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# Subduction to the lower mantle – a comparison between geodynamic and tomographic models

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#### Abstract

It is generally believed that subduction of lithospheric slabs is a major contribution to thermal heterogeneity in Earth's entire mantle and provides a main driving force for mantle flow. Mantle structure can, on the one hand, be inferred from plate tectonic <sup>5</sup> models of subduction history and geodynamic models of mantle flow. On the other hand, seismic tomography models provide important information on mantle hetero-geneity. Yet, the two kinds of models are only similar on the largest (1000s of km) scales and are quite different in their detailed structure. Here, we provide a quantitative assessment how good a fit can be currently achieved with a simple viscous flow geo-<sup>10</sup> dynamic model. The discrepancy between geodynamic and tomography models can indicate where further model refinement could possibly yield an improved fit. Our geodynamical model is based on 300 Myr of subduction history inferred from a global plate reconstruction. Density anomalies are inserted into the upper mantle beneath subduction zones, and flow and advection of these anomalies is calculated with a spheri-

- <sup>15</sup> cal harmonic code for a radial viscosity structure constrained by mineral physics and surface observations. Model viscosities in the upper mantle beneath the lithosphere are  $\sim 10^{20}$  Pas, and viscosity increases to  $\sim 10^{23}$  Pas in the lower mantle above D". Comparison with tomography models is assessed in terms of correlation, both overall and as a function of depth and spherical harmonic degree. We find that, compared
- to previous geodynamic and tomography models, correlation is improved significantly, presumably because of improvements in both plate reconstructions and mantle flow computation. However, high correlation is still limited to lowest spherical harmonic degrees. An important ingredient to achieve high correlation in particular at spherical harmonic degree two is a basal chemical layer. Subduction shapes this layer into two
- <sup>25</sup> rather stable hot but chemically dense "piles", corresponding to the Pacific and African Large Low Shear Velocity Provinces. Visual comparison along cross sections indicates that sinking speeds in the geodynamic model are somewhat too fast, and should be  $2 \pm 0.8$  cm yr<sup>-1</sup> to achieve a better fit.





# 1 Introduction

At convergent plate margins, slabs of subducted lithosphere start their journey toward the Earth's interior, and seismic tomography is arguably the best tool to track their sinking. Based on such seismic models, the opinion of most scientists is currently that <sup>5</sup> most slabs eventually sink to the base of the mantle (Grand et al., 1997; van der Hilst et al., 1997; van der Voo et al., 1999), where they accumulate. However, slab sinking trajectories are complicated through their interaction with phase transitions, particularly the spinel-perovskite transition at 660 km depth, where some slabs may lay flat for a while before they sink further (e.g. Fukao et al., 2001). While long-debated, the inter-<sup>10</sup> actions of slabs with the transition zone, and the degree of mass transport, are still unclear.

By comparing slab locations predicted from geodynamic models based on subduction history, both plate reconstructions (e.g. Bunge and Grand, 2000; Hafkenscheid et al., 2006; Zhang et al., 2010; Shepard et al., 2012) and geodynamic model parame-

- ters, such as slab sinking rates and mantle viscosity (e.g. Ricard et al., 1993; Lithgow-Bertelloni and Richards, 1998; Liu et al., 2008; van der Meer et al., 2010; Čížková et al., 2012), can be constrained. Towards that goal, we present here a simple geodynamic model of mantle density based on subduction history, and compare it to seismic tomography. We both visually compare along cross sections and compute formal corre-
- <sup>20</sup> lations (cf. Ray and Anderson, 1994; Becker and Boschi, 2002). Our work is essentially an update of Steinberger (2000), which we believe is appropriate now, as both models of seismic tomography and subduction history have changed since then. We also compare our results with those of a simple slab sinker approach (Ricard et al., 1993; Lithgow-Bertelloni and Richards, 1998), as well as an updated slab sinker approach headed on our comparison biotecomercial and the sinking rate of 0 answ<sup>-1</sup> informal
- <sup>25</sup> based on our own subduction history model and the sinking rate 1.2 cm yr<sup>-1</sup> inferred from van der Meer et al. (2010), in order to assess whether the geodynamic model in fact leads to an improvement.





Our geodynamic model also includes a thermochemical layer in the lowermost mantle, which is shaped by subduction into piles (e.g. McNamara and Zhong, 2005; Garnero and McNamara, 2008; Zhang et al., 2010). A similar global comparison has recently also been performed by Shephard et al. (2012). However, they use a different mantle convection code and mantle viscosity structure, their mantle model does not include thermochemical piles, they use a somewhat different subduction input model, and they also compare their results to a different tomography model.

## 2 Methods

# 2.1 Geodynamic models

- The geodynamic models are very similar to Steinberger and Torsvik (2012) mostly, unless noted otherwise, the thermo-chemical model, as shown in their Fig. 5 where the emphasis was on the creation of plumes, whereas here we focus on subduction. Since Steinberger and Torsvik (2012) describe their model in more detail, we give here only a brief description: it is based on 300 Myr of subduction history in the true polar wander (TPW) corrected, global hybrid reference frame (Steinberger and Torsvik, 2008; Torsvik et al., 2008) corrected in longitude according to van der Meer et al. (2010) (Fig. 1), i.e. density anomalies are added beneath the locations of subduction at each time step. The differences between Steinberger and Torsvik (2012) and the models st12den-1 and st12den-2 shown in this paper are the depth dependence of thermal density anomalies (for both models), where and when density anomalies are added (for st12den-2) and the consideration of phase boundaries (for st12den-2); we
- added (for st12den-2), and the consideration of phase boundaries (for st12den-2): we use here a stronger depth dependence of thermal expansivity  $\alpha$  than Steinberger and Torsvik (2012): we assume a radial profile of  $\alpha$  corresponding to the lower mantle profile of Steinberger and Calderwood (2006) but extrapolated to the surface. Consequently, density anomalies corresponding to subducted slabs become considerably smaller as





they sink through the lower mantle. In model st12den-2 we add density anomalies

14 Myr after subduction vertically below subduction zones at the two radial layers at depth 650 km and 700 km, i.e. near the top of the lower mantle, corresponding to a presumed slab sinking speed of ~5 cm yr<sup>-1</sup> in the upper mantle, similar to typical speeds of plate motion. Between their time of initial subduction and 14 Myr later, slabs are not included in the density anomalies that drive mantle flow. We make this simplification because in the upper mantle slabs are still mechanically attached to the plates and a viscous flow model may therefore not be fully appropriate (Becker and O'Connell, 2001;

Conrad and Lithgow-Bertelloni, 2002). Also, we are here mainly concerned with what happens to slabs after they enter the lower mantle. However, we consider upper mantle slabs for the correlation computations and the cross section plots: slabs subducted at 14 Ma are assigned to depth 650 and 700 km, slabs from 12 Ma to 550 and 600 km, and so on, upward, until slabs from 2 Ma are assigned to the 50 and 100 km layers.

