

Author's response to the second review of the manuscript entitle "Deciphering the metamorphic evolution of the Pulo do Lobo metasedimentary domain (SW Iberian Variscides)" by Pérez Cáceres *et al.* (Manuscript number se-2019-143):

We acknowledge the positive overview and constructive comments raised by Prof. J.B. Murphy, as well as his formal corrections that contributed to clarify the final manuscript. All the suggestions and corrections have been attended, which will increase the prominence and dissemination of our paper. Below the marked-up revised manuscript version is included.

1 Deciphering the metamorphic evolution of the Pulo do 2 Lobo metasedimentary domain (SW Iberian Variscides)

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20 21 22 **Abstract**

23 The Pulo do Lobo domain is one of the units ~~related to~~exposed within the orogenic suture
24 zone between the Ossa-Morena and the South Portuguese zones in the SW Iberian
25 Variscides. This metasedimentary unit has been classically interpreted as a Rheic subduction-
26 related accretionary prism formed during ~~the~~ pre-Carboniferous convergence and eventual
27 collision between the South Portuguese Zone (part of Avalonia) and the Ossa-Morena Zone
28 (peri-Gondwanan terrane). Discrete mafic intrusions also occur ~~in~~within the dominant Pulo
29 do Lobo metapelites, related to an intraorogenic Mississippian transtensional and magmatic
30 event that had a significant thermal input. Three different approaches have been applied to
31 the Devonian/Carboniferous phyllites and slates of the Pulo do Lobo domain in order to
32 study their poorly known low-grade metamorphic evolution. X-Ray diffraction (XRD) was
33 used to ~~unravel~~identify the mineralogy and measure crystallographic parameters (illite
34 “crystallinity” and K-white mica *b*-cell dimension). Compositional maps of selected samples
35 were obtained from electron probe microanalysis, which allowed processing with

36 XmapTools software, and chlorite semi-empirical and thermodynamic geothermometry was
37 performed. Thermometry based on Raman spectroscopy of carbonaceous material (RSCM)
38 was used to obtain peak temperatures.

39 The microstructural study shows the existence of two phyllosilicate growth events ~~at~~in the
40 chlorite zone, the main one (M₁) related to the development of a Devonian foliation S₁, and
41 a minor one (M₂) associated with a crenulation cleavage (S₂) developed ~~at~~in middle/upper
42 Carboniferous time. M₁ entered well into epizone (greenschist facies) conditions. M₂
43 conditions were at lower temperature, reaching the anchizone/epizone boundary. These data
44 accord well with the angular unconformity that separates the Devonian and Carboniferous
45 formations of the Pulo do Lobo domain. The varied results obtained by the different
46 approaches followed, combined with microstructural analysis, ~~are indicative of~~provide
47 different snapshots of the metamorphic history. Thus, RSCM temperatures are higher in
48 comparison with the other methods applied, which is interpreted ~~as reflecting~~to reflect a
49 faster reequilibration during the short-lived thermal Mississippian event. Regarding the
50 metamorphic pressure, the data are very homogeneous: very low celadonite content (0-10
51 %) in muscovite (and low values of K-white mica *b*-cell dimension (8.995 Å mean value),
52 indicating a low pressure/temperature gradient, which is unexpected in a subduction-related
53 accretionary prism.

54

55 **Keywords**

56 Pulo do Lobo metapelites

57 Low-pressure gradient

58 Illite “crystallinity”

59 Chlorite geothermometry

60 Raman spectroscopy of carbonaceous material

61

62 **Highlights**

63 A multidisciplinary approach has been applied to study the metamorphism of the Pulo do
64 Lobo metapelites.

65 Devonian metamorphism entered epizone conditions.

66 Carboniferous metamorphism reached the anchizone/epizone boundary.

67 The inferred low-pressure gradient is incompatible with a subduction-related accretionary
68 prism.

69 1. Introduction

70 The ~~knowledge~~determination of temperature and pressure conditions reached by the low-
71 grade metasedimentary units stacked hinterlands of orogens ~~helps to better~~
72 ~~interpret~~constrains their tectonometamorphic evolution (e.g., Goffé and Velde, 1984;
73 Franceschelli et al., 1986; Ernst, 1988; Gutiérrez-Alonso and Nieto, 1996; Frey and
74 Robinson, 1999; Bousquet et al., 2008; Lanari et al., 2012). ~~In this regard, the~~The various
75 results derived from the application of diverse geothermometric and/or geobarometric
76 methods may also allow the identification and characterization of superposed
77 tectonometamorphic events, thus improving the knowledge of the P-T paths and their
78 tectonic significance (e.g., Brown, 1993; Crouzet et al., 2007; Ali, 2010; Lanari et al., 2012;
79 Airaghi et al., 2017).

