

1 **Variations of the crustal thickness in Nepal Himalayas based on tomographic inversion of**  
2 **regional earthquake data**

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26 **Abstract**

27 We estimate variations of the crustal thickness beneath the Nepal Himalayas based on tomographic  
28 inversion of regional earthquake data. We have obtained a low-velocity anomaly in the upper part of  
29 the model down to depths of 40 to 80 km and proposed that the lower limit of this anomaly  
30 represents variations of the Moho depth. This statement was supported by results of synthetic  
31 modeling. The obtained variations of crustal thickness match fairly well with the free-air gravity  
32 anomalies: thinner crust patterns correspond to lower gravity values and vice versa. There is also  
33 some correlation with magnetic field: higher magnetic values correspond to the major areas of  
34 thicker crust. We propose that elevated magnetic values can be associated with more rigid segments  
35 of the incoming Indian crust which cause more compression in the thrust zone and leads to stronger  
36 crustal thickening.

37

38 **1. Introduction**

39 Collision processes are related to the convergence of continental blocks and lead to  
40 significant shortening and thickening of the crust. The collision zones with strong seismic  
41 activity often coincide with highly populated areas, leading to damage and destruction of human  
42 habitation and suffering of population. The Himalayas, which are the highest mountain chain on  
43 the Earth, has been formed due to the collision of the Indian and Asian plates. The mechanisms  
44 of mountain building in Himalayas and Tibet are extensively discussed by many authors for  
45 decades (e.g., [Dewey and Bird, 1970](#); [Seeber et al., 1981](#); [Molnar and Tapponier, 1975](#); [Allegre  
46 et al. 1984](#)). As the Indian landmass moved northwards the sedimentary piles with its older  
47 crystalline foundation complexly folded, faulted and thrust, that caused varied crustal structure  
48 all along the 2500 km long Himalayan arc from west to east. According to the most popular  
49 tectonic model of Himalayan collision ([Seeber et al., 1981](#)), the Indian plate underthrusts the  
50 Asian plate along a gentle north dipping (4-10° N) detachment plane, called the Main Himalayan  
51 Thrust (MHT) ([Figure 1](#)). Most of the Himalayan earthquakes are shallow and occur at 15 to 20

52 km depth on the MHT. Several recent seismological studies, however, suggest that the tectonic  
53 model varies from west to east. For example, earthquakes in the eastern Himalaya tend to be  
54 much deeper than in the western part (Kayal, 2001 and 2010; Mukhopadhyay and Sharma,  
55 2010). More definitive geodynamic concepts can only be constructed based on reliable  
56 information on the deep structures in the crust and the mantle. However, due to many political  
57 and natural reasons the Himalayas is a difficult region to make a detailed study with most of  
58 geophysical methods.

59 The Mohorovicic (Moho) discontinuity depth is one of the key types of information  
60 which is directly linked with the major geodynamical processes. For the Himalayas, the existing  
61 Moho depth models are either too generalized or too fragmentary. For example, an existing  
62 global model CRUST2.0 (Bassin et al., 2000) provides an over smoothed solution in Himalayas  
63 with extrapolation in some parts. The gravity modeling in the Himalayas also provides fairly  
64 smooth variations of the Moho depth (e.g., Tenzer and Chen, 2014). Another gravity study by Jin  
65 et al., (1996) reported that the Moho depth varies from 38 km below Indo-Gangetic Alluvial  
66 Plains (IGAP) to about 75 km below high Himalayas. The regional tomography models though  
67 depict reliable images of the lithospheric behavior beneath Himalayas and Tibet (e.g., Li et al.,  
68 2008; Koulakov, 2011), but they cannot provide much constraints on the crustal structures. On  
69 local scale, the existing receiver function sites, deep sounding profiles and local earthquake  
70 tomography results from east to west of the Himalayas and Tibet (e.g. Kind et al., 2002; Kumar  
71 et al., 2005; Galve et al., 2002; Hauck et al., 1998; Mitra et al., 2005; Ramesh et al., 2005;  
72 Schulte-Pelkum et al., 2005; Rai et al., 2006, Mukhopadhyay and Sharma, 2010) provide  
73 reliable, but too local and sparse information which is hard to be used to build a generalized  
74 crustal model for the entire Himalayas.

75 In this paper we make an attempt to estimate the variations of the crustal thickness based  
76 on results of tomography inversion using the travel time data recorded by the networks of Nepal

77 and northern India. In most cases, seismic tomography is used to derive smooth velocity  
78 distributions and appears to be not sensitive to sharp first-order interfaces. However, in some  
79 cases it can provide useful information to estimate the variations of the main interfaces. For  
80 example, [Koulakov and Sobolev \(2006\)](#) provided the map of the Moho depth beneath the Middle  
81 East area based on the inversion of the regional travel time data from the ISC catalogue. This  
82 model is fairly corroborated by later studies based on receiver functions and active seismic  
83 profiles ([Mechie et al., 2013](#)). [Koulakov and Sobolev \(2006\)](#), however, put forward some  
84 conditions which would make possible studying the Moho depth using travel time tomography:  
85 (1) stations in the study area should be distributed densely and uniformly as much as possible;  
86 (2) sufficient amount of sources should be located inside the study region; (3) size of the area  
87 should be in the range of 150-500 km; (4) both travel times of crustal (Pg, Sg) and mantle (Pn,  
88 Sn) rays should be presented in the dataset. To some extent, all these conditions are fulfilled in  
89 the Nepal Himalayas region. Thus, we claim that the tomographic results in this study provide  
90 new information on the variations of the Moho depth beneath the Nepal Himalayas.

91

## 92 **2. Data analysis and tomography model**

93 We have combined the data of regional networks in northern India (run by India  
94 Meteorological Department, IMD) and Nepal (run by the Department of Mines and Geology,  
95 Nepal, DMN) along with the global ISC catalogue for the years of 2004-2014. In total, we used  
96 the information from 78 seismic stations installed in India and Nepal. The data selection was  
97 based on three criteria: (1) the residuals for the P- and S data after location of sources in the 1D  
98 model should not exceed 2 s and 3 s, respectively; (2) the number of picks per event should not  
99 be less than 8; (3) the distance from an event to the nearest recording station should not be more  
100 than 250 km. In total, 10864 P- and 5293 S arrival times from 821 events in the study region  
101 were selected for this study (on average, almost 20 picks per event). The distributions of stations

102 and selected events used for computations are shown in [Figure 1](#). Note that only in Nepal we  
103 have fairly dense distributions of both stations and earthquakes. In China to the north, there were  
104 many events, but no stations were available; in India, there were some stations available, but a  
105 very few events were reported.

