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Asymmetry of high-velocity lower crust on the South Atlantic rifted margins and implications for the interplay of magmatism and tectonics in continental break-up

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Abstract

High-velocity lower crust (HVLC) and seaward dipping reflector sequences (SDRs) are typical features of volcanic rifted margins. However, the nature and origin of HVLC is under discussion. Here we provide a comprehensive analysis of deep crustal structures

- in the southern segment of the South Atlantic and an assessment of HVLC along the margins. Two new seismic refraction lines off South America fill a gap in the data coverage and together with five existing velocity models allow a detailed investigation of the lower crustal properties on both margins. An important finding is the major asymmetry in volumes of HVLC on the conjugate margins. The seismic refraction lines across the
- South African margin reveal four times larger cross sectional areas of HVLC than at the South American margin, a finding that is in sharp contrast to the distribution of the flood basalts in the Paraná-Etendeka Large Igneous Provinces (LIP). Also, the position of the HVLC with respect to the seaward dipping reflector sequences varies consistently along both margins. Close to the Falkland-Agulhas Fracture Zone a small body
- of HVLC is not accompanied by seaward dipping reflectors. In the central portion of both margins, the HVLC is below the inner seaward dipping reflector wedges while in the northern area, closer to the Rio Grande Rise/Walvis Ridge, large volumes of HVLC extend far seawards of the inner seaward dipping reflectors.

This challenges the concept of a simple extrusive/intrusive relationship between sea-²⁰ ward dipping reflector sequences and HVLC, and it provides evidence for formation of the HVLC at different times during the rifting and break-up process. We suggest that the drastically different HVLC volumes are caused by asymmetric rifting in a simple shear dominated extension.



1 Introduction

A lower crustal zone with high density and high seismic P wave velocity is part of the magmatic "trinity" that characterizes volcanic rifted margins: continental flood basalts, seaward dipping reflector sequences (SDRs) and high-velocity lower crust (e.g., Men-

- ⁵ zies et al., 2002; White et al., 1987; Talwani and Abreu, 2000). There is no set definition for "high velocity" in this context, but here we use a cutoff value of 7 km s⁻¹ as "normal", based on the Vp values of unaltered gabbroic oceanic crust (layer 3: 6.8–7.1 km s⁻¹ after, Mooney et al., 1998). Thus, the high-velocity lower crust (HVLC) has a *P* wave velocity (and density) greater than that of typical oceanic lower crust. In general, HVLC
- at volcanic rifted margins is thought to represent magmatic (gabbroic) intrusions and related cumulate layers (Farnetani et al., 1996; Furlong and Fountain, 1986; Kelemen and Holbrook, 1995; White and McKenzie, 1989; Thybo and Artemieva, 2013). Alternatively HVLCs may represent serpentinized peridotite (O'Reilly et al., 1996) or dense metamorphic rocks (Gernigon et al., 2004; Mjelde et al., 2013). Lower continental crust
 of cratons or shields may have velocities exceeding 7 km s⁻¹ (Rudnick and Fountain, 1995).

Alternatively, it might be speculated that portions of the HVLC form post-rift (Franke, 2013). A common assumption describes a close spatial relationship between the SDRs and HVLC, whereby the HVLC represents intrusive equivalents of the erupted lavas
which form the SDRs (White et al., 2008; White and Smith, 2009; Blaich et al., 2009). HVLC can make up a large part of the total magmatic output along volcanic rifted margins, and as studies in the North Atlantic have shown, variations in size and physical properties of the HVLC in these settings hold important clues to mantle melting scenarios (Fernàndez et al., 2010; Kelemen and Holbrook, 1995; Korenaga et al., 2002;
Ridley and Richards, 2010; Voss et al., 2009; White et al., 2008).

The presence of HVLC bodies along the Southern Atlantic rifted margins is well established from seismic and gravity studies (e. g., Bauer et al., 2000; Blaich et al., 2011; Franke, 2013; Franke et al., 2010; Maystrenko et al., 2013; Schnabel et al., 2008)



but they have not yet been studied for areal extent and rifting process determination. That is the purpose of the study reported here. We provide seismic velocity and gravity models for two new margin profiles in South America and integrate them with five others into a regional interpretation of break-up and magmatism in the South Atlantic. The emphasis is on variations in the size and *P* wave velocities of the HVLC along the South American and African margins, and on their distribution relative to the SDRs. We demonstrate a much stronger development of HVLC bodies on the African margin than on the conjugate margin of South America, whereas the distribution of surface volcanism in the Paraná-Etendeka flood basalt province shows exactly the opposite sense of asymmetry. The development of SDR sequences is roughly symmetrical and there are variations in the relative position of HVLC bodies with respect to the SDRs

which question a simple intrusive vs. extrusive relationship and has implications for the timing of HVLC formation relative to rifting and break-up.

2 Geologic framework

- In order to interpret the origin of HVLC we made an inventory control of the location and composition of the cratons and possible metamorphic rocks, which both were found not in immediate distance. Neoproterozoic to early Paleozoic fold belts at the coast border at older cratonic provinces farther inland. The Dom Feliciano Belt developed at the rim of the Rio del la Plata craton. The coast-parallel Kaoko-Gariep Belt and younger
- ²⁰ Cape fold belt are separated by the Mesoproterozoic Namaqua Mobile Belt from the Kalahari Craton. Metamorphic and magmatic events rework the sediments, which were deposited in the Kaoko, Demara and Gariep belts on the western African margin in the Proterozoic or old Paleozoic. Rift magmatism in the Proterozoic led to the intrusions of felsic magmatism or igneous plutons in the basement, which can be found in form of slivers of granitoids and orthogneisses.