80 The metamorphism of the Iberian Variscides has been mostly studied on intensely high grade
81 metamorphosed rocks in order to characterize ~~high-grade events~~ and obtain the P-T-t paths
82 of suture-related units (e.g., Gil Ibarguchi et al., 1990; Abalos et al., 1991; Escuder Viruete et
83 al., 1994; Barbero, 1995; Arenas et al., 1997; Fonseca et al., 1999; López-Carmona et al., 2013;
84 Martínez Catalán et al., 2014). The low- to very low-grade units have been also studied (e.g.,
85 Martínez Catalán, 1985; Bastida et al., 1986, 2002; López Munguira et al., 1991; Gutiérrez-
86 Alonso and Nieto, 1996; Abad et al., 2001, 2002, 2003a; Martínez Poyatos et al., 2001; Nieto
87 et al., 2005; Vázquez et al., 2007), despite the scarcity of appropriate robust methodologies
88 to apply in these kind of rocks. Obtaining new results from the low-grade rocks of the Pulo
89 do Lobo domain, a suture-related low-grade unit in SW Iberia, is of capital prime importance
90 in order to understand its significance and tectonometamorphic evolution, ~~that whose~~
91 interpretations have been cause of discrepancies, and to reconstruct the overall history of the
92 SW Iberian Variscides.

93 In this work, three different methodologies are applied to a number of samples of the Pulo
94 do Lobo domain (Fig. 1): (i) X-Ray Diffraction (XRD) in order to identify minerals not easily
95 recognizable with optical microscopy (fine-grained muscovite, paragonite, mixed-layer
96 phyllosilicates, etc.) and obtain thermobarometric information via the measurement of
97 crystallographic parameters (illite “crystallinity” and *b*-cell dimension); (ii) Compositional
98 maps derived from electron probe microanalysis (EPMA), which enable the recognition of
99 different tectonometamorphic events by combining mineral composition and microtextural
100 features (e.g., Airaghi et al., 2017), as well as the application of geothermobarometers based
101 on chlorite and K-white mica compositions; and (iii) Raman spectroscopy of carbonaceous
102 material (RSCM) to estimate peak temperatures ~~thanks to using~~ an adapted thermometric
103 calibration. ~~Firstly, the~~The results obtained enables discussing the tectonometamorphic
104 evolution of the Pulo do Lobo domain. Moreover, the comparison of the different
105 approaches ~~allows to know further~~tests their reliability and sensitivity to characterize
106 different geological processes.

107

108 2. Geological setting

109 The SW Iberian Variscides resulted from the Devonian-Carboniferous left-lateral oblique
110 collision of three different terranes: the Central Iberian Zone (CIZ), the Ossa-Morena Zone

111 (OMZ) and the South Portuguese Zone (SPZ) (Fig. 1a). ~~The~~, and the boundaries between
112 these terranes are considered as orogenic sutures (e.g., Quesada, ~~1990~~1991; Pérez-Estaún
113 and Bea, 2004; Pérez-Cáceres et al., 2016). Besides the dominant left-lateral shortening
114 kinematics, SW Iberia also attests Mississippian synorogenic sedimentary basins, widespread
115 mafic magmatism and high-temperature metamorphic areas, which ~~altogether reveal~~
116 ~~suggest~~ an intraorogenic transtensional stage (Simancas et al., 2003, 2006; Pereira et al., 2012;
117 Azor et al., 2019).

118 The OMZ is commonly interpreted as a ~~fragment of~~ continental ~~piece~~crust that drifted from
119 the CIZ (i.e., north Gondwana) in early Paleozoic times (Matte, 2001). The OMZ/CIZ suture
120 (Badajoz-Córdoba Shear Zone) includes early Paleozoic amphibolites with oceanic affinity,
121 eclogite relicts and intense high- to low-grade left-lateral shear ~~imprint~~deformation (Burg et
122 al., 1981; Abalos et al., ~~1991~~; Quesada, 1991; Azor et al., 1994; Ordóñez-Casado, 1998; López
123 Sánchez-Vizcaíno et al., 2003; Pereira et al., 2010). Ediacaran to Carboniferous sedimentary
124 successions with an ~~angular~~ unconformity at the base ~~or of~~ the Lower Carboniferous
125 characterize the OMZ. Low-grade regional metamorphism dominates the OMZ, though
126 there are areas of high-temperature / low-pressure metamorphism associated with Early
127 Carboniferous magmatism (e.g. Bard, 1977; Crespo-Blanc, 1991; Díaz Azpiroz et al., 2006;
128 Pereira et al., 2009).