106 The analysis of data is performed using the iterative tomographic algorithm LOTOS  
107 ([Koulakov, 2009](#)). Because of the large size of the area, we have modified the code by taking  
108 into account the sphericity of the Earth. All the calculations are performed in the Cartesian  
109 coordinates. However, the reference model is kept radially symmetric, and Z-coordinates for the  
110 events and stations are corrected according to the spherical shape of the Earth. In other aspects,  
111 the workflow of the analysis was similar to that used in other studies based on this algorithm  
112 (e.g., [Koulakov et al., 2010](#)). The processing starts with preliminary source locations with the use  
113 of reference table containing travel times in the 1D model. In the next step, the sources are re-  
114 located using 3D algorithm of ray tracing based on bending method. The velocity distributions  
115 are parameterized with nodes distributed inside the study area according to the ray density. To  
116 avoid any bias of the model due to predefined parameters of the grid, we performed the  
117 inversions for four different grids with different basic orientations. [Examples of node](#)  
118 [distributions for two grids with basic orientations of 0° and 45° are shown in Figure 2.](#) Note that  
119 [in map view, the node projections look regularly spaced. However, along the vertical lines, the](#)  
120 [number of nodes and spacing depend on the data distribution \(cases with denser node](#)  
121 [distributions are depicted with darker points in Figure 2\).](#) The inversion was performed  
122 simultaneously for the 3D P- and S-velocity distributions, source parameters and station  
123 correction. The matrix was inverted using the LSQR method ([Paige and Saunders, 1982; Nolet,](#)  
124 [1987](#)). The inversion results obtained using differently oriented grids are averaged into one  
125 model which then used to update the 3D model for the next iteration. [In total, for the analysis of](#)

126 both synthetic and observed data, we used three iterations, each of which included the steps of  
127 source locations in the updated velocity models, matrix calculations and inversions.

128 To avoid any predefinition for the Moho depth, we set the reference model without any  
129 sharp interfaces and even without high gradient levels. We defined a constant  $V_p/V_s$  ratio equal  
130 to 1.75 and set the P-velocity values at different depth levels: 5 km/s at -1 km, 6 km/s at 25 km,  
131 7.2 km/s at 40 km, 7.7 km/s at 65 km , 8 km/s at 120 km and 8.2 km/s at 210 km depth. Between  
132 these levels, the velocity was linearly interpolated. Starting velocities for P and S models are  
133 shown in Figure 3 with dotted lines.

134 Unlike the tomography algorithms used by Koulakov and Sobolev (2006) for studying  
135 the Moho depth in the Middle East, here we do not parameterize the Moho as a sharp first order  
136 interface with variable depth. Instead, we derive the geometry of Moho based on consideration  
137 of velocity anomalies and absolute velocities. In the starting 1D model the velocity around the  
138 Moho depths was faster than expected crustal velocities and slower than mantle velocities. As a  
139 result, the crust was revealed as low-velocity anomaly, whereas the uppermost mantle is  
140 associated with high-velocity anomaly. The variation of thickness of the crust-related low-  
141 velocity anomaly may represent the perturbations of the Moho depth. After inversion, absolute  
142 velocity forms a zone of higher gradient around the presumed location of the Moho interface, as  
143 seen in representation of average absolute velocities in Figure 3. Note that in this case we present  
144 an average for a long profile, for which the crustal thickness may vary significantly. For local  
145 points, the Moho-related high-gradient zone is seen more prominently.

146 To examine the adequacy of the detection of crustal thickness, we have performed a  
147 series of synthetic tests. In Figures 4A and 5 we present result of one of the tests. The synthetic  
148 model was defined as a superposition of a reference 1D velocity model and a low-velocity  
149 anomaly with the amplitude of -15% of variable thickness representing the crust. The lower limit  
150 of this anomaly is indicated in vertical sections in Figure 5 with solid lines. In map view (Figure

151 4A), a zone of variable thicker crust is highlighted with dotted line. Outside this zone, the Moho  
152 is set at 30 km depth. Note that starting 1D model for the synthetic reconstruction was different  
153 of the “true” reference models. This represents the realistic situation in the case of observed data  
154 analysis when the true reference model is unknown.

155 To compute the synthetic data, we have used the same source-receiver pairs as in the case  
156 of the real data analysis. The computed synthetic travel times were perturbed with random noise  
157 having average deviation of 0.1 s which enables approximately the same variance reduction as in  
158 the case of observed data inversion. After computing synthetic travel times using the 3D ray  
159 tracer, we “forgot” all information on the velocity distributions and source locations. Then we  
160 performed the full data processing including the steps of source locations. The restored  
161 anomalies at 35 km depth are shown in Figure 4A. In vertical sections in Figure 5, we present the  
162 restored relative anomalies and absolute velocities. It is seen that in Section 1 along Himalayas,  
163 the thickness of the derived low P-velocity anomalies correctly represents the undulations of the  
164 Moho interface in the input model, especially in the eastern part of the profile. We see that the  
165 location of the “true” Moho better corresponds to the velocity of 7.4 km/s (yellow zone) in the  
166 restored absolute velocity model.

167 For the Sections 2 to 5, the low-velocity anomaly is visible only beneath Nepal. Neither  
168 in the Indian nor in the Tibetan side, the crust-related anomaly is restored. This test shows that  
169 the robust reconstruction of the Moho depth using the tomographic reconstruction can be  
170 achieved only in case of coexistence of stations and events in a sufficiently large area. Just  
171 availability of only stations (like in India) or only seismicity (in Tibet) is not enough for this  
172 purpose. The resolved area can be estimated from this test that roughly coincide with the shape  
173 of the restored crust-related anomaly.

174 The horizontal resolution is examined with another synthetic test which is presented in  
175 Figure S1 of Supplementary materials.

176 As was earlier mentioned, for the inversion of the observed data, we used three iterations.  
177 During the inversion procedure, the residuals reduced from 0.807 s to 0.500 s for the P-data  
178 (37.99% of reduction) and from 1.57 s to 0.83 s for the S-data (47.14% of variance reduction).  
179 Note that the similar remnant residuals were obtained in the final synthetic model discussed  
180 above.

181 The results of tomographic inversions for P-velocity are presented in five vertical  
182 sections: one along and four sections across Himalayas. We present both relative deviations in  
183 respect to the starting model (Figure 6) and absolute velocities (Figure 7). In addition, one  
184 horizontal section of P-velocity anomalies at 35 km depth is presented in Figure 4B. More  
185 horizontal sections, as well as S-velocity anomalies, are presented in supplementary materials in  
186 [Figure S2](#). We have defined the resolved area according to the results of various synthetic tests,  
187 mainly based on the capacity to retrieve crustal related anomalies in the test shown in Figure 4.  
188 Areas with lower resolution in sections 2 to 5 are shaded in the resulting plots in Figures 6 and 7.

189 As we see from the results of the synthetic test, thickness of the low-velocity anomaly  
190 beneath Nepal may represent the depth variations of Moho. We can also identify Moho in a zone  
191 of generally higher gradient of the absolute velocity, which is observed in plots of absolute  
192 velocities in Figure 7. Based on the results of synthetic modeling, we have identified the Moho  
193 depth approximately at the contour line of 7.4 km/s (yellow zones in Figure 7).

