the Almopias zone by Vergely and Mercier (2000), although it is considered Tertiary in age by the quoted authors. However, our results differ from previous studies that have documented progressive deepening from the Aptian up to flysch-like Maastrichtian to Paleogene sediments (Mercier and Vergely, 2002) or the development of thick Santonian-Campanian carbonates followed by a foredeep succession indicating dramatic subsidence in the Late Maastrichtian (Sharp and Robertson, 2006). The ob-

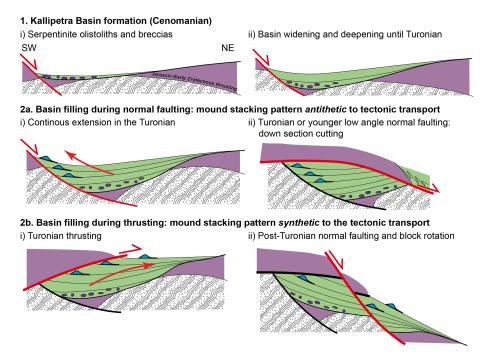
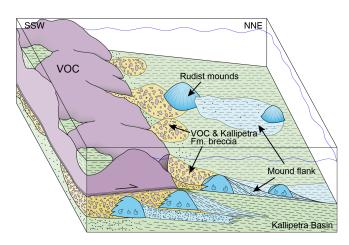


Figure 12. A series of sketches to demonstrate the opening of the Kallipetra Basin in the Cenomanian and the closure of the Kallipetra Basin under normal faulting or thrust faulting conditions. Only a NE-verging thrust of the VOC found at the southwestern margin of the basin can explain the observed rudist mound stacking pattern, as well as the serpentinite breccias and the northeast mound shadow. Colors and patterns correspond to those on the geological map (Fig. 3).



**Figure 13.** Schematic diagram showing the sedimentary and tectonic environment of the Kallipetra Basin during the Turonian and the overriding of rudist mounds by the resumed thrusting of the VOC. Colors correspond to those on the geological map (Fig. 3).

served stratigraphy from this study, such as olistoliths, breccias, and slump deposits, suggests deposition very close to steep basin margins. Therefore, it is plausible that the closure of the Kallipetra Basin observed here only affected the margins of a larger-scale basin, which experienced continuous deepening and transgression as observed to the north of the study area by Mercier and Vergely (2002). Alternatively,

the Kallipetra Basin may not be laterally continuous, and its birth and closure would have only had local significance.

During the Late Cretaceous to Eocene, along the eastern Pelagonian margin, the dominant deformation at the regional scale is SW-verging thrusting (e.g., Schenker et al., 2015, and references therein). Bivergent thrusting occurred locally but later in time during the late Late Cretaceous or Tertiary (Vergely and Mercier, 2000; Brown and Robertson, 2003; Katrivanos et al., 2013). Thus, the sealing of the Kallipetra Basin occurred earlier than or in the very early phase of this regional deformation event, although the direction of tectonic transport of the VOC above this basin is opposite to the common SW vergence of thrusting. This apparent difference may be explained by a localized basin inversion rather than a regional tectonic event that predated the start of the regional convergence in the Late Cretaceous, i.e., the Campanian at the earliest (Aubouin, 1973; Baumgartner, 1985; Godfriaux and Ricou, 1991; Bonneau et al., 1994; Papanikolaou, 1997; Brown and Robertson, 2003; Grubić et al., 2009; Ustaszewski et al., 2009; Kilias et al., 2010; Katrivanos et al., 2013; Schmid et al., 2020).