129 The SPZ is a continental piece considered as a fragment of Avalonia, and thus the
130 OMZ/SPZ boundary is usually interpreted as the Rheic Ocean suture (Crespo-Blanc and
131 Orozco; 1988; Eden and Andrews, 1990; Silva et al., 1990; Quesada et al., 1994; Braid et al.,
132 2011; Pérez-Cáceres et al., 2015, 2017). This boundary is delineated by the Beja-Acebuches
133 Amphibolites (Fig. 1b), a narrow strip of metamafic rocks that resembles a dismembered
134 ophiolitic succession (from greenschists to metagabbros and locally ultramafic rocks) (e.g.,
135 Bard, 1977; Crespo-Blanc, 1991; Quesada et al., 1994). This unit was interpreted as a Rheic
136 ophiolite (Munhá et al., 1986; Crespo-Blanc, 1991; Fonseca and Ribeiro, 1993; Quesada et
137 al., 1994; Castro et al., 1996), ~~though this~~. ~~This~~ idea was reconsidered based on the
138 Mississippian age of the mafic protholits (≈ 340 Ma; Azor et al., 2008). ~~At present, the Beja-~~
139 ~~Acebuches unit is better interpreted as an outstanding~~ and is more likely evidence of the
140 early Carboniferous intraorogenic, lithospheric-scale transtensional and magmatic episode
141 that ~~here~~ obscures the previous suture-related features of the OMZ/SPZ boundary (Pérez-
142 Cáceres et al., 2015 and references therein). Nevertheless, there is also the alternative
143 explanation that the OMZ/SPZ boundary was a protected tract of Rheic oceanic lithosphere
144 that did not close until Carboniferous time (Murphy et al., 2016; Braid et al., 2018; Quesada
145 et al., 2019). The rocks of the Beja-Acebuches Amphibolites were affected by a left-lateral
146 ductile shearing ~~that~~which occurred at granulite to greenschist facies conditions, though
147 amphibolite facies conditions were dominant (e.g., Quesada et al., 1994; Castro et al., 1996;
148 Castro et al., 1999; Díaz Azpiroz et al., 2006). This metamorphism has been dated at 345-
149 330 Ma (Dallmeyer et al., 1993; Castro et al., 1999), thus suggesting that it started very shortly
150 after the magmatic emplacement.

151 North of the Beja-Acebuches Amphibolites, the allochthonous Cubito-Moura unit might be
152 the only ~~witness~~vestige of the Rheic Ocean suture (Fonseca et al., 1999; Araújo et al., 2005;
153 Pérez-Cáceres et al., 2015). This unit was emplaced onto the southern OMZ border (Fig. 1b)
154 with a left-lateral top-to-the-ENE kinematics (Ponce et al., 2012). It contains Ediacaran-

155 Lower Paleozoic ~~metasediments~~metasedimentary rocks and Ordovician MORB-featured
156 mafic rocks (≈ 480 Ma; Pedro et al., 2010) transformed into high-pressure blueschists and
157 eclogites at ≈ 370 Ma (Moita et al., 2005). The high-pressure metamorphism has also been
158 studied by using white mica and chlorite (and chloritoid pseudomorphs) mineral equilibria
159 (Booth-Rea et al., 2006; Ponce et al., 2012; Rubio Pascual et al., 2013), yielding peak
160 conditions of 1 GPa at 450 °C.

161 South of the Beja-Acebuches Amphibolites, low- to very low-grade successions crop out:
162 Devonian siliciclastics, earliest Carboniferous volcano-sedimentary rocks, and a south-
163 migrating Carboniferous flysch (e.g., Oliveira, 1990). These rocks are usually grouped into
164 two geological domains: the Pulo do Lobo domain to the north, and the SPZ s. str. to the
165 south (Fig. 1a, b). The deformation in the SPZ consists ~~in~~of a south- to southwest-vergent
166 fold and thrust belt with decreasing strain intensity and age southwards (Oliveira, 1990;
167 Simancas et al., 2004). The metamorphic grade also decreases southwards, from epizone to
168 ~~diagenesis~~diagenetic, through the SPZ (Munhá, 1990; Abad et al., 2001). The Pulo do Lobo
169 domain, which has been traditionally considered as a suture-related unit (see below) is the
170 focus of this work.

171

172 **2.1. Pulo do Lobo domain**

173 The Pulo do Lobo domain constitutes a polydeformed structure affecting low-grade
174 Devonian-Carboniferous sedimentary formations. These formations are, from bottom to top
175 (Fig. 1b-c):

176 (i) The Pulo do Lobo ~~formation~~Formation (s. str.), which is constituted by a succession of
177 satiny black to grey phyllites and fine-grained schists with minor intercalations of quartz
178 sandstones (Fig. 2a). The presence of abundant segregated quartz veins (pre- to post-folding)
179 is common. The palynological content suggests a middle Frasnian age (Pereira et al., 2018).