In model st12den-2 the Clapeyron slope of phase boundaries is considered by adding mass anomalies at the depth of the phase boundary. Specifically, we add here a mass per area at depth 660 km that is equal and opposite to the density anomaly just below 660 km (here at depth 700 km) multiplied with a 77 km thickness. This thickness corresponds to a product of Clapeyron slope and density difference of  $300 (\text{MPa K}^{-1}) (\text{kg m}^{-3})$  as suggested by Akaogi and Ito (1999), and a thermal expansivity of  $\alpha = 2.1 \, 10^{-5} \, \text{K}^{-1}$  (Steinberger and Calderwood, 2006). In the same way, we use a thickness 105 km (and equal, not opposite, density anomaly) for modeling the effect of the "410", however this plays almost no role, as density anomalies are inserted into the model below that depth.

These slight modifications were made because a slab that has a realistic width of 100 km at the depth of the "660" interacts with the phase transition to slow down sinking into the lower mantle. We expected that, combined with the delay of 14 Myr of inserting slabs and the stronger decrease of thermal expansivity with depth, an increased sinking time and hence improved correspondence between predicted slab locations and tomography could result. However, the effect of these modifications turned out to be rather minor.





Flow is modeled with a spherical harmonic approach (Hager and O'Connell, 1979, 1981). We use radial viscosity profile M2b of Steinberger and Calderwood (2006) – that is, the shape of the profile in each layer (upper mantle, transition zone, lower mantle) is inferred from mineral physics, and absolute viscosity values in each layer are adjusted to optimize fit to geoid and global heat flux (for a flow model based on seismic tomog-

- to optimize fit to geoid and global heat flux (for a flow model based on seismic tomography) and satisfy the "Haskell" constraint from postglacial rebound (Mitrovica, 1996). We use a tangential stress free surface as upper and lower boundary condition, include diffusion of heat through the CMB, but apply an "isolating" upper surface, in order to not include twice the same effect, which is already considered by the explicit addition of subducted slabs into the model. Our model also includes a chemical layer at the base
- of the mantle, which is initially 70 km thick at the time of model initiation at 300 Ma.

As our model output, we compute the present-day temperature or thermal density field, which is compared with a seismic tomography model as discussed in the next section. As a model enhancement, we have introduced slab tracers. They are also

- added vertically below subduction zones at depths 650 and 700 km and advected with the flow using a 4th order Runge-Kutta scheme (Press et al., 1986). They carry as information their weight, location and time of insertion. In this way, we can more directly compute sinking times and speeds at specific locations, as well as averages and variability of sinking times and speeds and lateral motion, and compare these modeling
   results with observation-based estimates (e.g. van der Meer et al., 2010) to comple-
- results with observation-based estimates (e.g. van der Meer et al., 2010) to comple ment our own comparison with tomography.

#### 2.2 Tomography model

A great number of tomography models have been published over recent years. Here we mainly compare to the SMEAN composite S-wave tomography model of Becker and Boschi (2002) (Fig. 2) which was obtained by RMS-weighted averaging of S20RTS (Ritsema and van Heijst, 2000), NGRAND (a 2001 update of Grand et al., 1997), and SB4L18 (Masters et al., 2000). SMEAN has been found to outperform other models



in geodynamic tests (Steinberger and Calderwood, 2006) and yields good variance reductions when put to the test with actual seismic waveforms (Qin et al., 2009).

We also constructed an update, which we call SMEAN2, where we replaced S20RTS by S40RTS (Ritsema et al., 2011) and NGRAND by TX2008 (the purely seismic model

- of Simmons et al., 2009). Differences between SMEAN2 and SMEAN are rather minor: correlations are close to unity up to spherical harmonic degrees / ~ 14, and average correlation up to degree / = 20 is 0.95, and we also find that different mean models such as a simple average of these three models, or the average of Buiter et al. (2012), or similar averages where we only include those models that have global coverage, or
- only S-wave models with global coverage, look all very similar. Correlation and visual agreement among different S-wave models and even more so among the various mean models constructed is generally much higher than between tomographic and geodynamic models (which will be discussed below; cf. Becker and Boschi, 2002).

# 2.3 Ways of comparison between geodynamic and tomographic models

<sup>15</sup> We assume here that S-wave velocity anomalies are linearly related to the density anomalies we infer from our slab models, implying that all chemical heterogeneity is captured by the chemical piles in the lower mantle. This assumption ignores the potential complexities which may arise from mineral physics and thermodynamic considerations of the shear wave to density anomaly scaling in a heterogeneous mantle (e.g. Ricard et al., 2005; Stixrude and Lithgow-Bertelloni, 2010).

We compare geodynamic and seismic models both visually along cross sections and in map view and formally in terms of correlation coefficients. Formal correlation is computed and displayed in the same way as in Becker and Boschi (2002): that is, we plot or give global correlation,

- as a function of depth and spherical harmonic degree;
  - as a function of depth for all spherical harmonics up to I = 8 or I = 20;





- and for the entire mantle up to / = 8 or / = 20.

Correlation is a convenient, but perhaps not always the most meaningful way of comparison between models (e.g. Ray and Anderson, 1994). For example, if the predicted location of a subducted slab is only slightly higher or lower than observed through tomography, or – in other words – at the same depth is slightly offset to the side, correlation may be low, although the pattern may be quite similar. This issue has been addressed in Becker and Boschi (2002) by vertically stretching subduction models.