194 We have manually traced the lower limit of this low-velocity anomaly in 21 vertical  
195 sections passing across the Himalayan chain (see [Figure S3](#) in supplementary) and created the  
196 2D surface of this limit beneath Nepal ([Figure 8A](#)). Projections of this surface to the vertical  
197 sections below Nepal, where a satisfactory resolution is achieved, are shown in [Figures 6 and 7](#)  
198 with solid lines. It should be noted that the unambiguous tracing of the Moho is not possible  
199 everywhere. For example in Sections 5 to 7 below the main low-velocity anomaly, there is  
200 another low-velocity pattern which appears to be weaker and separated from the upper one. This

201 transitional anomaly may or may not be included to the crust. In case it is included, the total  
202 thickness of the crust in the frontal zone beneath Himalayas may reach 80 km (see [Figures S3](#)  
203 [and S4](#) of supplementary materials) that seems to be less plausible. Thus the latter model may be  
204 accepted as more realistic model; the former one is shown in supplementary. Thick transitional  
205 zone in this area may represent underthrusting of one continental block underneath another one  
206 which results at doubling of the crustal thickness. However, as shown by synthetic tests, the  
207 resolution in this part of the area is not high; thus, we should be prudent and avoid too  
208 speculative interpretations.

209 In sections 4 and 5, we can compare our results with Moho depth determinations obtained  
210 in other studies. In Section 4, dotted line depicts the Moho depth derived from combined  
211 interpretation of receiver function, gravity and thermomechanical modeling by (Hetényi et al.,  
212 2006). Same as in our results, they depict flat Moho in the area where our model is resolved. To  
213 the north, their model shows gradual thickening of the crust, however this trend is observed  
214 outside the resolved area of our model. In Section 5, we can compare with the Moho depth  
215 determined from receiver functions by (Schulte-Pelkum et al., 2005). This model shows  
216 generally same dipping trend as in our result. In both profiles, Moho interface in our results  
217 appear to be deeper to 5-7 km than one determined by other authors. We should emphasize that  
218 the absolute values of Moho depth derived from tomography should be considered with prudence  
219 because of unambiguity of the conversion of continuous seismic anomalies into the interface.  
220 The relative variations of the crustal thickness, however, appear to be correct.

221

### 222 **3. Discussion**

223 The variations of the crustal thickness in the frontal zone of the Himalayan thrust belt, as  
224 seen in our tomographic model ([Figure 8A](#)), may be attributed to variable mechanical properties  
225 of the collided plates. The existence of weaker or more rigid segments in the underlying Indian

226 plate may cause weaker or stronger folding in the Himalayan thrust zone. However, due to  
227 several reasons, it is not easy to quantify this correlation because the Indian plate is mostly  
228 covered by thick sediments of the Gangetic alluvium, which hide major tectonic features.

229 To identify hidden crustal structures, the observations of potential fields might be useful.  
230 In [Figure 8B](#) we show the free-air gravity anomalies for the Nepal Himalayas region extracted  
231 from the global model by [Andersen et al. \(2010\)](#); the smoothed anomaly is obtained using a  
232 Pseudo-Gaussian weight function with the characteristic radius of 10 km. In the Nepal  
233 Himalayas, the gravity field demonstrates very strong variations. To the south of the Himalayas,  
234 there is a strong negative anomaly, which is partly caused by isostatic compensation related to  
235 the mountain growth, and it might also reflect thick sediments of the Gangetic alluvium brought  
236 from the Himalayas due to very fast erosion. The maximum value of the free-air gravity field is  
237 observed in the higher Himalayas along the Nepal - China border. It is clear that these strong  
238 variations across the Himalayan thrust zone are mostly due to abrupt Moho dipping from  
239 relatively thinner crust in the Indian Plate to almost doubled crust beneath the Himalayas and  
240 Tibet. Along the Himalayas we also observe strong variations of the gravity anomalies which  
241 might be associated with laterally inhomogeneous thickness of the crust. The lateral variations in  
242 gravity anomalies in Nepal correlate rather well with our estimates of the crustal thickness. For  
243 example, areas of thinner crust indicated with “1”, “3”, “6” and “8” correspond to lower-gravity  
244 anomaly patterns. On the contrary, thicker crust segments numbered with “2”, “4”, “5”, “7” and  
245 “9” are associated with higher values of gravity anomalies.

246 Here we also examine the magnetic anomalies extracted from the global compilation by  
247 [Maus et al. \(2009\)](#). Besides the map for the Nepal Himalayas and adjacent areas in [Figure 8c](#), we  
248 present the map of magnetic anomalies for a much larger area in [Figure S5](#) of supplementary. In  
249 the Nepal Himalayas, the correlation of crustal thickness with the magnetic anomalies is not as  
250 clear as found with the gravity map (partly due to non-availability of high-resolution magnetic

251 data in Nepal). However, it is worth noting that the largest pattern of thinner crust “6”  
252 corresponds to the negative magnetic anomaly. On the contrary, the positive anomaly “4” in the  
253 frontal thrust zone is located close to the positive magnetic anomaly to the south.

254         Based on the observed correlation of the seismic model with gravity and magnetic  
255 anomalies, we propose a mechanism which may explain the variability of crustal thickness along  
256 the Nepal Himalayas. As observed in regional magnetic map of India ([Figure S5 of  
257 supplementary](#)), the crust of the Indian plate appears to be much heterogeneous. One of the  
258 reasons for strong magnetic anomalies may be the presence of large provinces affected by relict  
259 igneous processes. Although these occurred in the geological past, the large magmatic intrusions  
260 might considerably strengthen the crust. The negative magnetic anomalies, on the other hand,  
261 may be explained by thicker sediments. In case of collision, these two types of the crust behave  
262 differently and cause different mechanical effect. The crust affected by igneous processes is  
263 stronger and thus the compression of the overlying crust in the collision zone would be more  
264 prominent. This would explain the thicker crust in anomaly “4” close to the contact area, which  
265 can be explained by stronger pushing effect of more rigid partition of the Indian plate. The  
266 presence of southward curve of the mountain limit line in front of the thicker crustal pattern “4”  
267 may represent a broader shortening zone produced by the more rigid incoming block. The  
268 segment of the anomalously thinner crust “6” may be explained by smaller compression rate of  
269 the crust because of weaker incoming crust and lubricating effect of thicker sediments having  
270 lower magnetic effect. In this case, the mountain front line is curved northward indicating less  
271 intensive shortening.

272

#### 273         **4. Conclusions**

274         In most tomographic studies, the main target is the smooth distribution of seismic  
275 properties which is not sensitive to the geometry of first-order interfaces. In this study we made

276 an attempt to reconstruct variations of the Moho interface beneath the Nepal Himalayas where  
277 more or less uniform distribution of stations and sources take place and travel times of both  
278 crustal (Pg, Sg) and mantle rays (Pn, Sn) are available. Based on synthetic modeling, we found  
279 that for the most of the Nepal Himalayas area, the crustal thickness variations can be robustly  
280 retrieved. For the surrounding areas, like northern India and Tibet, crustal structures cannot be  
281 resolved with the available data.