180 (ii) The Ribeira de Limas ~~formation~~Formation, which is constituted by phyllites with thin
181 beds of quartz sandstones and arkoses (Fig. 2b). The presence of palynomorphs also suggests
182 a middle Frasnian age for this formation (Pereira et al., 2018). The contact with the
183 underlying Pulo do Lobo ~~formation~~Formation is gradual, with a progressive increase of
184 sandstones and a decrease of phyllites upwards. For that reason, we will refer to the Pulo do
185 Lobo and Ribeira de Limas formations as the lower formations of the Pulo do Lobo domain.
186 Furthermore, these lower formations share the same ~~structure~~structural fabrics consisting
187 ~~in~~of three fold-related foliations (Fig. 2a-b; Pérez-Cáceres et al., 2015). The first foliation of
188 the lower formations (S_1) is preserved inside microlithons of the second foliation (S_2); usually,
189 the angle between these two foliations is high. S_2 is the main foliation and consists in a
190 crenulation-dissolution cleavage that ~~frequently~~commonly appears as a milimetric- to
191 centimetric-spaced tectonic banding. This foliation is axial-~~plane~~planar to north-vergent
192 folds. The third foliation (S_3) is a spaced crenulation-dissolution cleavage that
193 ~~sometimes~~typically develops a characteristic decimetric- to metric-scale tectonic banding. S_3
194 is associated with upright to slightly south-vergent folds.

195 (iii) The Santa Iría ~~formation~~Formation, which is composed by alternating beds of slates and
196 greywackes (Fig. 2c). The greywacke beds show normal grading and an erosive base.

197 Paleontological and palynostratigraphic studies suggest an Upper Famennian age for this
198 formation (Pereira et al., 2008; 2018). However, an early Carboniferous age is ~~much~~also
199 plausible, since more than 90% of the palynomorphs correspond to reworked material
200 (Lopes et al., 2014) and the younger detrital zircon population is early Carboniferous (Braid
201 et al., 2011; Pérez-Cáceres et al., 2017; Pereira et al., 2019). The Santa Iría
202 ~~formation~~Formation only shows two foliations, correlative with the last two deformation
203 phases in the lower formations. Therefore, an unconformity between them is inferred, which
204 ~~also agrees~~is consistent with the age and flysch character of the Santa Iría
205 ~~formation~~Formation (Pérez-Cáceres et al., 2015). S₂ is observed as a penetrative slaty
206 cleavage, while S₃ is a disjunctive crenulation cleavage.

207 According to Silva et al. (1990) and Pérez-Cáceres et al. (2015), the two main foliations (S₂
208 and S₃) in the Pulo do Lobo domain resulted from the middle/upper Carboniferous collision
209 between the OMZ and SPZ. ~~On the contrary, the~~The first foliation (S₁) in the Pulo do Lobo
210 domain might have formed during the ~~vanishing stages~~convergence of ~~the~~ Rheic Ocean
211 subduction and/or the ~~starting~~beginning of the Variscan collision, probably at Late
212 Devonian time.

213 The Pulo do Lobo domain contains some decimetric- to metric-scale lenticular bodies of
214 MORB-featured metamafic rocks. At some outcrops, the mafic rocks are embedded in a
215 greenish detrital matrix, thus suggesting an olistostromic origin (the Peramora Olistostrome;
216 Eden and Andrews, 1990). These rocks are tectonically imbricated with the phyllites of the
217 Pulo do Lobo ~~formation~~Formation and hence forming a tectonic mélangé (the so-called
218 Peramora Mélangé; Fig. 1b-c; Eden, 1991; Dahn et al., 2014). Based on this aspect and on
219 the supposedly Rheic Ocean derived greenschists, the Pulo do Lobo domain has been
220 classically interpreted as a pre-collisional subduction-related accretionary prism (Eden and
221 Andrews, 1990; Silva et al., 1990; Eden, 1991; Braid et al., 2010; Ribeiro et al., 2010; Dahn et
222 al., 2014; Quesada et al., 2019). However, the recently obtained Mississippian U/Pb zircon
223 ages from the metamafic rocks (Dahn et al., 2014; Pérez-Cáceres et al., 2015) make ~~it~~difficult
224 to maintain such ~~an~~ hypothesis. More properly, they can be interpreted as mafic
225 intrusions/extrusions in the frame of the intraorogenic transtensional magmatic event that
226 prevailed in SW Iberia during the Mississippian. The metamafic rocks display a foliation
227 (equivalent to the S₂ of the enveloping metasediments) developed ~~at~~during loosely
228 constrained greenschist facies ~~conditions~~metamorphism. These rocks would have been
229 imbricated with the Pulo do Lobo ~~metasediments~~metasedimentary rocks during the second
230 deformation phase which caused S₂ (Pérez-Cáceres et al., 2015). Our multidisciplinary
231 metamorphic study of the Pulo do Lobo metasediments provides ~~with~~crucial data
232 concerning the tectonic significance of this domain.