Besides this formal comparison, we therefore here also give a visual comparison in map view and along cross sections, which can also give an indication of how predicted and observed slabs are offset, and what kind of model modification could improve the fit.

#### 3 Results

10

#### 3.1 Sinking speeds and lateral displacement of slabs

Results regarding modeled slab sinking speed and lateral displacement are summa-<sup>15</sup> rized in Fig. 3. The left panel shows an average sinking speed of ~ 2 cm yr<sup>-1</sup> in the lower mantle. However, it also indicates variability: for example, at 60 Ma, most slabs are in a depth range 1700–2300 km, indicating sinking speeds varying between ~ 2.2 and 3.5 cm yr<sup>-1</sup> in the upper part of the lower mantle. Somewhat lower sinking speeds (average ~ 1.6 cm yr<sup>-1</sup>) are computed if we restart the computation at 300 Ma with the <sup>20</sup> computed present-day density model. This might occur because the lower mantle beneath subduction zones already contains cold material from the previous model run, so the density difference between newly subducted slabs and surrounding mantle is less, leading to slower sinking.

The center and right panel show that most slabs sink close to vertical. For example, at 1850 km depth, the peak of the histogram is at only 1 degree of arc, and the larger amount of slabs get advected less than 2.4 degrees. Only at the very base of



the mantle, older slab particles will obviously move horizontal above the CMB, and distances get larger.

# 3.2 Visual comparison between geodynamic and tomography models

# 3.2.1 Map views of slabs in the entire lower mantle

25

- <sup>5</sup> Figure 4 shows the distribution of slabs predicted by our forward model and displayed in analogous way to the tomography model in Fig. 2. On the positive side, we note that there is an overall agreement between regions where slabs occur in both models. In particular, the agreement is quite good in the lowermost mantle (violet colors). Of course, this does not come as a surprise, as slabs in our model do not move very
   <sup>10</sup> much laterally until they reach the lowermost mantle (see Sect. 3.1), and the agreement of subduction zone locations through geologic history (Fig. 1) and tomography of the lowermost mantle (Fig. 2) has been noted early on (e.g. Richards and Engebretson, 1992). On the downside, the maps look quite different on a smaller scale: the tomography model looks more "blobby" whereas the geodynamic model generally
   <sup>15</sup> shows linear anomalies often connected to surface subduction zones and getting more
- shows linear anomalies often connected to surface subduction zones and getting more diffuse and less strong further down in the mantle, before spreading out above the core-mantle boundary. It is also not straightforward to "match" individual slabs in both models. Part of the problem is that we have to choose a certain cutoff in both figures, and the figures change with cutoff, which is why we also compare the models along cross sections in Sect. 3.2.2. But it is also clear that trying to match slabs inferred from
- <sup>20</sup> cross sections in Sect. 3.2.2. But it is also clear that trying to match slabs inferred from subduction history and tomography requires a very dedicated effort (e.g. van der Meer et al., 2010).

The dependence of fit on wavelength will be more formally discussed in Sect. 3.3; expectedly, the agreement is looking much better if you step back and don't look at details.





#### 3.2.2 Cross sections

While map views as presented in Sect. 3.2.1 enable an overall comparison, the problem of having to choose a certain cutoff, and the fact that slabs higher up in the mantle hide slabs beneath makes cross sections more suitable for a detailed comparison try-

<sup>5</sup> ing to match slabs in both models. Such a comparison of cross sections through the mantle beneath subduction zones of the past 300 Myr is given in Fig. 5. While the cross sections beneath the most of the Americas, Australia and Antarctica mostly cross one subduction zone for a given time, those beneath Indonesia, Eurasia and Alaska typically cross two – one in the circum-Pacific Ring and one corresponding to Tethys,
<sup>10</sup> Mediterranean and Indian Ocean subduction at the southern margin of Eurasia. Cross sections are shown for model st12den-2. Cross section for s12den-1 have more diffuse slabs – as the slab input is distributed over a larger depth interval – but otherwise look similar.

Tomographic and geodynamic cross sections show overall agreement on the large scale, with – in most cases – the central parts of the cross sections being dominated by cold or seismically fast material that can be attributed to subduction, while on one (240–285°), or both (90–225°) of the sides often hot or seismically slow material appears, due to upwellings that may occur along the margins of the Large Low Shear Velocity Provinces (LLSVPs; Garnero and McNamara, 2008). A different pattern occurs under

- <sup>20</sup> Asia, with subduction and corresponding slab anomalies (Circum-Pacific and Tethys) on either side, and a slow anomaly (seen in the tomography cross section especially at 345°, in the model cross section at 345° and 330°) in between. A slow anomaly in the lowermost mantle beneath Russia and Kazakhstan appears in most recent tomography models and seems to be a robust feature. The tomography cross section at 345° and
- <sup>25</sup> qualitative agreement with modeling results (see also Steinberger and Torsvik, 2012) indicate that it may well be overlain by a mantle plume.





A more detailed comparison of slow regions, except for those corresponding to the LLSVPs of the lowermost mantle (see Sect. 3.1.3) is not attempted here, for the following reasons:

a. Our code does not consider lateral viscosity variations, and therefore our resulting upwellings are probably unrealistically wide.