282         The obtained crustal thickness varies from 40 to 75 km along the Nepal Himalayas. There  
283 is a fair correlation of the derived crustal structures with the observed gravity and magnetic  
284 anomalies. The areas of thicker crust are associated with higher values of the free-air gravity  
285 field and vice versa. This correlation is a good argument to prove the reliability of our findings.  
286 The magnetic anomalies may provide important information on the mechanical properties of the  
287 crust. We see that different segments of the Indian crust behave differently leading to various  
288 collision rates. We expect that thicker crust in the frontal thrust zone can be associated with the  
289 more rigid incoming crust. Weaker crust segments may penetrate underneath overlying plate  
290 with less resistance, and thus the weaker compression rate leads to thinner crust in the frontal  
291 thrust zone. The presence of thick sediments may have a lubricating effect and thus may also  
292 reduce the shortening of the crust.

293         This study gives us a fair understanding of the Moho configuration beneath the central  
294 Himalayas. However, western and eastern parts of the Himalayas are not yet well studied. Such  
295 comprehensive study based on joint consideration of seismic, gravity and magnetic data for the  
296 entire Himalayas will make possible better understanding the mechanisms of the India – Asia  
297 collision.

298

299         **Authors contribution:**

300 J. R. and S. M. provided seismic data and performed its preliminary analysis and  
301 preparation. G. M. with help of A. J. performed all tomographic calculations. A. V. provided  
302 necessary data and information on magnetic and gravity fields used in the paper. J. R. K., S. M.  
303 and I. K. provided geodynamical interpretation of presented results. A. J. prepared presented  
304 graphic materials. I. K. prepared manuscript with contributions of all co-authors.

305

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404

405 **Figure captions:**

406 Figure 1. Map of the study area and data distributions. Background is topography. Yellow dots  
407 are the earthquakes, and red triangles are the stations used in this study. Locations of three  
408 profiles used for visualization of the results are shown. Blue lines indicate the Main  
409 Central Thrust (MCT) and the Main Boundary Thrust (MBT). IGAP is the Indo-Gangetic  
410 Alluvial Plains. Inset shows the location of the region.

411 Figure 2. Distribution of P-rays and two parameterization grids corresponding to basic  
412 orientations of 0 and 45 degrees. Intensity of the points represents the number of nodes in  
413 Z-direction corresponding to the current XY coordinates.

414 Figure 3. P- and S-velocity versus depth. Dotted lines depict starting 1D velocity models, solid  
415 line represents average velocities along section 1 for the main result shown in Figures 6-7.

416 Figure 4. Velocity anomalies at 35 km depth after inversion of synthetic (A) and observed data  
417 (B). For the case of synthetic modeling, the limits of “thick crust” are marked with dotted  
418 line. Triangles denote stations. Locations of profiles presented in Figures 3 to 5 are shown.

419 Figure 5. Synthetic test with reconstruction of the “variable Moho” model in relative anomalies  
420 (upper part) and absolute P-velocities (lower part). Locations of the profiles are shown in

421 Figure 2. The configuration of the synthetic “Moho” is indicated with dotted line. Vertical  
422 lines with numbers mark locations where sections cross each other.

423 Figure 6. Vertical sections of the resulting P-velocity anomalies. Locations of sections are  
424 indicated in Figure 4. Above each section, exaggerated topography is shown. Vertical lines  
425 indicate locations where sections cross each other. Areas with poorer resolution in sections  
426 2 to 5 are shaded. Moho interface (black line) is traced on the bottom of the low-velocity  
427 anomaly. Dashed line in Sections 1 and 2 indicates an alternative interpretation which is  
428 less plausible. Dotted lines in Sections 4 and 5 indicate Moho depth determinations from  
429 (*Hetényi et al., 2006*) and (*Schulte-Pelkum et al., 2005*), respectively. More sections are  
430 shown in Supplementary.

431 Figure 7. Same as Figure 6, but for absolute P-velocity.

432 Figure 8. Map of estimated Moho depth beneath the Nepal Himalayas (A) together with free-air  
433 gravity anomalies (*Andersen et al., 2010*) (B) and magnetic anomalies (*Maus et al., 2009*)  
434 (C). Red numbers indicate the locations discussed in the text. In Plot A, the locations of  
435 profiles used for presenting the main results are shown.