233

234 3. Samples and analytical methods

235 Eighteen samples were collected from well-exposed outcrops of phyllosilicate-rich detrital
236 rocks of the Pulo do Lobo domain along two north-south transects perpendicular to the
237 structural trend. Five samples belong to the Santa Iría ~~formation~~Formation (unconformable
238 upper formation) and thirteen to the lower formations (Pulo do Lobo and Ribeira de Limas
239 formations) (location of samples are ~~shown~~ in the map and cross-sections of Fig. 1b-c and

240 the UTM coordinates are given in supplementary information). As a whole, the samples were
241 selected in non-altered fresh outcrops, far from faults and joints, and were ~~taken~~ as
242 homogeneous and representative as possible. Sampling design strategy was ~~intended~~ to collect
243 representative sites, both of the overall stratigraphic succession and along ~~the~~ two transects.
244 We also aimed to characterize the unconformity between the lower and upper formations
245 from a metamorphic point of view, since “crystallinity” aspect at first sight seems to the naked
246 eye appears to be lower in the Santa Iría ~~formation~~ Formation. Some samples from the
247 lowermost Pulo do Lobo ~~formation~~ Formation were collected not far ≈ 200 m from the
248 metabasite lenses of the Peramora Mélange.

249 Samples were examined under the optical microscope and scanning electron microscope
250 (SEM) for overall mineralogy, deformation textures and minerals/foliations relationships
251 using an environmental scanning electron microscope FEI model Quanta 400, operating at
252 15–20 keV (Centro de Instrumentación Científica-CIC, University of Granada, Spain).

253

254 3.1. X-Ray diffraction

255 Sample preparation and analysis by XRD were done in the laboratories of the Department
256 of Mineralogy and Petrology of the University of Granada (Spain). After washing and
257 cleaning of patinas and oxides, samples were crushed to a <2 mm fraction. The <2 μm
258 fractions were separated by repeated extraction of supernatant liquid after centrifugation,
259 according to the Stokes' law. Oriented aggregates were prepared by sedimentation on glass
260 slides of whole-rock and <2 μm fractions (the latter aims to minimize the content of detrital
261 micas non-re-equilibrated during very low-grade metamorphism, which are generally larger
262 than 2 μm; Moore and Reynolds, 1997). Samples were also treated with ethylene glycol
263 (EGC) to identify illite/smectite or chlorite/smectite mixed-layers on the basis of their
264 expansibility-expandibility. Samples were analyzed using a PANalytical X'Pert Pro powder
265 diffractometer equipped with an X'Celerator detector, CuKα radiation, operated at 45 kV
266 and 40mA, Ni filter and 0.25° divergence slit. The resulting diffraction diagrams were
267 examined to extract information on mineralogy based on their characteristic reflections and
268 white mica crystal data.

269 The Illite “Crystallinity” index (Kübler Index; KI; Kübler, 1968) has been estimated from
270 the measurement of the full peak-width of K-white mica at half maximum intensity (FWHM
271 values), expressed as $\Delta^{\circ}2\theta$ of the Bragg angle. Preparation of samples and experimental
272 conditions were carried out according to IGCP 294 IC Working Group recommendations
273 (Kisch, 1991). A step increment of 0.008° 2θ and a counting time of 52 s/step were used in
274 the diffractometer. The KI has been was measured in all samples for both the 5 and 10 Å
275 reflection peaks of K-white mica in order to identify possible effects of other overlapping
276 phases (Nieto and Sánchez-Navas, 1994; Battaglia et al., 2004). Some XRD traces showing
277 complex mixture of mixed-layered minerals were decomposed with the MacDiff software
278 (Petschick, 2004). The FWHM values obtained in the laboratory (x) have been were
279 transformed to Crystallinity Index Standard (CIS) values (y) using the equation $y=0.972x +$
280 0.1096 ($R^2 = 0.942$), obtained from the measure in our lab laboratory of the international
281 standards of Warr and Rice (1994). Finally, they have been these values are expressed in
282 term terms of traditional KI values using the equation of Warr and Ferreiro Mähnlmann

283 (2015; 'CIS' = 1.1523*Kübler index 'Basel lab' + 0.036). The lower and upper boundaries of
284 the anchizone in the KI scale are 0.42 and 0.25 °2θ, respectively (Warr and Ferreiro
285 Mähnlmann, 2015). The thermal range for the anchizone is estimated in c. 200-300 °C,
286 though the KI cannot be considered as a true geothermometer (Frey, 1987; Kisch, 1987).

287 The *b*-cell parameter of white mica was obtained from the (060) reflection peak measured
288 with quartz as internal standard on polished rock-slices cut normal to the sample main
289 foliation S_2 (Sassi and Scolari, 1974). The *b*-cell dimension of K-white mica is ~~often~~ typically
290 proportional to the ~~magnitude~~ extent of phengitic substitution and therefore ~~is~~ considered ~~as~~
291 a proxy of the pressure conditions during its crystallization. Thus, Guidotti and Sassi (1986)
292 have shown that *b* values lower than 9.000 Å are typical of low-pressure facies conditions,
293 while *b* values higher than 9.040 Å are related to ~~rather~~ relatively high-pressure facies
294 metamorphism. Precise measurements of the basal spacing of white mica (d_{001}) have also
295 been made, using quartz from the sample itself as internal standard. d_{001} is related to the
296 paragonitic Na/K substitution (Guidotti et al., 1992), thereby approximately reflecting the
297 temperature of white-mica formation (Guidotti et al., 1994).