5

- b. A statistical comparison of seismically slow regions and predicted plume conduits from geodynamic modeling has already been shown to display good agreement for some deep-rooted plumes being connected to hotspots (Boschi et al., 2007, 2008).
- c. Although we generally find a pattern of plumes along LLSVP margins (Steinberger and Torsvik, 2012) their locations do not exactly correspond to observed hotspots.
  - d. Individual, thermal plumes may be hard to detect seismically (e.g. Hwang et al., 2011).
- If we attempt to match individual features in the seismically fast regions, we find they are generally less deep in the tomography model. We begin the comparison under South America (135–90°) where it is perhaps most straightforward, and then move counterclockwise around the Pacific before discussing subduction at the southern margin of Eurasia: at 135°, we find a gap (corresponding to smaller amounts of subduction) at a radius ~ 0.7–0.75 (normalized to Earth's radius) in the geodynamic model,
- whereas a similar gap occurs in the tomography model at about 0.73–0.85. At 120°, this gap occurs at larger depth in both cases:  $\sim 0.6-0.7$  in the geodynamic vs.  $\sim 0.63-0.8$  in the tomographic model. Further north, we observe more or less continuous fast or cold material from the top to bottom of the mantle in both cases. However, maxima (corresponding to largest amounts of subduction) occur again at larger depths  $\sim 0.8$  in
- the geodynamic model vs. ~ 0.85 in tomography, at 105–75°. We also find that modeled maxima typically occur ~ 10° further west than seismically observed, which may,





at least in the southern part, be caused by flat slab subduction (Isacks, 1988) not accounted for in our model, where we assume vertical sinking.

Beneath North America (15–60°), disagreement becomes more prominent: still we find maxima generally deeper (~0.65–0.7) in the geodynamic model than in the to-<sup>5</sup> mography model (~0.72–0.8) and further to the west. Again, flat slab subduction (Bird, 1988) could, at least partly, be responsible for this lateral offset. Accordingly, Bunge and Grand (2000) invoked low-angle subduction from a comparison of geodynamic modeling results and tomography. It has been shown to be difficult to achieve agreement between geodynamic and tomographic models in this region without ad-hoc

assumptions such as a stress guide (Liu et al., 2008). However, beneath North America also the shape of the anomaly disagrees. Whereas tomography shows a slab dipping from west to east (especially in the sections at 60° and 45° in Fig. 5), at an approximately constant dip angle, hence indicating an approximately constant slab sinking speed of ~ 1–1.5 cm yr<sup>-1</sup> (Grand et al., 1997), the modeled slab has a more compli cated shape, reaching a maximum depth for slabs of about 80 Ma, whereas older slabs are less deep in the mantle.

Qualitatively, we can understand this shape due to high rates of subduction beneath North America in our model in the late Mesozoic (darker colors in Fig. 1). These not only cause relatively fast slab sinking speeds, but also an upward return flow to the side of it, further enhanced by active upwelling (indicated by red colors beneath the slab in the cross sections) thus hindering the older slabs from sinking further or even pushing

them up again. However, the observation that this shape is not seen may indicate that rates of Farallon subduction beneath North America before ~ 80 Ma were higher than in our model. This could be the case, as the absolute Pacific plate motion is not well

20

known before the age of the Hawaii and Louisville hotspot chains, and hence Circum-Pacific rates of subduction are not well constrained before ~ 80 Ma. Larger amounts of Farallon subduction could also account for the discrepancy between seismically fast material at the base of the mantle at 60° and 45° beneath North America, and the absence of corresponding material in the geodynamic model.





Beneath Alaska and the Aleutians (left part of cross section at 0° and right part at 345°) a larger modeled anomaly centered at radius ~0.8 might correspond to a smaller observed anomaly at radius ~0.82. A reasonably good match is found beneath Kamchatka and especially Japan (right parts at 330° and 315°). The main difference is that the slab beneath Japan appears to stagnate in the transition zone (Fukao et al., 2001) and is therefore observed at a depth shallower than modeled. The fact that the observed slab is further to the west could be due to the slab moving (possibly being pushed by the subducting plate) further westward while it is stagnating.

Further south, disagreement becomes more prominent again: at 300° (Izu-Bonin arc) the model slab is still continuous, whereas tomography actually shows slow anomalies in the mid-mantle, separating fast anomalies below and above. Such a separation between slabs in the upper part of the lower mantle, and in its lowermost part occurs in the geodynamic model only further south, at 285° (Marianas arc). This pattern of fast or cold anomalies mainly in the upper part of the lower mantle and in the lowermost

- <sup>15</sup> mantle in both the geodynamic and tomographic model continues south until 240° and 225° (Tonga-Kermadec). This is also the only region where the slab, as seen in the tomography model, occurs at larger depth than in the geodynamic model. This is likely partly because due to fast spreading in the Lau back-arc basin and resulting trench retreat, actual convergence rates at the Tonga trench are up to 24 cm yr<sup>-1</sup> (Bevis et
- al., 1995). This is much higher than in our simplified plate reconstruction, where we use Pacific and Australian plate motions to compute convergence, and larger amounts of subducted material may lead to faster sinking. A comparatively smaller effect is that in our geodynamic model, sinking is counteracted by upward flow in the lower mantle, whereas the tomographic model only shows further to the east slow wave-speeds
   indicative of upward flow.

Another discrepancy is that further to the west, beneath Australia and the Australian Antarctic Discordance south of Australia, there are considerable amounts of seismically slow material throughout the lower mantle in the tomography model, whereas in the geodynamic model most of the cold material is in the lowermost mantle. Again,





this could be because sinking speeds are too fast in the model, but it has also been suggested that an ancient slab is being drawn up beneath the discordance (Gurnis and Müller, 2003).

In the regions further south, across and near Antarctica (cross sections at 210° and 180°) agreement is good again to the extent that in both models the most prominent fast or cold anomalies only occur in the lowermost mantle, because subduction has terminated in that region at about 80 Ma. At the 180° (south polar) cross section it can also be seen that, once subduction has stopped, the last slabs sink at a considerably slower rate than average shown in Fig. 2: from the computed present-day depth ~ 1300 km and the insertion depth 650–700 km at 66 Ma an average sinking speed of ~ 1 cm yr<sup>-1</sup> is inferred. Given that the tomographic anomaly reaches up to about the same depth, one can infer a similar sinking speed also from tomography.

Slabs subducted at the southern margin of Eurasia can be seen at the right side of the cross sections from 30° to 0° and at the left side from 345° to 270°. We can again mostly match fast anomalies in the seismic with cold anomalies in the geodynamic

- <sup>15</sup> mostly match fast anomalies in the seismic with cold anomalies in the geodynamic cross sections. Again, anomalies in the geodynamic cross sections occur often (such as in the 315° cross section) at greater depth than in the seismic ones. There is no strong lateral offset, except at 270°. In that equatorial cross section beneath Indonesia, the fit is very poor: the geodynamic model predicts cold anomalies throughout the lower
- <sup>20</sup> mantle, whereas the seismic model shows a strong anomaly laterally displaced in the upper mantle and upper part of the lower mantle. This misfit is probably in part because the plate kinematic history in that region is very complicated and not adequately matched by our simplified model.