The basement composition of the Kaoko belt exhibits turbidites in a highgrade amphibolite facies in contrast to that of the Gariep belt which contains a oceanic crust



sheet in a low metamorphic grade. This belt also includes mélange, metagreywacke and metabasalts. The Dom Feliciono belt includes three main sequences: a schist belt with metasediments and metavolcanics in greenschist to low amphibolites grade, a granite belt with granitoids and basement inliers of Proterozoic age (Rapela et al., 2011).

The Ventania/Cape fold belt is the result of a northwards directed propagation of Patagonia with Gondwana in Permian and Carboniferous times. Sedimentary Late Paleozoic to Mesozoic strata like the South African Karoo Supergroup or the Colorado Fortin Formation overlie the eroded old fold belts. The break-up of SW Gondwana was long considered to have followed an earlier suture marking the collision of those two cratons but more recent studies of the Neoproterozoic belts on the two margins suggest that the South Atlantic rift followed the position of back-arc basins located to the east of the Pan-African suture (Basei et al., 2005, 2008; Frimmel et al., 2013).

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The South Atlantic Ocean formed as a consequence of the break-up of western Gondwana in Early Cretaceous times (ca. 140–130 Ma). South of the Walvis Ridge-Rio Grande Rise, most of the continental margins of the South Atlantic are of the typical volcanic rifted type, with well-developed SDR wedges and HVLC bodies mentioned above. However, the nature of the margins changes abruptly beyond these limits. North of the Walvis Ridge-Rio Grande Rise, the margin architecture resembles the end mem-

²⁰ ber of a magma-poor margin (Mohriak and Leroy, 2012). HVLC at the West African Angolan margin thus was interpreted as being likely unrelated to break-up (Contrucci et al., 2004). Still another style of margin occurs at the Falkland-Agulhas Fracture Zone (FAFZ) in the south (Fig. 1), where the South Atlantic rifting likely was accommodated by strike-slip displacement to form a margin that is lacking SDRs and other signs of magmatism (Becker et al., 2012; Franke et al., 2007; Koopmann et al., 2014).

In contrast to the rifted volcanic margins in the North Atlantic, the geology and spreading history of the South Atlantic (between the Falkland-Agulhas Fracture Zone and the Walvis Ridge-Rio Grande Rise) is comparatively simple. There are no major ridge jumps and the opposing continental margins have a broadly similar geologic his-



tory. The timing of South Atlantic opening was diachronous, progressing from south to north (Austin and Uchupi, 1982; Blaich et al., 2011; Franke, 2013; Franke et al., 2007; Jackson et al., 2000; Rabinowitz and LaBrecque, 1979; Sibuet et al., 1984; Koopmann et al., 2014). The use of seafloor magnetic anomalies to date this process is compli-

- ⁵ cated by uncertainties in anomaly picks for the southern margins with different interpretations. A recent detailed investigation by Koopmann et al. (2014) proposed that the oldest magnetic anomaly offshore Argentina and South Africa related to oceanic spreading is M9r (ca. 135 Ma). Rabinowitz and LaBrecque (1979) suggested that M9N and M11 (133 and 136 Ma, resp.) are the oldest spreading anomalies whereas others
- ¹⁰ mentioned M7 as the earliest anomaly. There is less debate about spreading anomalies in the northern margin areas of the Walvis Basin where M4 (ca. 130 Ma) is the oldest spreading anomaly. Anomaly M4 can be mapped along the entire margin (Fig. 1), and this will serve in the present study as a time marker for the regime of seafloor-spreading in oceanic lithosphere.
- ¹⁵ Onshore, by far the largest concentration of magmatic activity was in the Paraná-Etendeka province of Brazil and Namibia, where considerably more than 10⁶ km³ of continental flood basalts and related silicic volcanic units were erupted in the time span of about 134–126 Ma (Peate, 1997). We do not discuss the complex topic of magma generation and evolution here, but point out that there is a great deal of information and
- ²⁰ much controversy about the source of magmas in the Paraná-Etendeka province and the role of the Tristan plume in producing them (see Peate et al., 1999; Hawkesworth et al., 1999; Trumbull et al., 2003; Ewart et al., 1998; Comin-Chiaramonti et al., 2011). The distribution of lavas offshore, represented by the SDR wedges, shows that magmatism was not concentrated only in the north but extended along both margins to
- the Colorado-Cape Fracture Zone (Franke et al., 2007; Becker et al., 2012; Koopmann et al., 2014; Gladczenko et al., 1997) (Fig. 1). On the South African margin onshore, mafic dikes of break-up age occur as far south as Cape Town. A comparative geochemical study of mafic dikes along the African margin (Trumbull et al., 2007) demonstrated a north-south decrease in crystallization temperatures by about 150 °C, which is impor-



tant when discussing the variations and possible origin of HVLC in a later section. Finally, widespread but sporadic magmatic activity continued well after break-up (80 Ma and younger) in southern Africa and Brazil (Gibson et al., 1995; Comin-Chiaramonti et al., 2011). The most common expression of this are alkaline intrusions, which are locally numerous (e.g., kimberlite fields) but involve much smaller volumes compared with the Early Cretaceous activity.