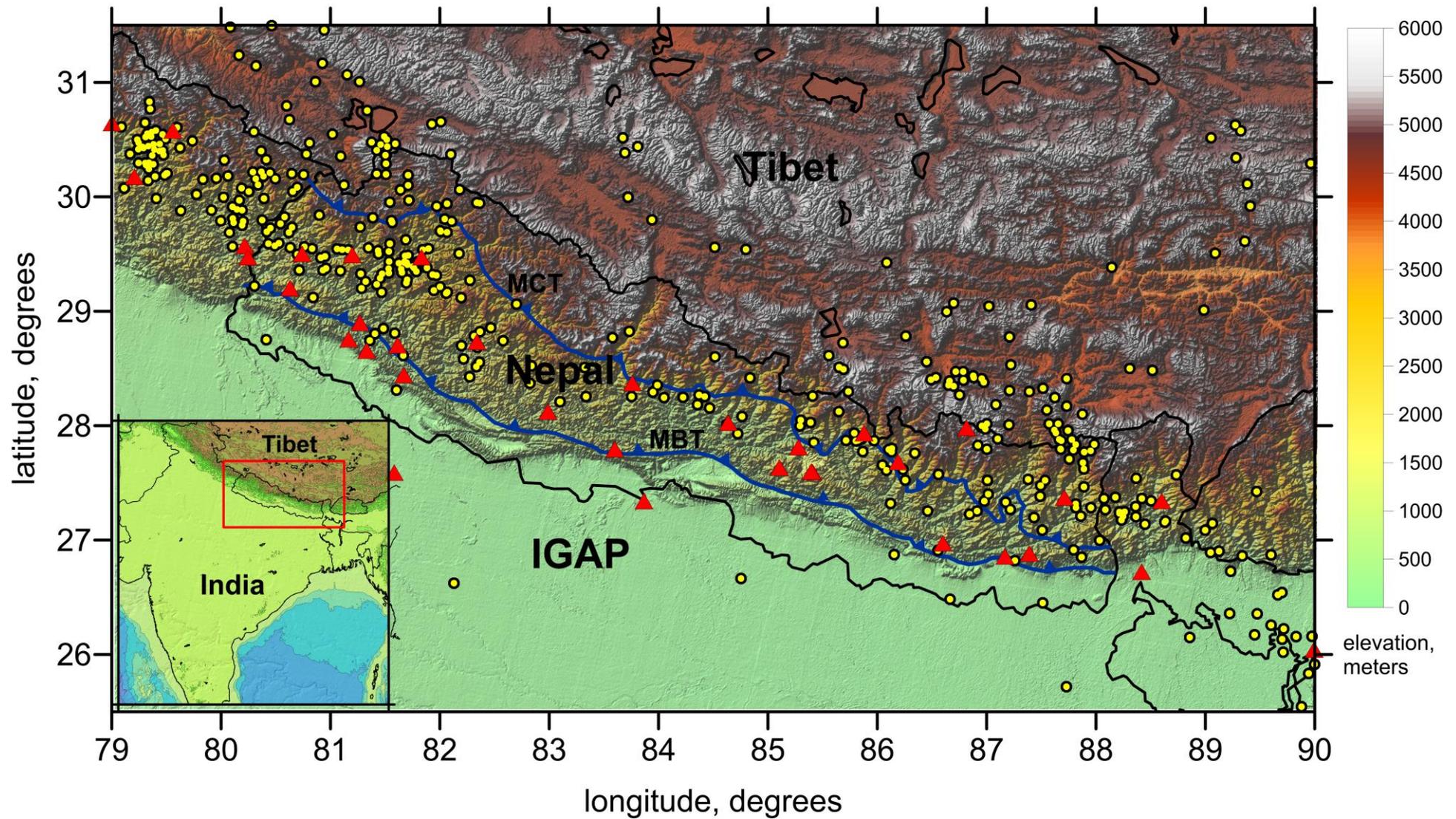


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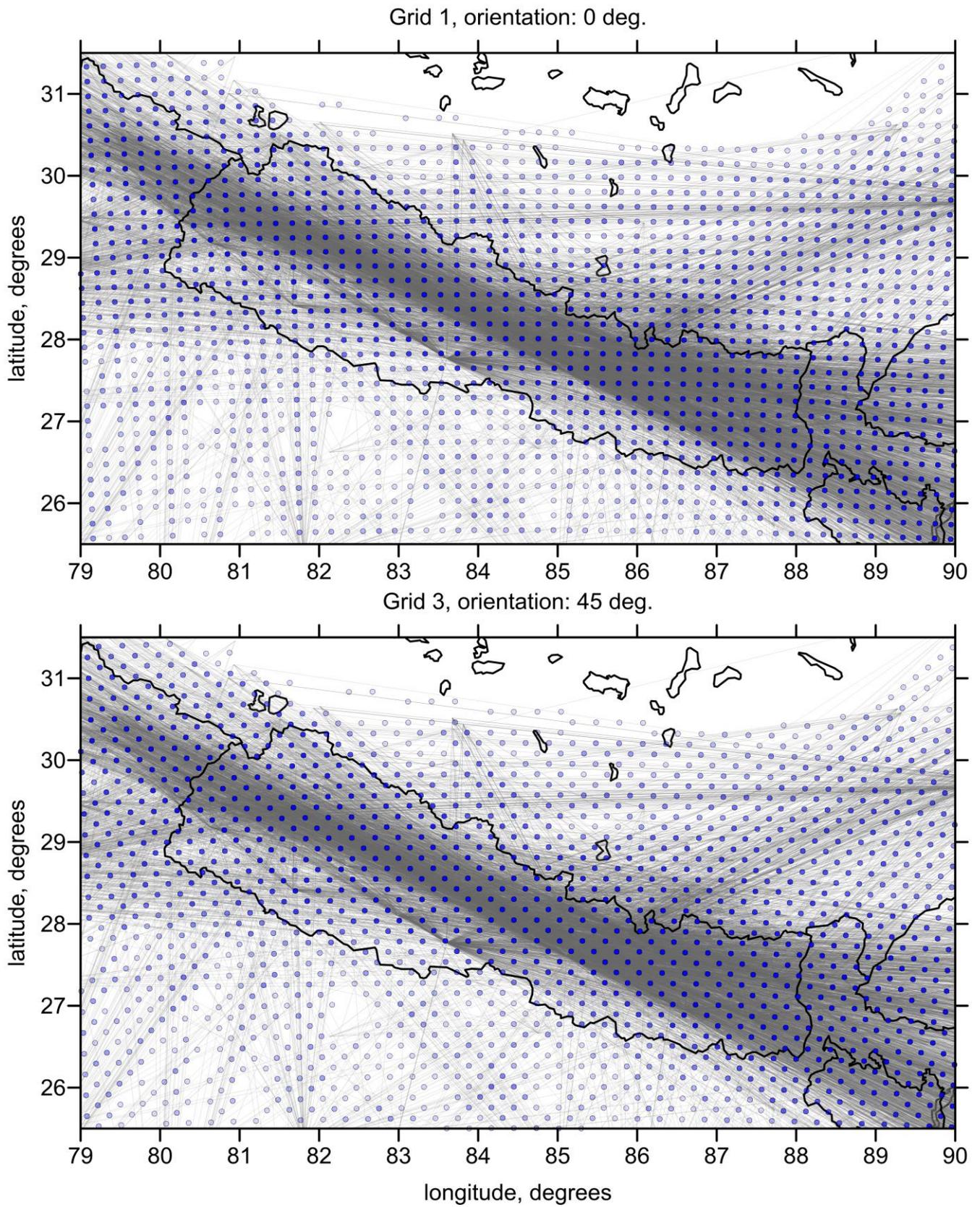


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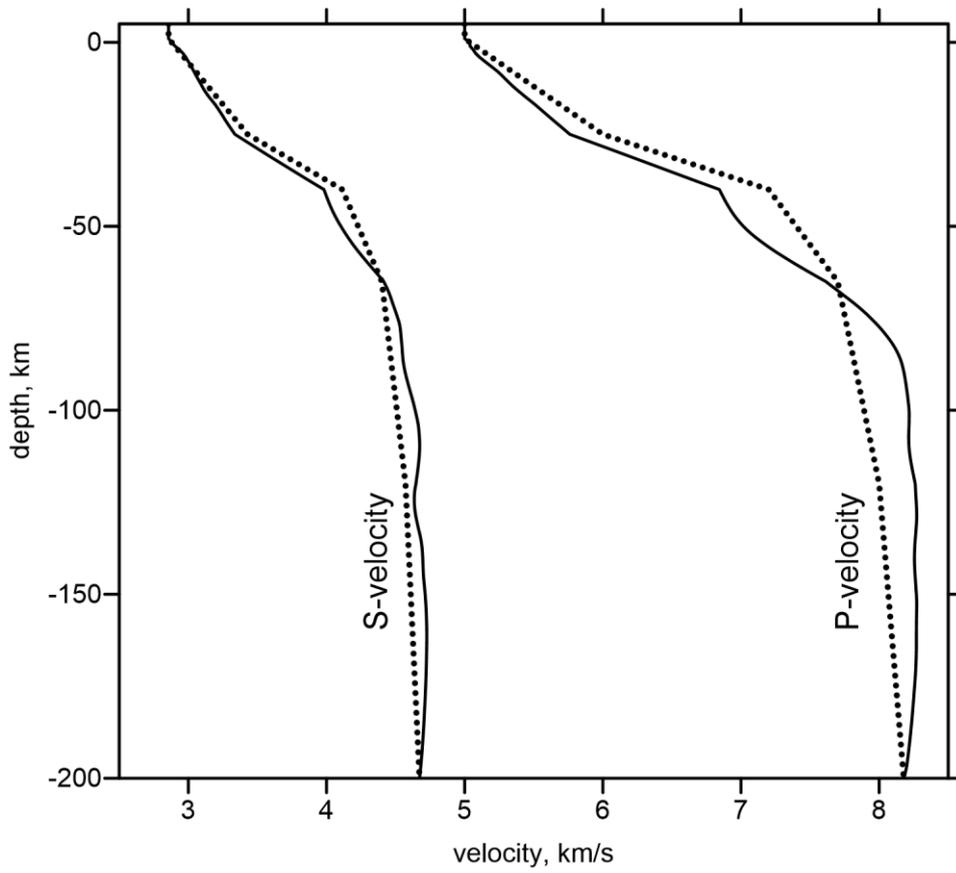