Finally, cross sections beneath Asia, especially at 330°, show in their central parts
the remains of the Mongol-Okhotsk subduction zone (van der Voo et al., 1999). Like in the other case where subduction has subsequently stopped – beneath Antarctica – the geodynamic model predicts here comparatively slow sinking of the final slabs subducted in that region, such that the geodynamic model predicts cold anomalies at a similar depth to the observed fast seismic anomalies.





#### 3.2.3 Map views of the lowermost mantle

We now compare model predictions with and without a thermo-chemical layer at the base of the mantle. If – as in the model shown in Figs. 4 and 5 – such a layer is included it is shaped into thermo-chemical piles, essentially being pushed away by subducted slabs and piled up beneath the regions where no subduction has occurred since 300 Ma (McNamara and Zhong, 2005). In the map view shown in the top panel of Fig. 6, these piles appear as large regions of negative thermal density anomaly, corresponding to high temperature: they are chemically denser and hence remain near the CMB, where they become hot. These regions correspond in location and shape approximately to the LLSVPs seen in tomography (bottom panel), but are somewhat larger in size. The size could however be adjusted by varying the chemical density contrast. On the other hand, in the purely thermal model, a larger number of smaller anomalies similar in size, and one of them similar in location, to the seismic anomaly seen beneath Russia and Kazakhstan are predicted.

#### **3.3** Correlations between geodynamic models and tomography models

Our more qualitative visual comparison in the previous section has already shown that agreement is best on the largest scales, whereas finer details, if matching pairs can be found at all, are often shifted laterally or radially. The correlation plots in Fig. 7 provide a more quantitative description of this finding: for st12den-1 and st12den-2 the

<sup>20</sup> center panels show that correlation is good throughout the lower mantle until / ~ 4, corresponding to a half-wavelength, or size of anomalies, of ~ 5000 km. However, since lowermost mantle structure is dominated by large-scale structure, this still corresponds to a total correlation ~ 0.7 for both  $r_8$  and  $r_{20}$ , i.e. up to degree / = 8 or 20. In the middle part of the lower mantle, large-scale structure becomes less dominant, and therefore <sup>25</sup> overall correlation is lower, particularly for  $r_{20}$ . From 2000 km up to 670 km, correlation becomes better at gradually higher degree. At 670 km, it is reasonably good until de-





offset of anomalies in the geodynamic vs. seismic model of typically less than that, in accord with the visual assessment in the previous section. Correspondingly, overall correlation in the upper part of the lower mantle is also higher than in its mid-part, particularly for  $r_{20}$ . The right panels show that through much of the lower mantle the

- <sup>5</sup> rms density anomaly of the geodynamic model is about one quarter of the RMS wavespeed anomaly of SMEAN, similar to the value expected if both are due to temperature anomalies (e.g. Steinberger and Calderwood, 2006). In the lowermost mantle, where the assumption of thermal density anomalies holds less well, the difference in rms anomaly becomes somewhat less. Overall, SMEAN is somewhat better correlated with
- st12den-1 than with st12den-2. This is probably partly because slabs in st12den-1 are inserted at shallower depth. Partly it may also reflect that slabs in st12den-1 are inserted over a larger depth interval and therefore more diffuse, and tomography models also tend to be "smeared out".

The overall appearance of the correlation plots remains similar, and total correlation becomes only slightly lower if we replace SMEAN by any of the other mean models, or by S40RTS (Ritsema et al., 2011), or if we replace our geodynamic model by the model of Steinberger and Torsvik (2012) (shown in their Fig. 5) with a less strong decrease in thermal expansivity with depth, compared to st12den-1.

Whereas in models st12den-1 and st12den-2 subduction zones are in a TPW corrected global hybrid reference frame (Steinberger and Torsvik, 2008; Torsvik et al., 2008) additionally corrected in longitude (van der Meer et al., 2010), we find somewhat lower correlations  $r_8 = 0.29$  compared to 0.35 and  $r_{20} = 0.18$  compared to 0.21 without the longitude correction. This is not surprising given that the correction of van der Meer et al. (2010) by construction is meant to optimize the fit between slab locations based on subduction zones and tomography.

Correlation is also substantially reduced throughout the mantle if the model does not include thermo-chemical piles (cf. middle panel of Fig. 6). We attribute that difference throughout the mantle and not only at the depths where the thermo-chemical piles





occur to the fact that upwellings in the model with piles are generated at locations that are less different from regions of low seismic wavespeed.

In contrast, if we include neither thermo-chemical piles nor diffusion of heat across the bottom thermal boundary layer in the model (case st12den-7 in Fig. 7), we obtain

- <sup>5</sup> an even higher correlation except in the lowermost mantle: including diffusion of heat across the CMB improves correlation in the lowermost mantle, because model slabs push hot material toward locations corresponding to LLSVPs. However, higher up in the mantle, correlation gets worse, as upwellings in the model form at locations that generally do not match well with actual upwellings.
- <sup>10</sup> We find that correlations have improved compared to the earlier slab model of Steinberger (2000), which gives  $r_8 = 0.3$  and  $r_{20} = 0.21$  with SMEAN (Becker and Boschi, 2002). In this case, like for our new slabs-only model st12den-7 (top right panel of Fig. 7), correlation remains at a similar level throughout the mantle; and is slightly higher in the upper part of the lower mantle than at its base. Correlations of SMEAN <sup>15</sup> with the simple slab sinker model (vertical sinking at a prescribed speed) of Lithgow-Bertellani and Bisharda (1000) are also similar to correlations with Stainbarger (2000)
- Bertelloni and Richards (1998) are also similar to correlations with Steinberger (2000)  $-(r_8 = 0.33 \text{ and } r_{20} = 0.18)$ . For comparison we have also devised a slab sinker model based on our own subduction model, both in the TPW corrected global hybrid reference frame (Steinberger and Torsvik, 2008; Torsvik et al., 2008) and in a reference frame ad-
- ditionally corrected in longitude (van der Meer et al., 2010) shown in Fig. 1, and vertical sinking of 1.2 cm yr<sup>-1</sup> (van der Meer et al., 2010). We find that despite the update in plate reconstruction model, correlations of the slab sinker model with tomography remain low, on a similar level to the model of Lithgow-Bertelloni and Richards (1998). The fact that the "slabs-only" geodynamic model st12den-7 gives much higher correlations
- than the slab sinker approach emphasizes the importance of mantle flow modeling which predicts (although rather small) lateral advection of slabs, and variable slab sinking speeds (Fig. 3).