3 Geophysical coverage

3.1 Existing profiles and interpretation

The upper crustal structure on both margins is well constrained by multichannel reflection seismic data, and this has been used for mapping the distribution of SDRs and their segmentation along the margins (e.g., Bauer et al., 2000; Franke et al., 2007; Koopmann et al., 2014; Gladczenko et al., 1998). In contrast to the reasonable spatial coverage of seismic reflection data, wide-angle seismic lines are few, especially on the South American margin, which motivated the new studies reported below. To some extent, gaps in the seismic coverage can be compensated by regional gravity interpretations (e.g., Blaich et al., 2009, 2011; Dragoi-Stavar and Hall, 2009; Maystrenko et al., 2013; Franke et al., 2006; Hirsch et al., 2009).

Prior to this study, five velocity profiles were available from wide-angle seismic studies and four of them are on the African margin (Fig. 1). Bauer et al. (2000) presented seismic velocity and gravity models for two seismic refraction traverses of the Namibian margin at 22–24° S (Fig. 1, lines 4 and 5), which show thick bodies of HVLC beneath a broad zone of SDRs (inner wedge, flat-lying flows and outer wedge). The third seismic traverse on this margin crosses the Namibian coastline near the Orange River (Fig. 1, line 6) (Schinkel, 2006). The fourth traverse is located at about 30° S (Fig. 1, line 7) and the seismic velocity profile derived by Hirsch et al. (2009) shows a well-

²⁵ line 7) and the seismic velocity profile derived by Hirsch et al. (2009) shows a welldeveloped body of HVLC below SDRs. Finally, it is worth mentioning for reference that



Stankiewicz et al. (2008) published a seismic velocity profile (Fig. 1, line 8) across the sheared South African margin east of the Cape Peninsula in South Africa. This seismic profile across the Falkland-Agulhas Fracture Zone (FAFZ) shows no evidence of magmatic features at the continent-ocean boundary but there are small HVLC bodies well ⁵ inland, which Stankiewicz et al. (2008) attributed to igneous crust formed during the 180 Ma Karoo event, but may also be interpreted as high-density metamorphic rocks (garnet amphibolite, mafic granulite) in the continental basement (as in Norway, see

Gernigon et al., 2006).

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On the South American margin, Schnabel et al. (2008) identified HVLC underlying SDRs along a traverse of the Argentina margin at latitude 44° S (Fig. 1 line 2). They interpreted this HVLC as magmatic underplating at the Moho with intruded lower crust above.

3.2 New profiles

3.2.1 Seismic velocity

- For this study we have calculated *P* wave velocity-depth models for two previously unpublished seismic lines at the South American margin. Line BGR04-REFR02 (Fig. 1 line 3) crosses the Argentine margin at about 47° S, and line BGR04-REFR01 (Fig. 1 line 1) is located at 35–36° S (Fig. 1). In addition to seven Ocean Bottom Hydrophones (OBH) one three component seismometer was deployed to collect the data (Fig. 2).
- ²⁰ An array of 20 airguns with a total volume 51.2 L was used as seismic source. Refraction line 1 and line 3 had a mean shot distance of 125 m. The seismic *P* wave tomography models were computed using the TOMO2d refraction and reflection travel time inversion routine described by Korenaga et al. (2000). The evaluation procedure started with a very simple model consisting of the bathymetry and a 1-D velocity model
- to the basement as constrained by coincident seismic reflection data. For the resulting velocity model we inverted the first arrivals which covered the sedimentary layers as well as the crust down to the crust-mantle boundary. Primary phases from the re-



fracted waves were observed at all stations, whereas coverage by reflected waves from the crust-mantle boundary (PmP phases) was slightly lower. The Moho as a reflector was sampled every 2 km (black line in Fig. 3). Schnabel et al. (2008) tested the velocity depth ambiguity for a similar data set on the Argentine margin (line 2) which has
a comparable acquisition and processing procedure as compared to line 1 and line 3.

Variations of the weighting parameter had no significant influence on the depth of the Moho.

For the inversion of line 1, we used 3576 refracted travel times. A model resulting from an inversion of these crustal phases (Pg) formed the starting model for a joint refraction/reflection inversion. The final models, shown in Fig. 3, include an inversion of the PmP phases in addition to the Pg phases. Line 1 is constrained by 2531 reflected travel times from the PmP phases and the resulting velocity model has a RMS-misfit for the Pg phases of 42 ms and for the PmP phases of 41 ms.

For the inversion of line 3, 2376 Pg and 727 PmP travel times were used from 8 ¹⁵ common receiver gathers. After 25 iterations the RMS travel time misfit reduced to 40 ms with a corresponding χ^2 of 650 ms. On line 3 the seaward extent of the HVLC cannot be fully constraint due to poor ray coverage between 210 and 250 km distance (Fig. 3b). The eastern parts of both lines show distinct regions of HVLC, with velocities between 7 and maximal 7.5 km s⁻¹ at the base of the crust (Fig. 3).