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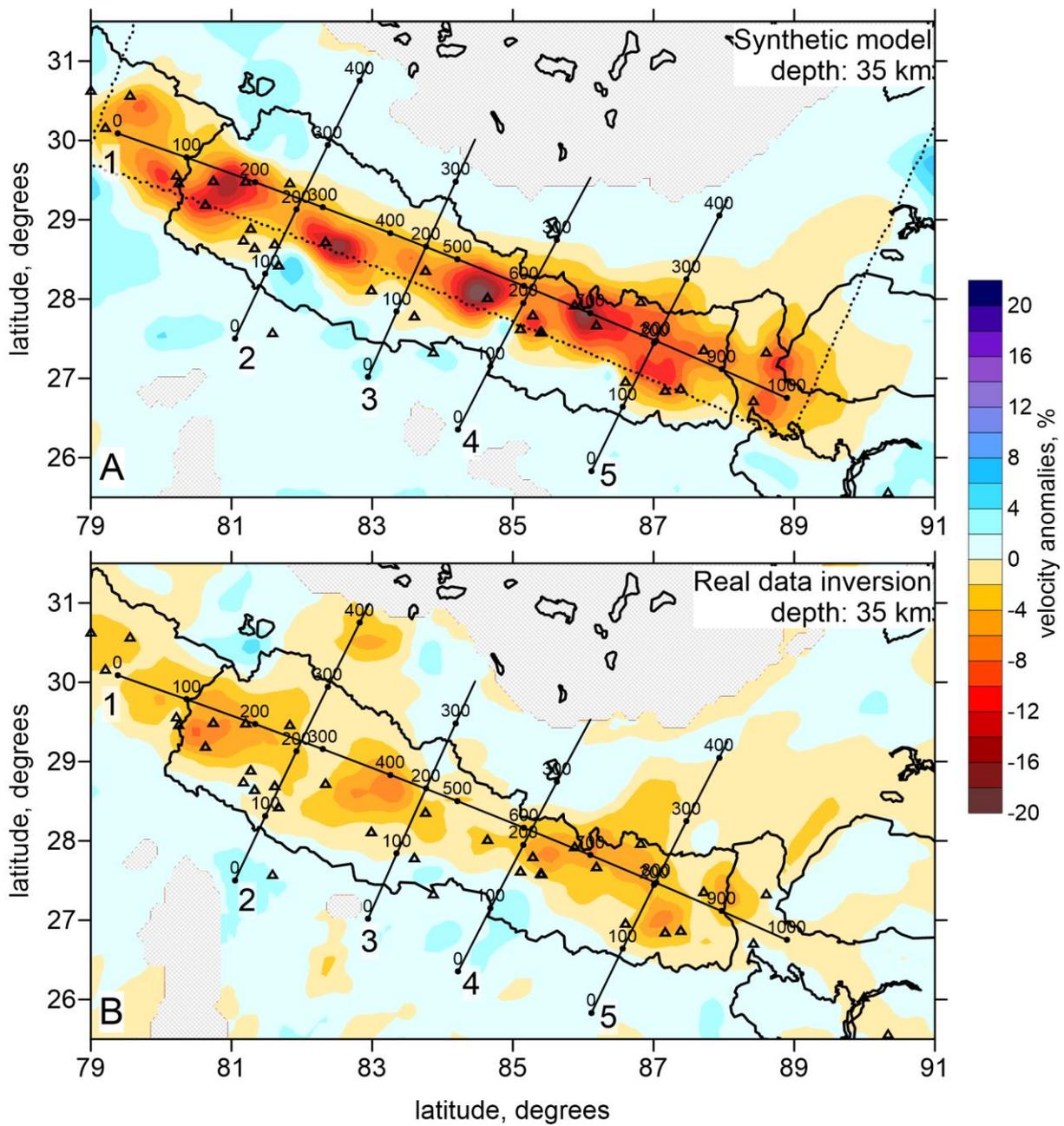