298

299 3.2. EPMA-derived X-Ray compositional maps and chlorite thermometry

300 From all of the collected samples, we selected those with the larger phyllosilicate grain-size
301 for electron probe microanalysis (EPMA). Thus, three carbon-coated polished thin-sections
302 were studied. The selected samples (PLB-84, PLB-88 and PLB-93) belong to the lower
303 formations of the Pulo do Lobo domain (Fig. 2d-e). The Santa Iria samples could not be
304 studied due to the tiny grain size of the ~~slaty~~ minerals (commonly less than 3 μm).

305 Compositional maps and accurate spot analyses were performed on a JEOL JXA-8230
306 EPMA at the Institut des Sciences de la Terre (ISTerre) in Grenoble (France), according to
307 the analytical procedure proposed by de Andrade et al. (2006) and Lanari et al. (2014a). The
308 data acquisition was made in wavelength dispersive spectrometry mode (WDS). Ten
309 elements (Si, Ca, Al, K, Mn, Na, P, Ti, Fe and Mg) were analyzed using five WD
310 spectrometers: TAP crystal for Si and Al, PETL for Ti and P, TAPH for Na and Mg, PETH
311 for K and Ca, and LIFH for Mn and Fe. The standardization was made by using certified
312 natural minerals and synthetic oxides: wollastonite (Si, Ca), corundum (Al), orthoclase (K),
313 rhodonite (Mn), albite (Na), apatite (P), rutile (Ti), hematite (Fe), and periclase (Mg). X-Ray
314 maps were obtained by adding successive adjacent profiles. Beam current of 100 nA and
315 beam size spot (focused) were used. The step (pixel) size was 1 μm and ~~the~~ dwell time was
316 200-300 msec per pixel. Spot analyses were obtained along the profiles within the mapping
317 at 15 kV accelerating voltage, 12 nA beam current and 2 μm beam size spot (focused). The
318 on-peak counting time was 30 sec for each element and 30 sec for two background
319 measurements at ~~both sides~~ either side of ~~the~~ each peak. ZAF correction procedure was
320 applied. The internal standards were orthoclase and/or chromium-augite (Jarosewich et al.,
321 1980), which were run (3 points on each standard) after each profile in order to monitor
322 instrumental drift and estimate analytical accuracy. Drift correction was made, if necessary,
323 using the corresponding regression equation.

324 The WDS X-Ray maps were then processed with XMapTools
325 (<http://www.xmaptools.com>), a MATLAB©-based graphical user interface program to
326 process the chemical maps, link them to thermobarometric models and estimate the
327 pressure-temperature conditions of crystallization of minerals in metamorphic rocks (Lanari
328 et al., 2014a). The compositional maps were standardized with the spot analyses measured
329 along the profiles and mineral compositions were plotted into binary and ternary diagrams
330 using the interface modules *Chem2D* and *Triplot3D*. Chemical maps of amount of tetrahedral
331 aluminum (Al^{IV}) of chlorites were acquired, because ~~is at the base of these amounts are used~~
332 in many empirical chlorite thermometers (e.g. Cathelineau and Nieva, 1985; Cathelineau,
333 1988). The temperature conditions were estimated for each chlorite pixel of the maps using
334 the chlorite thermometer of Lanari et al. (2014b), as well as the approaches of Vidal et al.
335 (2006) and Bourdelle et al. (2013), which are summarized in the supplementary information.

336 In addition to the above mentioned compositional maps, white micas from seven carbon-
337 coated thin sections of the lower formations of the Pulo do Lobo domain were analyzed
338 before with a Jeol WDS four-spectrometer microprobe (JXA-8200 Superprobe) at the
339 University of Huelva (Spain). A combination of silicates and oxides were used for calibration:
340 standards used were wollastonite (Si and Ca), potassium feldspar (Al, K and Na), forsterite
341 (Mg) and fayalite (Fe). Single point analyses were obtained with 20 nA probe current, 1-5 μ m
342 spot size and 15 kV of acceleration voltage, with 5 s counting times.

343

344 3.3. Raman Spectroscopy of carbonaceous material

345 Raman Spectroscopy of Carbonaceous Material (RSCM), is based on the observation that
346 sedimentary carbonaceous material (CM) is progressively transformed into graphite at
347 increasing temperature. Because of the irreversible character of graphitization, CM structure
348 is not sensitive to the retrograde path during exhumation of rocks, but ~~only~~ only depends
349 on the maximum temperature reached during metamorphism (Beysac et al., 2002a).
350 Temperature ~~can be was~~ determined in the range 330-650°C with a calibration-attached
351 accuracy of ± 50 °C due to uncertainties ~~on in the~~ petrologic data used for the calibration:
352 (Beysac et al., 2002a). Relative uncertainties ~~on in~~ temperature ~~are, however, much~~
353 smaller estimates were later reduced (around 10-15 °C; Beysac et al., 2004). For temperature
354 below 330 °C, Lahfid et al. (2010) performed a systematic study of the evolution of the
355 Raman spectrum of CM in low-grade metamorphic rocks in the Glarus Alps (Switzerland).
356 They showed that the Raman spectrum of CM is slightly different from the spectrum
357 observed at higher temperature and they established a quantitative correlation between the
358 degree of ordering/structuration of CM and temperature.