20 3.2.2 Checkerboard tests and resolution

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We performed checkerboard tests to examine the resolution of the data imaging the crust-mantle boundary. The velocities of the final model were perturbed by variations of ±5% resulting in a checkerboard with cells of dimension 25 km × 9 km for line 1 (see Fig. 4a upper panel), and 30 km × 12.5 km for line 3 (Fig. 4c). Synthetic traveltimes are calculated using this input model and the given source receiver geometry. We assumed an uncertainty of 50 ms in the identification of the travel times and inverted these data to recover the undisturbed anomaly pattern (Fig. 4a, c, lower panel). The results show that the resolution is good at shallow crustal depth. To resolve the artificial



velocity perturbation in the lower crust, we conducted a second test, where we placed three different anomalies (characterized by Gaussian spikes, ellipses with a = 10 km, b = 1.5 km) in the lower crust (Fig. 4b and d, upper panel). The obtained recovery displayed in Fig. 4b) confirms that we are also able to resolve variations in the lower crust $_{5}$ exceeding 10 km horizontal and 1.5 km vertical dimension.

3.2.3 Gravity models

The seismic modeling was complemented and extended by 2-D gravity modeling (Figs. 5 and 6) based on shipboard gravity data, which were processed and modeled using the software GM-SYS. For the density-depth model, structural constraints are given by the seafloor, basement and the Moho from the seismic refraction experiment. These boundaries were fixed throughout modeling. The densities of model blocks represent a mean value following the velocity-density conversion law by Ludwig et al. (1970). In general, two layers were used to represent the Cenozoic and Cretaceous postrift sediments in the Argentine Basin, with densities between 2.15 and 2.4 g cm⁻³. In the area of the shallow continental shelf the thin postrift sediments are modeled jointly as one layer. The crust is represented by three layers: upper continental, lower continental and oceanic crust. Wedges of SDRs were included in the gravity model as a single body. To set the boundaries for the HVLC we followed the 7 km s⁻¹

- isoline in the 2-D *P* wave velocity models. From the resulting deviation of the calculated gravity for models with and without a high density lower crustal body (see red and green lines in Figs. 5 and 6), we conclude that a high density lower crustal body within the velocity range above 7 km s⁻¹ determined by the tomography is necessary to fit the observed gravity values. The dimension of the HVLC bodies is verified by refraction seismics. We can observe a strong lateral gradient in the velocities to values below 7 km s⁻¹ eastwards of the HVLC, where we have a lack of data in line 1 and 3 due
- to low quality. Gravity modeling confirms the extent of the HVLC within the distance range given by the tomography. A refraction seismic line across the Colorado basin also shows that the HVLC is limited to an area close to the continent-ocean transition



and gives evidence that there is no continuation of the HVLC to the west (Franke et al., 2006).

4 Distribution and geometric analysis of the HVLC

HVLC bodies (with Vp > 7.0 km s⁻¹) are identified in all deep seismic lines on the conjugate margins of the South Atlantic (Figs. 7 and 8). However, the HVLC bodies differ greatly in thickness and velocity, and the emphasis of this study is to document and interpret these differences. Useful points of reference are the Rio Grande Rise/Walvis Ridge in the north and the Agulhas-Falkland Fracture Zone in the south (Fig. 1), as well as the seafloor-spreading anomalies M4 and M0 which are mapped on both con-

- jugate margins (Fig. 1). To define the size and seismic properties of the HVLC bodies along the traverses, we divided the velocity profiles into vertical sections with 20 km width and compiled for each section the thickness and the average Vp of HVLC. Due to the use of different modeling approaches for the eastern and western margin we have to differentiate between the respective uncertainties. For the western margin, the
- estimated uncertainty in average Vp estimates is probably better than 0.1 km s⁻¹, and the uncertainty in thickness is on the order of 1 km. All relevant lines on the African margin were modeled with a forward modeling routine. Thus the uncertainties relating the eastern margin are higher.

The South American margin at the latitude of line 3 is classified as magma poor because there are no SDRs and no records of Cretaceous igneous rocks onshore. However, the seismic velocity and gravity models do indicate small bodies of HVLC with a total estimated cross-sectional area of about 120 km², a maximum thickness of 4.4 km and maximum average Vp of nearly 7.3 km s⁻¹ (one segment has a local maximum of 7.6 km s⁻¹). The lack of SDRs above the HVLC body calls into question whether it can be interpreted as a magmatic feature, related to break-up. Farther north along the margin, on lines 2 and 1, the area of HVLC increases to 334 and 586 km², respectively.



The corresponding values for maximum thickness are 4 km and 6 km, and the average Vp values reach 7.4 and 7.3 km s⁻¹, respectively. The seismic profiles in South Africa and Namibia show much greater amounts of HVLC (Fig. 9). The southernmost traverse, line 7 at about 30° S, has a cross-sectional area of HVLC of 1340 km², a maximum thickness of about 10 km and highest average Vp of 7.4 km s⁻¹. The next traverse to the north, at the Orange River (line 6) has well-developed HVLC with a cross-section area of about 1900 km², maximum thickness of 12 km and maximum average Vp of 7.3 km s⁻¹. Finally, the two northern refraction lines in Namibia (4 and 5), about 100 km apart and 500 km north of the Orange River, show thick and broad bodies of HVLC.
The HVLC in line 5 (Mamba-2) has a cross-section area of 2530 km², average Vp of 7.2 km s⁻¹ and maximum thickness of 20.5 km. The HVLC on line 4 (Mamba-1) has an

7.2 km s⁻¹ and maximum thickness of 20.5 km. The HVLC on line 4 (Mamba-1) has an even larger cross-section area of 3240 km^2 , maximum thickness of 18 km and average Vp reaching 7.3 km s⁻¹.

In summary, the HVLC bodies are 2–3 times thicker and about 4 times larger in cross-sectional area on the African margin compared to the South American margin, and this contrast is maintained along the entire N–S extent of the margins. A striking feature of the African margin profiles is a systematic increase in size of the HVLC from south to north (Fig. 9).