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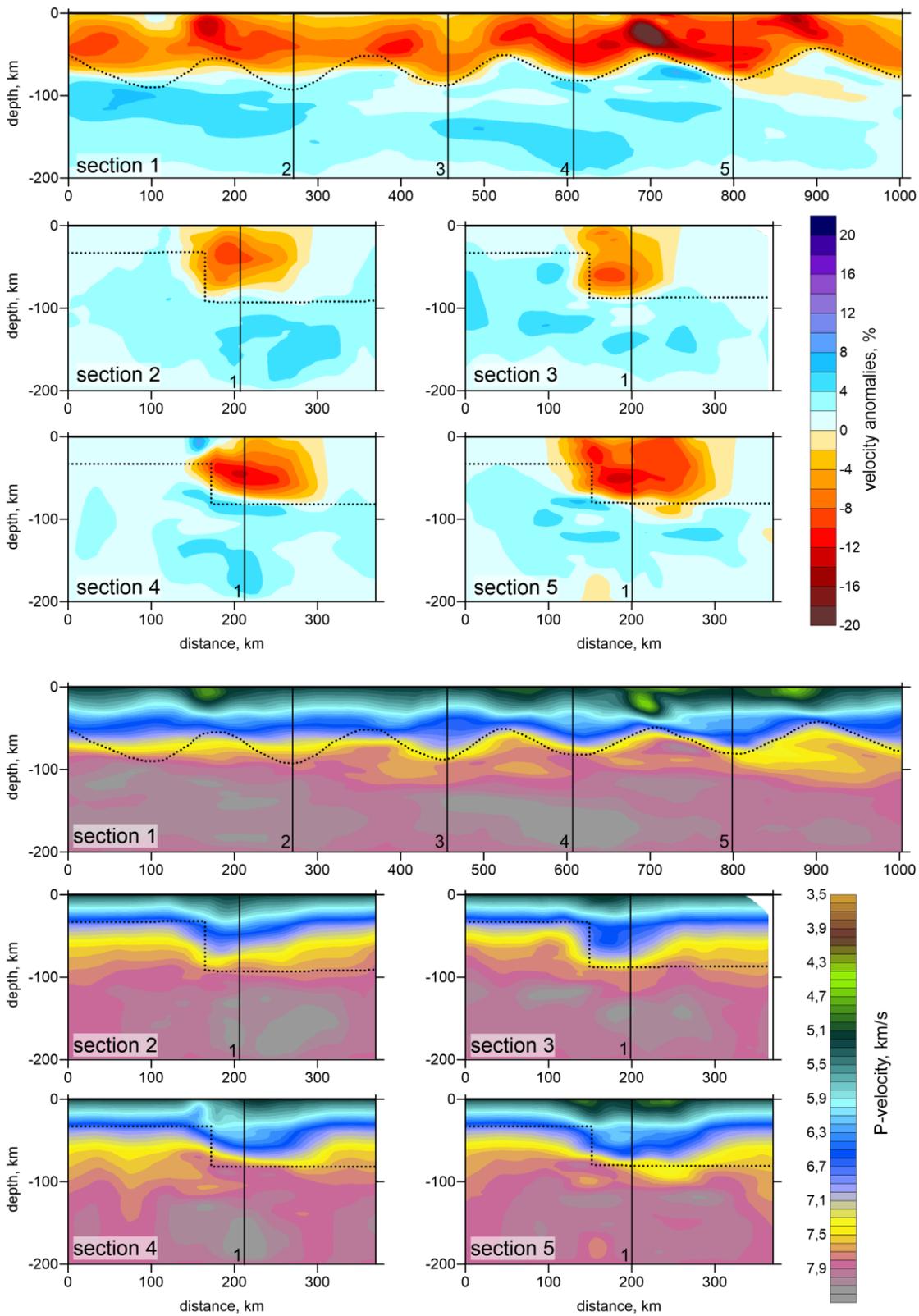


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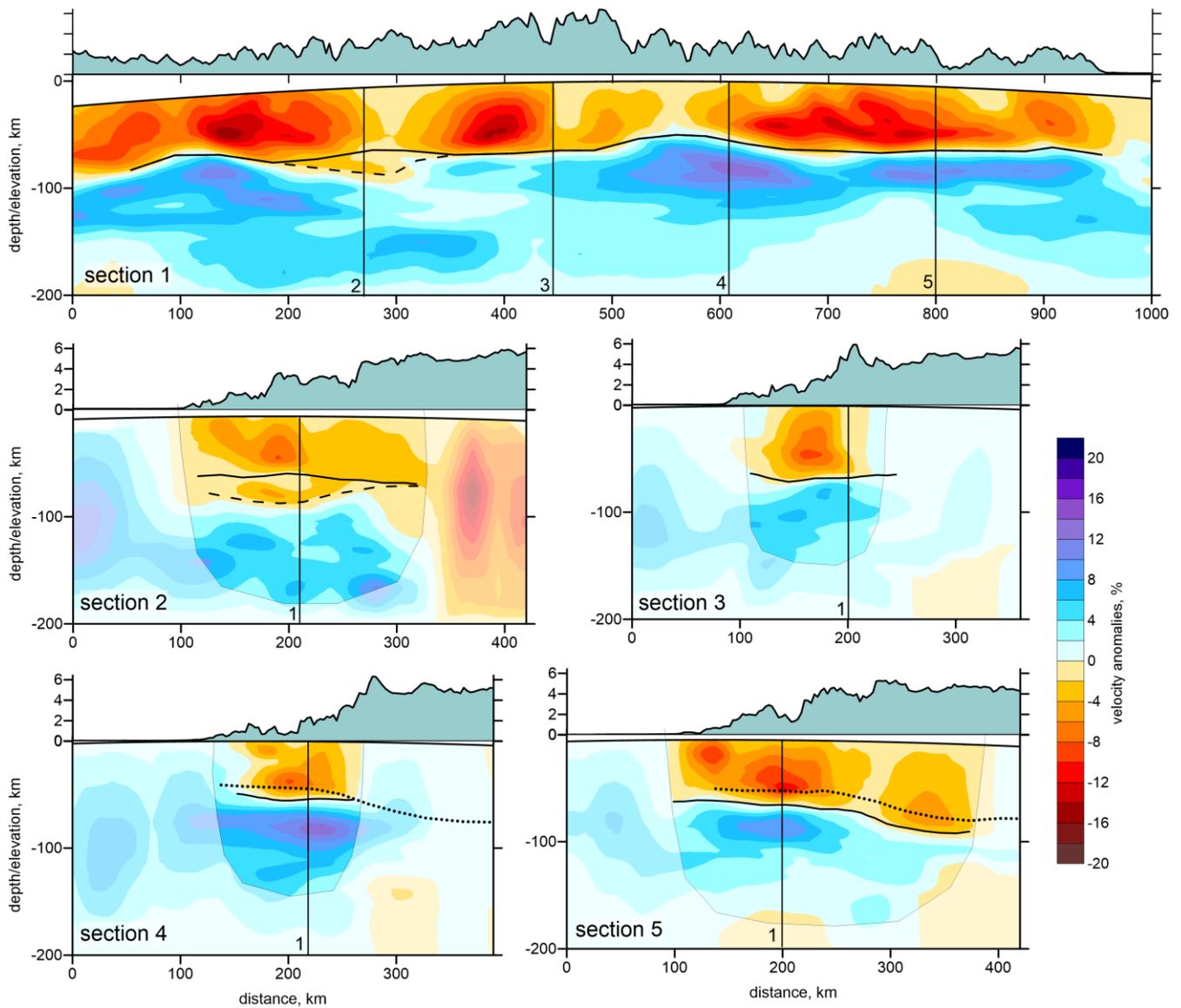