Overall, the thermo-tectono-sedimentary history documents a basin that likely formed in the early Cenomanian over a suture with accreted Jurassic ophiolites in the hanging-wall and an Aptian metamorphic basement in the footwall. This attests its intracontinental position within an orogenic wedge in a tectonic scenario similar to the model proposed

by Schenker et al. (2014) (Fig. 2). The Kallipetra Basin cuts and covers the Pelagonian basement with an Aptian regional metamorphism that cannot be associated with the metamorphic sole beneath the Jurassic-Early Cretaceous obduction but rather with a Barrovian metamorphism formed during the collision of the Pelagonian zone with the Rhodope (Schenker et al., 2014). Accordingly, geodynamic models (cf. Fig. 2) that invoke a Jurassic-Early Cretaceous obduction without an Aptian crustal-scale accretion (e.g., Sharp and Roberson, 2006; Froitzheim et al., 2014) or the distal position from the Pelagonian zone of the Rhodope during the Early Cretaceous (e.g., Ricou et al., 1998; Froitzheim et al., 2014) do not agree with the thermo-tectonic history of the Pelagonian zone prior to the formation of the Kallipetra Basin. The extensional phase that opened the Kallipetra Basin remains enigmatic and may be associated with an isostatic re-equilibration of the orogenic wedge or with a far-field plate tectonic reorganization (e.g., Matthews et al., 2012). The closure of the basin anticipates the beginning of resumed ophiolitic imbrication in this sector of the Internal Hellenides in the Turonian and is considered a local basin inversion. If the actual Hellenic subduction is considered active since at least the Early Cretaceous (van Hinsbergen et al., 2005), the closure of the Kallipetra Basin could be seen as early evidence in the upper crust of the initiation of the Hellenic slab.

#### 6 Conclusions

The evolution of the Kallipetra Basin documents the transition from extension to compression during the early Late Cretaceous along the eastern margin of the Pelagonian zone in northern Greece. The history of the Kallipetra Basin can be summarized as follows:

- The sediments of the Kallipetra Basin were deposited between the early Cenomanian ( $\sim 100\,\mathrm{Ma}$ ) and the latest Turonian ( $\sim 90\,\mathrm{Ma}$ ) over the VOC and the Pelagonian basement in a depression deepening to the east and northeast. The depression formed initially by extension as testified by normal faults at the base of the basin.
- As the basin widened, a topographic high located to the NW and exposing Pelagonian basement rocks became the main source of siliciclastic detritus to the basin. The basin widened and deepened to the point when no clastic input reached it. This time might correlate with the global Cenomanian-Turonian sea level transgression. Carbonate sediments were produced by pelagic organisms and by rudist-rich microbial mounds growing on the southwestern slopes of the basin (Fig. 13).
- The terrigenous input was later renewed, and the main source were ophiolitic rocks to the south or southwest, which provided breccias stacking up against the southern flanks of the rudist mounds. The progressive in-

crease in detrital input restricted the environments of the rudist mounds.

The ophiolitic rocks overrode the Kallipetra Basin from the SW, causing uneven deformation of its sediments (Fig. 12). Thrusting was associated with a high inverted geothermal gradient that caused illitization, crystallization of chlorite, and partial-to-total annealing of the fission tracks in detrital zircons close to and increasing towards the top of the basin. Deformation, illitization, and zircon fission track annealing occurred during the Turonian and were followed by cooling in the late Late Cretaceous, anticipating the beginning of the resumed tectonics in this sector of the Internal Hellenides by about 10 Ma.

Data availability. Additional data are available from the corresponding author upon request.

*Supplement.* The supplement related to this article is available online at: https://doi.org/10.5194/se-11-2463-2020-supplement.

Author contributions. LRB is the primary author, conducted the study for her MSc thesis at ETH Zurich, and wrote the paper. FLS conceptualized the original research goals and aims of this study, and contributed to the interpretations, introduction, and background sections of this paper. MGF conducted the fission track analysis, provided contributions to Sects. 1, 3.3, and 4.5 and to interpretations of data and also provided mentorship. MC prepared samples for and conducted all planktonic foraminifera analysis for this study. TA prepared samples for illite crystallinity analysis and conducted the latter as well as giving constructive suggestions on the final paper draft. VP contributed to the interpretation and provided supervision and mentorship in all aspects of this study.

Competing interests. The authors declare that they have no conflict of interest.

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