5 Discussion

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20 5.1 Origin of the HVLC bodies

For better visibility the HVLC sections shown in Figs. 7 and 8 were separated into two zones: one upper HVLC with velocities ranging between 7 and 7.3 km s^{-1} and a lower HVLC with velocities ranging between 7.3 and 7.8 km s^{-1} . The cratons are separated by orogenic belts from the coast and are located further inland. Thus they are unlikely to be the origin of the HVLC. Along line 7, which crosses the Falkland-Agulhas Fracture Zone, we do not exclude the possibility that the landwards part of the HVLC



represents metamorphic rocks like eclogites as suggested by Mjelde et al. (2013). Stankiewicz et al. (2008) interpreted the landward part of profile 7 as the base crust of the Namaqa Natal mobile belt (NNMB). The NNMB with velocities around 6.8 km s⁻¹ runs towards the north at a distance of 100–150 km landwards of the HVLC. The NNMB ⁵ is very unlikely the origin of the HVLC. Similarly the Cape Fold Belt does not intersect

- profile 7, and can be ruled out as origin for the HVLC (Stankiewicz et al., 2008). Blaich et al. (2011) suggested the accretion of terranes against South Gondwana as the origin for the HVLC on the South American margin. This is proposed to have resulted in the formation of the Cape/Ventana Fold Zone and a northwards directed collisional deformation (Blaich et al., 2009), which on the South American margin may solely 10 affect line 1. if anv.

The next possible explanation is serpentinization of mantle peridotites. This takes place when the crust is thinned to a maximum thickness of 5 km (Mielde et al., 2002), if crustal faults provide pathways for downward fluid migration. We consider such con-

- cept as unlikely in the South Atlantic case for the following reasons. If major serpen-15 tinization took place this occurred before the emplacement of the SDRs, which seal potential deep reaching faults. However the HVLC of the northern lines 1, 4 and 5 developed maximum thickness at the continent ocean transition and reach well into oceanic crust. This implies, first that they contain practically no continental crust and
- second at least for the seaward HVLC a formation that postdate the emplacement of 20 the SDRs. Landward, the vertical extent of the HVLC by far exceeds values for typically serpentinized mantle. We do not see a process that may explain 10 km thick serpentinized peridotites. Further the model of serpentinization fails to explain the observed systematic HVLC variations.

The conjugate margins on the South Atlantic were classified as volcanic because of 25 the emplacement of the SDRs. Due to the close spatial correlation of SDRs and HVLC a magmatic origin for the latter seems likely, with underplating or mafic intrusions as the most probable reasons for the anomalous velocity layer. In its original meaning underplating described the accumulation of magmatic material at the base of the crust.



Nowadays the term "underplating" combines both processes, mafic intrusions and underplating (Thybo and Artemieva, 2013; Mjelde et al., 2002). Bauer et al. (2000) favor sensu strict accreted igneous material over intrusions as an explanation for the HVLC in line 4 and 5. Intrusions related to the Cretaceous Cape Cross complex with velocities

⁵ of 6.9–7.2 km s⁻¹ fail to explain the part of the HVLV with velocities above 7.2 km s⁻¹ (displayed in orange in Fig. 8) and differ in shape from the HVLC bodies. Further, a sharp vertical velocity gradient above the HVLC in line 1 may argue for underplated material against small scale intruded crust.

The HVLC along central lines 2, 6 and 7 extends over the total width of the SDRs.
 Previously, the HVLC was interpreted as a combination of magmatic underplating and heavily intruded crust (Schnabel et al., 2008; Hirsch et al., 2009). A low velocity gradient above the HVLC on line 2 suggests intruded continental crust above the HVLC. Densities exceeding normal crustal values argue for intruded continental crust above the HVLC in the central lines 2, 6 and 7 (Schinkel, 2006; Hirsch et al., 2009; Schnabel et al., 2008) which implies the HVLC being intruded crust.

The lack of SDRs on line 3 and the small thickness of the HVLC make magmatic underplating unlikely for this section. This does not necessarily mean the HVLC is composed of serpentinized mantle. An extreme thin continental crust with a synrift basin above the HVLC, may have eased water entry to serpentinize mantle peridotites. Since

- ²⁰ serpentinization is a gradual process with no clear interface (Mjelde et al., 2002) the presence of Moho reflections argues against the model of serpentinized peridotites. A sharp vertical and lateral velocity gradient surround the HVLC as well as Moho reflections argue for one intrusive body. In case the HVLC of line 3 being magmatic material, the intrusion has not reached the surface to form SDRs. Based solely on seismic
- ²⁵ observations we cannot clearly decide for one hypothesis, magmatic intrusions or serpentinization.