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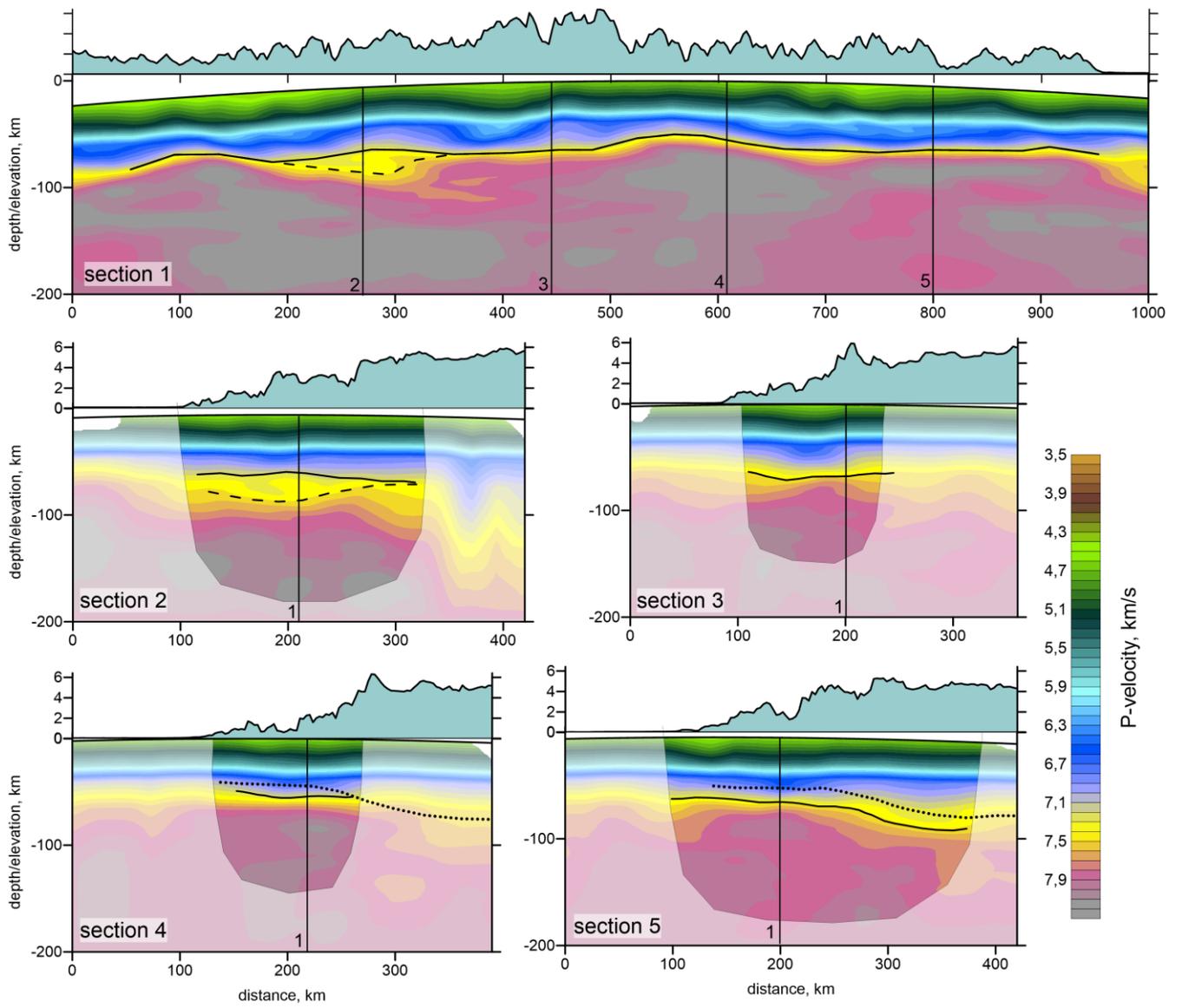


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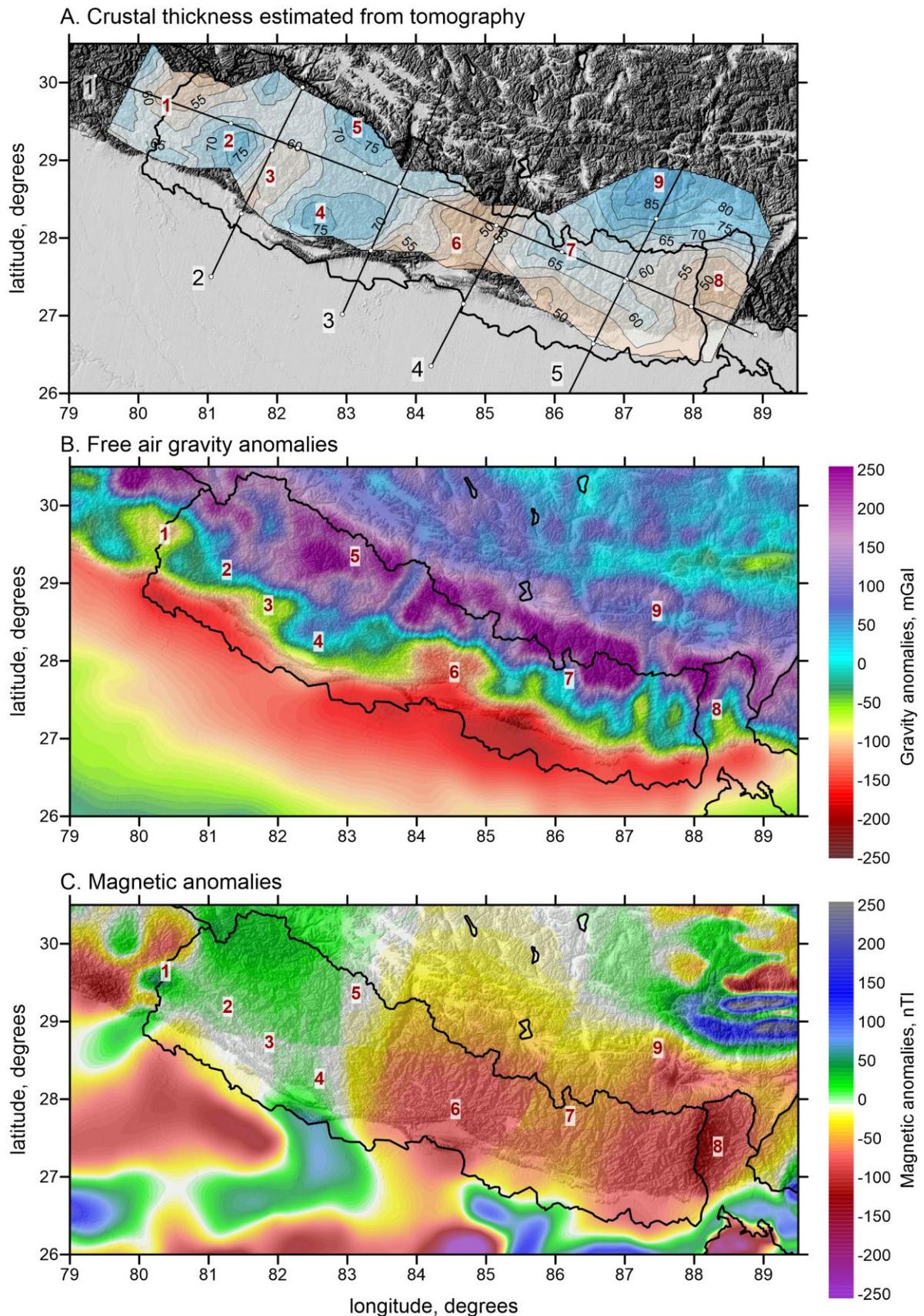


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