359 In this work, twelve representative thin-sections previously examined by optical microscopy
360 were selected. From ~~them these~~, ten samples were ~~finally~~ analyzed (according to their larger
361 CM grain-size and content): eight samples belong to the lower ~~formations~~ (Pulo do Lobo
362 and Ribeira de Limas formations), ~~while, and~~ the other two samples belong to the Santa Iría
363 formation. Polished thin-sections cut perpendicularly to the S₂ foliation were
364 analyzed at the Institut de Minéralogie, de Physique des Matériaux et de Cosmochimie at the
365 Sorbonne University of Paris (France). We followed closely the analytical procedure
366 described by Beysac et al. (2002a, b; 2003; see supplementary information). More than 15

367 Raman spectra (Fig. 3) were obtained for each sample using a Renishaw InVIA Reflex
368 microspectrometer equipped with a 514.5 nm Modulaser argon laser under circular
369 polarization. The laser was focused by a DMLM Leica microscope, and laser power was set
370 below 1 mW at the sample surface. The Rayleigh diffusion was eliminated by edge filters and
371 the signal was dispersed using a 1800 g/mm grating and finally analyzed by a Peltier cooled
372 RENCAM CCD detector. The recorded spectral window was large to ~~correctly set~~precisely
373 determine the background correction, from 700 to 2000 cm^{-1} in case of low-temperature
374 samples. Before each session, the spectrometer was calibrated with a silicon standard. CM
375 was systematically analyzed behind a transparent adjacent mineral, generally quartz or white
376 mica grains oriented along S_1 . For a full description of the temperature calculations see the
377 supplementary information.

378

379 4. Results

380 According to the petrographic study, all the samples correspond to slates or phyllites with
381 phyllosilicates smaller than 500 μm , composed of variable quartz + K-white mica \pm chlorite
382 \pm feldspar \pm ore and accessory minerals (Fig. 2d-f). Samples from the Santa Iría
383 ~~formation~~Formation have much smaller grain-size and apparently lower “crystallinity” (Fig.
384 2f). The first foliation S_1 is defined by the largest micas and chlorites (Fig. 2d-e), being folded
385 by microscopic- to centimetric-scale tight folds of the second deformation phase (Fig. 2a-b,
386 d-e). The second foliation S_2 is the main foliation at outcrop (Fig. 2a-c), but the ~~development~~
387 ~~of~~ phyllosilicates (mostly white mica) ~~is lesser~~are smaller than ~~those in~~ S_1 . The third
388 foliation S_3 is ~~much less penetrative (Fig. 2a-c) and a fracture cleavage that~~ does not develop
389 phyllosilicates: ~~(Fig 2a-c).~~

390

391 4.1. X-Ray diffraction

392 The mineralogy and crystal parameters of K-white mica obtained from the 18 samples of the
393 Pulo do Lobo domain are summarized in Table 1. The results of KI values, b-cell parameter
394 and d_{001} analyzed in K-white mica, obtained from whole-rock and $<2 \mu\text{m}$ fractions are very
395 similar, which suggests that detrital micas re-equilibrated during metamorphism.

396 The mineralogy of the samples is relatively simple: Qz + Ms + Fsp+ Chl \pm Pg \pm C/S. The
397 slates of the Santa Iría ~~formation~~Formation have quartz, muscovite and chlorite, with
398 chlorite/smectite interlayers (C/S) in some samples. In the lower formations, besides quartz
399 and muscovite, chlorite is present in almost all of the samples, paragonite appears in most of
400 them, and chlorite/smectite interlayers are occasional.

401 KI values measured in the 10 Å peak of white mica from the $<2 \mu\text{m}$ fraction are shown in
402 Table 1 and Fig. 1c with a relative colour bar from orange (lower values) to green (higher
403 values). Values of the Santa Iría samples (n=5) range from 0.20 to 0.26 $\Delta^\circ 2\theta$, the mean value
404 ~~being~~is 0.23 (standard deviation 0.02). ~~As for~~In the lower formations (n=12), KI values range
405 from 0.17 to 0.22, the mean value ~~being~~is 0.19 (standard deviation 0.02). KI values measured
406 in the 5 Å peak (not shown to avoid repetitions) are very similar to those of the 10 Å peak.