5.2 Relationship of high velocity lower crust and seaward dipping reflector sequences

All margin profiles except line 3 from southern Argentina show well-developed SDR wedges close to the HVLC. Previous studies demonstrate a common pattern of SDR sequences with distinctive "facies" consisting of an inner (landward) SDR wedge interpreted as subaerial lava flows, which is followed by a zone of flat SDRs and commonly an outer wedge that presumably formed in a submarine setting (Planke and Eldholm, 1994). The generation of the SDRs is assumed to be restricted to the break-up process and lies close to the continent ocean boundary (Mutter, 1985; Hinz, 1981). Astonish-

- ¹⁰ ingly, is a distinct seaward shift of the HVLC relative to the SDRs. While in the south the HVLC is situated below the SDRs, towards the north the HVLC formed seaward of the SDRs. The contrast in their distribution across the conjugate Atlantic margins, nearly symmetrically SDR (Koopmann et al., 2014) against asymmetrically HVLC, questions a simple intrusive vs. extrusive relationship between them. The formation of the HVLC
- ¹⁵ bodies seems to be more complex than merely a break-up related feature. This is indicated by the twofold HVLC as found along line 2, which could be explained by a formation during different stages of the rifting and break-up. From the position of HVLC bodies relative to the SDRs we try to infer the approximate timing of the HVLC emplacement.
- The HVLC of the southernmost line 3 was found in continental crust in a SDR free, magma limited environment. This challenges the intrinsic relationship between SDRs and HVLC. On the central lines where SDRs and HVLC are well developed (2, 6 and 7), the HVLC bodies are centered beneath the inner SDR wedge (Hinz et al., 1999; Blaich et al., 2009, 2011). The vertical coincidence suggests that the HVLC bodies are of synrift character. This resembles the classical architecture of volcanic rifted margins
- of synrift character. This resembles the classical architecture of volcanic rifted margins i.e. the heavily intruded crust under the SDRs of the central lines 2, 6 and 7 (Schnabel et al., 2008) may have acted as conduits providing magma for the thick volcanic flows imaged as SDRs (line 2 and line 7). On the northernmost margin profiles (lines 1, 4 and



5) the HVLC is located mainly seawards of the inner SDRs. The HVLC bodies show maximum thickness beneath the inner SDRs wedges and thin slowly with increasing plate separation. This means that all magmatic rocks of the northernmost lines were emplaced close to or after break-up. Indeed, on lines 4 and 5 (Namibia), the HVLC

extends 100 to 150 km seaward as far as magnetic anomaly M4 (130 Ma). This implies that formation of the HVLC initiated at break-up marked by the SDR inner wedge but continued into the phase of margin subsidence and steady seafloor-spreading.

The HVLC in the central segment of the margins appears to have formed near the time of magnetic anomalies M11/M9 (about 136 to 133 Ma), but before M7, which here

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marks early oceanic crust. In the north, the HVLC formation was later than anomaly M7 and continued until anomaly M4, at ca. 130 Ma (see Fig. 9). Thus, the seawards shifting HVLC suggest that the HVLC formation followed the northward migrating South Atlantic rift.

5.3 Asymmetry of magmatism and implications for the break-up process

- ¹⁵ Our study shows a 4-fold difference in the cross-sectional area of the HVLC between the South American and African margins (Fig. 9). A new assessment of SDR distribution and area by Koopmann et al. (2014) reveals a more symmetric distribution, with only 50 % difference in area (0.2×10^3 km² for South America vs. 0.3×10^3 km² for Africa), which is insignificant considering the uncertainties. Interestingly both the HVLC
- ²⁰ and SDR distributions are completely at odds with the onshore record in the Paraná-Etendeka Large Igneous Province, where the magma volumes in South America are on the order of 10 times greater than on the African side. The Paraná volcanic units cover at least 1.2×10^6 km² (Peate et al., 1992). The equivalent units on the African margin cover 0.08×10^6 km² (Erlank, 1984), to which we may add an equivalent area covered by baselting factors divergence are created (Trumbull at el. 2004, 2007). Taking
- ²⁵ by basaltic feeder dikes where lavas are eroded (Trumbull et al., 2004, 2007). Taking maximum thickness values for the Paraná-Etendeka sequences into account (1.7 km in Brazil vs. 0.9 km in Namibia from Peate et al., 1992) yields volumes of 1.4 × 10⁶ km³



and 0.14×10^6 km³ for South America and Africa, respectively, which are likely to be minimum estimates given the advanced state of erosion.

This brings up two important questions: how did the strong asymmetry of magmatism develop during rifting and break-up in the South Atlantic, and why is there such

- a difference in the sense of asymmetry from the offshore and onshore evidence? One point to consider is that the Paraná-Etendeka Large Igneous Province is a feature of the northern end of the margins only and not necessarily related to the Atlantic rifting which started in a magma-poor setting in the south. Franke (2013) points out that there was already a well-established seafloor-spreading system in the southern South At-
- ¹⁰ Iantic during the peak of Paraná-Etendeka volcanism onshore [130 \pm 2 Ma after Peate (1997)]. Nevertheless, the contrast in asymmetry of onshore volcanic rocks and HVLC is a general feature of the margins and must be explained in the context of the break-up process. One explanation for the smaller volume of volcanic rocks on the Namibian margin relative to South America could be a greater extent of post-rift uplift and erosion,
- ¹⁵ but fission-track and denudation studies on both margins do not support this (Gallagher and Hawkesworth, 1994; Gallagher et al., 1994). We suggest that South America possibly offered more favorable structures for magma ascent and extrusion than South Africa.

Potentially, the greater HVLC volumes on the African margin could reflect a misinterpretation of the HVLC as igneous crust. Arguments for an igneous origin of HVLC were given above, but we cannot claim that the HVLC bodies consist entirely of igneous material, especially for their landward border under inner SDR wedges. It seems, however, both unlikely and ad-hoc to suggest that the proportion of continental material in the African HVLC bodies should be many times greater than on the conjugate margin.