Finally, because it appears that including upwellings that dynamically form in our model always deteriorates correlation, we consider a model where, instead, we include





plumes with surface positions based on hotspots, and tilted plume conduits with moving source at the CMB, as in Boschi et al. (2007, 2008), based on the modeling approach of Steinberger and Antretter (2006). This approach is different in that here plumes are treated as essentially passive, not influencing large-scale flow. We find that in this case

<sup>5</sup> (bottom right panel of Fig. 7) correlations are further improved compared to the slabonly model. We note that here the amplitude scaling of plumes and hence the amplitude of the combined model is somewhat arbitrary, but resulting correlations depend on this scaling only slightly. Note, though, that the flow field used to advect plumes in this approach is not based on subduction, but inferred from tomography.

#### 10 4 Discussion

Slab sinking speeds in our model (Sect. 3.1) are significantly higher, by about a factor 2, than the estimate  $1.2\pm0.3$  cm yr<sup>-1</sup> of van der Meer et al. (2010) based on comparing reconstructed subduction zones with tomography. Our own comparison to tomography in Sect. 3.2 essentially confirms that model sinking speeds appear to be too high. Build-

ing upon this comparison we can give our own estimate of what would be appropriate slab sinking speeds to best explain tomography: we identify characteristic features that can be visually matched in the geodynamic model and tomography model cross sections. Based on the slab tracers, we determine the age of slabs corresponding to this feature. We then obtain our own observation-based sinking speed estimate by dividing
 the depth of the feature in the tomography model through this age.

We distinguish between the following three cases: (a) beginning of subduction, or substantial increase in the amount of subduction (e.g. beneath South America); (b) end of subduction (especially Mongol-Okhotsk subduction beneath Asia; Phoenix subduction beneath Antarctica); and (c) specific features in the middle of subduction (the

<sup>25</sup> bend in the slab beneath Japan). Results are plotted in Fig. 8. It appears that most data points plot around a straight line through the origin with a slope of about 2 cm yr<sup>-1</sup>, but with considerable spread, with most data points falling between lines with slope





1.2 cm yr<sup>-1</sup> and 2.8 cm yr<sup>-1</sup>. Exceptions are the two data points corresponding to the Mongol-Okhotsk slab, which would correspond to a much lower sinking speed: however, here our geodynamic model predicts an inverted age progression, with the oldest slabs on top, as subduction at two sides of it – at the southern and eastern margins of Asia – has pushed this slab up again. On the other hand the "225" data point (Tonga-

Kermandec) corresponds to sinking speed higher than 2.8 cm yr<sup>-1</sup>, which could well be caused by the fast convergence rate and corresponding large amount of subducted slab per time and subduction zone length in this region.

The discrepancy of our observation-based sinking-speed estimate with the 1.2 cm yr<sup>-1</sup> determined by van der Meer et al. (2010) is therefore somewhat marginal. We also note that our approach is somewhat biased toward high sinking speeds – at least when interpreting the lower end of a slab, as often slabs are bent in our geodynamic model, such that not the oldest slab is at the lower end.

Our modeled sinking speeds are higher than found by Čížková et al. (2012) for the case where they use a similar viscosity structure. We think that this difference occurs mainly because Čížková et al. (2012) model relatively short episodes of subduction whereas our model typically has subduction in the same region for a long time, leading to larger amounts of subducted slabs, and hence faster sinking.

The models of Shephard et al. (2012) are more similar to ours in that respect, as they are also based on actual subduction history. They find that sinking speeds in the lower mantle do not exceed  $1.5-2 \,\mathrm{cm}\,\mathrm{yr}^{-1}$ , but this difference is probably due to their viscosity being  $10^{23}$  Pas throughout the lower mantle, whereas in our model, such high viscosities are only reached in the lower part of the lower mantle. We find that our correlations between models and tomography are mostly higher than those of Shephard et

al. (2012), in their case with 200 Myr of subduction in the "subduction reference frame", which is their case most similar to our model. This is in part due to our comparison to a different tomography model (for example, we find somewhat lower correlations, if we compare our model with S40RTS rather than SMEAN) and, more importantly, because





we have included a thermo-chemical layer that is formed into "piles" at the base of the mantle.

Figure 6 suggests, at least to us, that the thermo-chemical model – with two large hot regions approximately corresponding to the two LLSVPs – fits tomography better
than the purely thermal one. In addition to the two LLSVPs, the tomography model shows one smaller low shear velocity province beneath north of the Caspian Sea. This feature occurs in many recent tomography models, and therefore appears to be robust. One might consider that if this feature – which is similar in size and location to one of the small hot anomalies in the thermal model – is resolved, tomography would generally resolve a pattern such as in the thermal model in the center panel of Fig. 6. However, more appropriately, the geodynamic models should also be "looked at" through a tomography filter (Megnin et al., 1997; Schuberth et al., 2009; Bull et al., 2009).

Our model provides a significant improvement compared to earlier models – both <sup>15</sup> based on a simple slab sinker approach (Lithgow-Bertelloni and Richards, 1998) and on mantle flow models (Steinberger, 2000). The remaining, and still quite substantial, misfit can help to find out how the model needs to be improved in order to come closer to the ultimate goal of a subduction-based model of mantle evolution that accurately explains present-day tomography.