Also, if we suppose that less than half of the African HVLC represents mafic intrusions, their seismic velocity would need to be much greater than 7.5 km s⁻¹ to explain the observed average Vp values of the HVLC bodies. Thus, it seems safe to assume that much of the HVLC observed along the South Atlantic margins represents magmatic material of mantle origin. However, there is little good evidence regarding the timing



of HVLC-forming intrusions. A break-up related context for the initial formation is suggested by the coincidence of the landward end of HVLC with the inner SDR wedge on most of the profiles. However, as pointed out above, the HVLC bodies on the northern margin profiles extend for 100–150 km seaward of the inner SDRs, and as far as

- ⁵ spreading anomaly M4. It might be speculated that part of the larger volumes of HVLC on the African margin are related to a longer duration of magma generation and accumulation. Indeed, it is clear from onshore geology that there were post-rift magmatic events in the late Cretaceous (90–70 Ma), late Eocene/early Oligocene (~ 40 Ma) and Miocene (Bailey, 1992, 1993; de Wit, 2007) in southern Africa. The surface expressions
- of post-rift magmatism must have an intrusive equivalent at depth that may contribute to the present volume of HVLC. This would only hold for the landward part of the HVLC. Post rift magmatism as a consequence of uplift cannot be made responsible for the seaward part of the HVLC. Further it fails to explain the systematic decrease in the cross sectional area of the HVLC to the south.
- ¹⁵ Despite many uncertainties of how and when the HVLC bodies formed, it is difficult to avoid the conclusion that much of their volume formed during and slightly after continental break-up, and therefore the reason for their asymmetric distribution should relate to the break-up process itself. Asymmetric rifting with a simple-shear component of stretching offers a mechanism to explain the differences in the HVLC distribution
- ²⁰ and was earlier proposed by Blaich et al. (2011) from analysis of structures across the South Atlantic margins.

Latest rift-related sedimentary basins are confined to the eastern margin, where the Orange, Luderitz and Walvis basins indicate regions of major crustal extension and thinning, in line with the break-up direction. In contrast the major sedimentary rift basins

on the South American margin, the Colorado, the Salado and the San Jorge basins, are oriented perpendicular to the rift axis, which negates a symmetric extension and support the simple-shear mode of extension as already suggested by Blaich et al. (2011) and Koopmann et al. (2014).



According to the simple shear model, rifting is accomplished by displacement along a major detachment, resulting in non-uniform extension of the lithosphere. This process producing asymmetric HVLC distribution occurred before the sea floor spreading, which is symmetrical on both sides.

5 6 Conclusions

Two new refraction seismic models complemented by gravity models fill a gap in the data coverage on the Argentine margin and prove the existence of high velocity lower crust (Vp > 7.0). Combined with the models from several publications a compilation of seven transects allows a comparative analysis of the deep crustal structures and physical properties along the conjugate margins of the South Atlantic. All but one of them (off southernmost Argentina) show a close spatial correlation of the HVLC with the SDR sequences in the upper crust, suggesting they are magmatic features. We cannot totally rule out the possibility that the landwards parts represent metamorphic rocks like eclogites, especially for line 3, but we think that serpentinization is unlikely,
¹⁵ since it fails to explain the systematic HVLC variations. A close spatial relationship of the HVLC to the SDRs, which were used to classify this margin as volcanic, suggests a magmatic origin for the central and porthern lines. That means predominantly sug-

a magmatic origin for the central and northern lines. That means predominantly suggest underplating and mafic intrusions in different combinations as the most probable origin for the HVLC's.

²⁰ Three seismic lines on the South American margin cover the change from a magma poor margin (lacking SDRs and magmatism) in the south to a well-developed volcanic rifted margin off Uruguay in the north and were compared with four transects across the South African margin.

In addition to the volume and the shape the relative position of the HVLC with regard to the SDR sequences varies in a systematic way from south to north. The southernmost small HVLC formed without associated SDRs. In the central sections, the HVLC underlies the inner SDR wedges. However, the northernmost HVLC are located sea-



ward of the inner SDRs. The northern profiles off Uruguay and Namibia show HVLC extending seaward as far as spreading anomaly M4 (130 Ma).

From the seawards migrating position of the HVLC with regard to the inner SDR wedges we infer implications for the formation of the HVLC during different stages of

- the rifting and break-up process. If of magmatic origin, the HVLC in the magma-starved segment was likely formed before break-up. The HVLC in central part of the margins was emplaced contemporaneously with the SDRs i.e. syn-rift. However in the northern-most margin segment the formation of the HVLC started after the emplacement of the SDRs. Thus, a causal relationship between SDRs and HVLC is questioned. The north ern HVLC may have been formed at the end of the break-up process and continued
 - until the earliest seafloor-spreading.

Concerning the distribution of the HVLC we observe an increase in cross-sectional area on the conjugate margins from south to north. Evident is an asymmetry in the cross-sectional area across the margins. The South African margin reveals about four

- times larger and 2–3 times thicker HVLC than the South American margin which stands in contrast to the onshore Paraná/Etendeka flood basalt province that shows the direct opposite picture. We attribute this asymmetry to asymmetric rifting in the simple shear mode. This asymmetry has implications on the magma volume output for the South Atlantic. There may be some analogies to the volcanic margin in the North Atlantic,
- ²⁰ where asymmetric HVLC can be observed at the conjugate East Greenland/Hatton Bank data. Similarly as on the South Atlantic the distribution of the HVLC occurring in continental and oceanic crust north and south of the island hot spot is heterogeneous and an interpretation not straight forward.