Firstly, we have tried to match certain features in the tomography model with corresponding features in the geodynamic model, and find that they occur in the geodynamic model generally too deep in the mantle. This might be compensated by assuming an even higher viscosity in the lower mantle. However, an even higher viscosity globally would presumably be difficult to reconcile with geoid constraints (e.g. Steinberger and

<sup>25</sup> Calderwood, 2006). This points towards a possibly important effect of lateral viscosity variations, which are not included in our model: if slabs have been subducted for a long time in a certain region, they cooled the lower mantle, leading to increased lower mantle viscosities in that region and thus slower slab sinking speed, while the global average





viscosity could remain compatible with geoid constraints (Yoshida and Nakakuki, 2009; Ghosh et al., 2010).

Another difference is that in our model, the slabs often appear bent (e.g. in the cross sections beneath North America), such that sometimes older slabs are less deep than <sup>5</sup> younger ones. In contrast, tomography in that region has been interpreted such that slabs sink at approximately constant speed such that a subduction zone migrating at a constant speed could give a slab with constant dip (Grand et al., 1997). Again, this discrepancy could possibly be due to lateral viscosity variations causing slabs to be stiffer than their surroundings: a stiffer slab would be less readily bent.

- Although in general our model slabs are too deep, the opposite case also occurs, namely beneath the Tonga-Kermadec subduction zone. This can be attributed to the fact that the convergence rate in our model is too low, and hence illustrates the importance of considering detailed, regional plate reconstructions. Also, a cross section where the fit is particularly poor is beneath Indonesia, which is also known to be a region of particularly complicated plate tectonic history. Clearly, it would be beyond
- the scope of any single paper to address this problem globally, so improvements here should be made region by region, possibly still within a global model, but with regionally refined plate reconstructions and focusing on a regional comparison.

Besides including more detailed plate reconstructions, it will also be a key issue to start with plate reconstructions further back in time. Our tentative model, where we re-ran starting from the modeled present-day structure for another 300 Myr yielded about 25 % slower slab sinking speeds. We think this occurs because of the accumulation of cold slab material in the lower mantle beneath subduction zones, leading to a reduced density contrast of newly subducted slabs and the surrounding lower mantle,

and thus a reduced sinking speed, even without considering lateral viscosity variations. Including plate reconstructions further back in time could hence yield a similar effect. Furthermore, it could help to keep thermo-chemical piles stable for a longer time: observational evidence indicates that they have been in a stable position since 200 Myr (Torsvik et al., 2006) and possibly much longer (Torsvik et al., 2010) whereas they form





more recently in our model, because subduction is initiated at 300 Ma. But it could be a challenge to keep them stable, particularly if earlier plate reconstructions feature subduction zones between different continents assembling to form Pangea, and these subduction zones directly overly the African LLSVP (cf. Zhang et al., 2010).

- <sup>5</sup> An improved fit could also come from comparison to other tomography models, particularly P-wave models which typically contain better resolved slabs (e.g. Bijwaard et al., 1998; Li et al., 2008). Yet, for example,  $r_8$  correlations of MITP-08 (Li et al., 2008) with st12den-2 is ~ 0.19, which is worse than for SMEAN on a global scale. The SMEAN model, which is an average over three S-tomography models and which we
- <sup>10</sup> mainly compared to here can be seen as a sort of "common denominator" that maintains robust features on global scales. On the other hand, some models may robustly resolve features that are not included in the mean model, by design, particularly on regional scales. Again, as for the use of more detailed plate reconstructions, a careful analysis and comparison should be done region by region.
- <sup>15</sup> Although not the focus of this paper, we note that including upwellings from a basal thermal boundary layer in the dynamic model always worsens the fit. Here we have shown that if positions of upwellings are based on surface hotspots, a much improved fit results. It will be a further challenge to improve models such that upwelling plumes self-consistently form at the right locations.
- <sup>20</sup> Although van der Meer et al. (2010) assume vertical sinking at constant speed, we confirm their approach to the extent that slab input in the reference frame that uses their longitude correction gives a better fit than without longitude correction. Future work should attempt to re-calibrate the longitude correction with a mantle flow based approach similar to this paper.

#### 25 5 Conclusions

We have devised a geodynamic model of the mantle based on 300 Myr of subduction. In the model, most slabs sink to the lowermost mantle in about 120 Myr, while they





typically move only a few degrees laterally. However, such lateral advection and the lateral variation of slab sinking speeds are relevant, and they lead to an improvement in model fit to tomography compared to models with slabs sinking vertically at constant speed. If a chemical layer is included in the model, it yields two thermo-chemical piles

<sup>5</sup> in the lowermost mantle, similar in shape and location to the Large Low Shear Velocity Provinces that are seen in tomography. This model correlates very well with the SMEAN composite tomography model up to spherical harmonic degree ~ 3–4.

Comparison along cross sections shows substantial differences between geodynamic and tomographic model, but allows to match certain "slab" features in either model with each other.

Corresponding features in the geodynamic model appear normally at greater depth than in the tomography model, indicating that modeled sinking speeds are too fast. Through such matching of features, we can obtain an observation-based slab sinking speed estimate of  $\sim 2 \text{ cm yr}^{-1}$ , varying mostly between 1.2 cm yr<sup>-1</sup> and 2.8 cm yr<sup>-1</sup>.

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Discussion Paper SED 4,851-887,2012 A comparison between geodynamic and tomographic **Discussion** Paper models B. Steinberger et al. **Title Page** Introduction Abstract **Discussion** Paper Conclusions References Tables Figures Back Close **Discussion** Paper Full Screen / Esc Printer-friendly Version Interactive Discussion





**Fig. 1.** Subduction history and locations of cross sections shown. The subduction history model is the same as in Steinberger and Torsvik (2012), but displayed in a different and hopefully intuitive way: color represents time of subduction before present, and darkness represents the amount of subducted material per time and per subduction zone length, expressed in terms of convergence rate (CR) times a thickness factor (TF), which accounts for the increase of lithosphere thickness with age. We use TF = (age/80 Myr)<sup>1/2</sup> for age < 80 Myr and TF = 1 for age > 80 Myr. Convergence rates are largely unconstrained before 140 Ma and therefore not considered – see Steinberger and Torsvik (2010) for details. Younger slabs are plotted on top of older ones, corresponding to a "view from above" for slabs sinking vertically at constant speed (cf. Sigloch, 2011).