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Figure 1. The southern South Atlantic Ocean aligned along magnetic seafloor-spreading anomaly M0 (dashed green line; ca. 125 Ma). The continental margins between the Falkland Agulhas Fracture Zone (FAFZ) and the Colorado-Cape transfer zone are lacking major magmatic extrusives and are indicated as magma-poor. The margins to the north of the transfer zone are volcanic rifted with extensive ward dipping reflector (SDR) wedges and high *P* wave velocities in the lower crust. Black lines show the location of the refraction seismic lines discussed in the text: line 1 (BGR04-01), Line 2 (BGR98-02), Line 3 (BGR04-02), Line 4 (Mamba 1), Line 5 (Mamba 2), Line 6 (Orange River), Line 7 (Springbok) and Line 8 (Agulhas Karoo transect). Magnetic anomalies are shown as green lines. The inlay shows the present-day South Atlantic Ocean with the green areas marking the distribution of the Paraná-Etendeka flood basalts.





Figure 2. Data example for OBH's from line 1 (a) OBH1.5.; (b) OBH1.2. and line 3 (c) OBH2.4; (d) OBH2.1. The reduction velocity is 6 km s^{-1} . Upper panel: seismic sections for OBH's. Middle panel: the panels in the middle represent the picked travel times as gray circles and the calculated travel times as black dots for the final model. Lower panel: the lower panels show the ray-coverage for the stations.



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Figure 3. The resulting *P* wave velocity models according to deep seismic refraction experiments along (a) line 1 and (b) line 3 of the South American margin. The location of the lines is shown in Fig. 1. The white line roughly resembles the 4.8 km s^{-1} isoline and marks the top of the basement, the black line is the Moho. Black points indicate OBH positions.





Figure 4. (**a** and **b**) Results of the checkerboard tests of line 1. (**a**) The upper panel shows the relative velocity perturbation of up to 5% in cells of $25 \text{ km} \times 9 \text{ km}$. In the lower left panel the recovery after 8 iterations is shown. (**b**) The right panels show that even small elliptical anomalies within the lower part of the crust can be resolved with the data. (**c**) and (**d**) Resolution of the checkerboard tests of line 3. (**c**) The upper panel shows the velocity perturbations of 5% in cells of 30×12.5 . (**d**) These panels show the resolution test for elliptical anomalies within the lower panels show the recovery after 8 iterations.





Figure 5. Two dimensional gravity model along northernmost line 1 from the western margin as adopted from velocity modeling. **(a)** Observed ship-borne gravimetric (red) and magnetic (green) data with tentative interpretation of magnetic anomalies M3 and M2. **(b)** Observed (blue) and calculated (red) gravity curves shown for the gravity model. The green line displays the curve for a gravity model without a HVLC (high velocity lower crustal) body. **(c)** Result of 2-D gravity modeling showing a deep crustal transect with a gradually thinning crust-mantle boundary towards the oceanic domain. The HVLC results in thick mafic crust close to the continent-ocean transition before a typical thickness of the oceanic crust is reached at the eastern end of the line. The density values are given in g cm⁻³. Black lines mark density polygons. SDRs seaward dipping reflectors, HVLC High velocity lower crust.





Figure 6. Two dimensional gravity model of the southernmost line 3 from the western margin as adopted from velocity modeling. (a) Observed gravimetric (red) and magnetic (green) data with interpreted magnetic anomalies M2 and M0. (b) Observed (blue) and calculated (red) gravity curves shown for the gravity model. The green line displays the curve for a gravity model without a HVLC body. (c) Result of 2-D gravity modeling showing an abrupt thinning of the continental crust when approaching the oceanic domain. The density values are given in g cm⁻³. SDRs seawards dipping reflectors, HVLC High velocity lower crust.





Figure 7. Interpretation of the crustal-scale lines 1–3 from North (top) to South (bottom) along the South American margin with emphasis on the HVLC. HVLC with *P* wave velocities above 7.0 km s^{-1} are indicated in red, HVLC exceeding 7.3 km s^{-1} is shown in orange. Line 2 is modified after Schnabel et al. (2008). The shape and distribution of SDRs as interpreted from coincident seismic reflection data are indicated as black lines.





Figure 8. Interpretation of the crustal-scale lines 4–7 from North (top) to South (bottom) along the South African margin with emphasis on the HVLC. HVLC with *P* wave velocities above 7.0 km s^{-1} are indicated in red, HVLC exceeding 7.3 km s^{-1} is shown in orange. Line 4 (Mamba 1) and line 5 (Mamba 2) are modified after Bauer et al. (2000), line 6 (Orange River mouth) is modified after Schinkel (2006), line 7 (Springbok) is modified after Hirsch et al. (2009). The shape and distribution of SDRs as interpreted from coincident seismic reflection data are indicated as black lines. Same scale as in Fig. 7.





Figure 9. Comparative N–S volumes of HVLC (High Velocity Lower Crust) in the South Atlantic along the crustal transects. HVLC on the South American margin is shown to the left and aligned according to the distance from the Rio Grande Rise (top). HVLC on the South African margins is shown to the right and aligned according to the distance from the Walvis Ridge (top). The timing of the emplacement of HVLC is indicated with respect to magnetic chrons M7, M4, and M2. Southern limit of the SDRs corresponds to the Colorado-Cape transfer zone as indicated in Fig. 1. The sketches to the right show the respective spatial relation of the HVLC to the SDRs at the relative margin position.