![](_page_29_Figure_0.jpeg)

**Fig. 2.** A representation of the SMEAN (Becker and Boschi, 2002) composite tomography model that is meant to show the depth and intensity of slab-related anomalies. For each location, we determine local maxima of the S-wavespeed anomaly vs. depth profile, as such maxima may correspond to the centers of subducted slabs. We plot the depth (represented by color) and S-wavespeed anomaly (represented by darkness) of such maxima, if they occur in the lower mantle (depth > 670 km) and exceed 0.5%. If at a given location more than one maximum satisfying these conditions is found, only the upper one or uppermost one is plotted, corresponding to a "view from above".

![](_page_29_Picture_2.jpeg)

![](_page_29_Picture_3.jpeg)

![](_page_30_Figure_0.jpeg)

**Fig. 3.** Left panel: depth of slab tracers vs. time of subduction. Center panel: lateral displacement of slab tracers vs. time of subduction. Right panel: lateral displacement vs. depth. All panels are for model st12den-2; red lines indicate the average, whereas histograms plotted for certain depths or times illustrate variability. Only those slabs subducted since 120 Ma are plotted, since older slabs tend to move laterally in the lowermost mantle, and may get heated up and rise again, and would hence make the picture less clear. Both average and histograms consider variable "mass" of slab tracers, which may differ because of variable convergence rate, age of subducted plate and spacing of tracers along subduction zones (which is kept nearly constant).

![](_page_30_Figure_2.jpeg)

![](_page_31_Figure_0.jpeg)

![](_page_31_Figure_1.jpeg)

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**Fig. 4.** Depth and intensity of slabs as inferred from our geodynamic forward model st12den-2, plotted in an analogous way as the tomography model in Fig. 2: for each location, we determine local maxima of the density anomaly vs. depth profile, as such maxima correspond to the centers of subducted slabs in our model. We plot the depth (represented by color) and density anomaly (represented by darkness) of such maxima, if they occur in the lower mantle (depth > 670 km) and exceed 0.125 %, else similar to Fig. 2.

![](_page_32_Figure_0.jpeg)

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**Discussion** Paper

Fig. 5. Caption on next page.

**Fig. 5.** Comparison of the geodynamic model st12den-2 (first and third columns showing predicted present-day thermal density anomalies) with the SMEAN (Becker and Boschi, 2002) composite tomography (second and fourth columns showing S-wavespeed anomalies) along the cross sections shown in Fig. 1. To make cross sections better comparable, the mean value at each depth layer is set to zero in both cases. Numbers at each cross section pair indicate the azimuth (clockwise from northward) of the cross section, i.e. the left two columns more or less correspond to the left part of Fig. 1, and the right two columns to the right part, and downward in this figure corresponds to southward. The orientation is such that mostly West is left and East is right; specifically in the left two columns, the African hemisphere is left and the Pacific hemisphere is right; in the right two columns it is the other way round. On the cross sections for thermal density anomalies we also plot the computed positions of any slab tracers within 1° of the cross section as black dots surrounded by colored dots. The size of these colored dots corresponds to amount of subducted material per time and per subduction zone length (see Fig. 1) and their color to slab age, i.e. its time of subduction (bottom left color bar).

![](_page_33_Picture_1.jpeg)

![](_page_34_Figure_0.jpeg)

**Fig. 6.** Map views of predicted thermal density anomalies and seismic wavespeed anomalies in the lowermost 100 km of the mantle. Top panel: density anomalies in the model st12den-2 with a thermochemical layer at the base of the mantle, described in Sect. 2 and shown in Figs. 4 and 5. Center panel: same for a model without thermochemical layer. No phase boundary is considered, and slabs are inserted at depths 600 and 650 km 12 Myr after subduction. However the latter two differences only change results in a minor way; the main difference is due to presence or absence of a thermochemical layer. Bottom panel: composite tomography SMEAN (Becker and Boschi, 2002). Again, mean values are set to zero in all cases.

![](_page_34_Figure_2.jpeg)

![](_page_34_Picture_3.jpeg)

![](_page_35_Figure_0.jpeg)

**Fig. 7.** Correlation between geodynamic models and composite tomographic model SMEAN (Becker and Boschi, 2002): the left parts of each panel give correlations  $r_8$  for all degrees up to I = 8 and  $r_{20}$  for all degrees up to I = 20, the center parts of each panel correlation as a function of depth and spherical harmonic degree, the right parts of each panel give the RMS anomaly of density (for geodynamic models) or seismic wavespeed (for tomography models) as a function of depth. Top left: model st12den-1, with input of slabs over a wider depth range (and further described in the text). Bottom left: st12den-2, the model shown in most other figures. Top right: st12den-7, the "slabs only" geodynamic model without thermo-chemical piles and without diffusion of heat across the CMB – otherwise same modeling assumptions as st12den-1. Bottom right: a model where tilted mantle plumes with moving source at the CMB according to Boschi et al. (2007) and based on the modeling procedure described in Steinberger and Antretter (2006) have been added to the slab model st12den-7.

![](_page_35_Figure_2.jpeg)

![](_page_35_Picture_3.jpeg)

![](_page_36_Figure_0.jpeg)

**Fig. 8.** Depth vs. time interpreted by matching specific features of tomography: if at a given cross section (indicated by the number plotted) a specific slab feature can be matched in the geodynamic model st12den-2 and the tomography model, the time of subduction is determined from the slab tracers in the geodynamic model, and plotted against the corresponding depth determined from the tomography model. Colors indicate the kind of feature: red – lower end of slab; violet – upper end of slab, if another slab is following; blue – upper end of slab, if no other slab is following; green – mid-slab feature: the characteristic slab bend beneath East Asia.

![](_page_36_Figure_2.jpeg)

![](_page_36_Picture_3.jpeg)