407 The measured *b*-cell parameter of white mica varies ~~in~~within a close range around 9 Å (8.991-
408 9.002). Mean value is 8.995 Å (standard deviation 0.003) for the Santa Iría
409 ~~formation~~Formation samples, and 8.997 Å (standard deviation 0.003) for the samples of the
410 lower formations. d_{001} values average 9.992 Å (standard deviation 0.004) and differ slightly
411 between upper and lower formations, ~~being~~with higher ~~values~~ in the upper formation.

412 The results obtained through X-Ray diffraction ~~denote~~reflect low-grade metamorphic
413 conditions due to the KI values between 0.17-0.26 $\Delta^{\circ}2\theta$. In addition, *b*-cell parameters are
414 lower than 9.000 Å which show a low-pressure metamorphic gradient (low
415 pressure/temperature metamorphic facies conditions; Guidotti and Sassi, 1986).

416

417 4.2. Compositional maps and chlorite thermometry

418 X-Ray maps show the distribution of major elements and allow ~~identifying~~identification of
419 white mica, chlorite, and some albite porphyroblasts, with ilmenite and rutile as accessory
420 minerals (Fig. 4a-b). Although quartz is abundant in all of the samples, the ~~zoomed selected~~
421 ~~areas~~areas selected for ~~detailed~~ X-ray mapping (composed mostly by phyllosilicates) do not
422 contain quartz (Fig. 4a-b). White mica is abundant along both S_1 and S_2 foliations (Fig. 2d-e
423 and 4b). Chlorite is found mostly along S_1 , ~~being~~but is very scarce and small-sized
424 ~~along~~within S_2 (Fig. 2e and 4b), with the exception of sample PLB-93 where chlorite is similar
425 in amount in both foliation domains (Fig. 4b).

426 Mapped compositions of end-members of white mica and chlorite have been plotted in the
427 ternary diagrams of Figure 5. The composition of white mica is similar in the three maps-~~It~~
428 ~~and~~ is close to muscovite, with 25% of pyrophyllite and very ~~searc~~eminor celadonite content
429 (0-10%; Fig. 5a). The high content of pyrophyllite (high amount of interlayer vacancies) is
430 typical of low-pressure illitic-mica compositions. Figure 6 shows white mica compositional
431 ratios, which can be related to P/T conditions: they present low degree of Na substitution
432 and low phengitic component, ~~thus being~~and so are close to the muscovite end-member ~~in~~
433 ~~composition~~. These results point to low-pressure conditions and agree well with XRD
434 results: low *b*-cell parameter (< 9.000 Å) and high d_{001} (> 9.985 Å; Table 1).

435 Chlorite compositions are variable, though all of them have in common $\approx 50\%$ clinocllore
436 + daphnite and $\approx 50\%$ amesite + sudoite (Fig. 5b). Chlorites in sample PLB-88 are ~~poor~~low
437 in amesite with a large variation of clinocllore + daphnite and sudoite ~~contents~~. In sample
438 PLB-84, chlorites, ~~variable compositions vary in composition~~ between amesite and sudoite
439 ~~indicate~~indicating a variation of Al^{IV} , which implies an increase of temperature from rims to
440 cores as shown in the chemical maps of Fig. 4c. Finally, PLB-93 chlorites are poor in sudoite
441 and higher in Al^{IV} content, thus suggesting higher average temperatures (Vidal et al., 2006-)
442 ~~than in PLB-84 and PLB-88~~. Altogether, chlorite compositional data ~~suggests~~show the
443 presence of two end-members: sudoite-rich low-temperature (PLB-88), and amesite-rich
444 high-temperature (PLB-93).

445 Maps of Al^{IV} in chlorites ~~have been represented~~are displayed in Fig. 4c. Sample PLB-88
446 ~~shows~~displays lower Al^{IV} content ($\approx 1.1-1.3$ apfu) than sample PLB-93 ($\approx 1.3-1.5$ apfu). In
447 sample PLB-84, some large chlorite grains oriented along S_1 are zoned, with higher Al^{IV}
448 content in the cores (≈ 1.4 apfu) than in the rims (≈ 1.0 apfu; see white square in Fig. 4c).

449 According to the empirical calibration of Cathelineau (1988), Al^{IV} in chlorites increases with
450 temperature. Thus, the Al^{IV} content in chlorites manifests different temperatures in different
451 samples, and also varies from core to rim in singular grains.

452 Temperature maps have been obtained with the semi-empirical thermometer of Lanari et al.
453 (2014b), assuming which assumes that Fe²⁺ is the Fe total (Fig. 4d). Temperatures range
454 between 100-200 °C in sample PLB-88, 150-350 °C in sample PLB-84, and 200-450 °C in
455 sample PLB-93. Tiny chlorites developed along S₂ show yield lower temperatures than larger
456 and more abundant chlorites along S₁, with the exception of sample PLB-93. Furthermore,
457 some large chlorites oriented along S₁ are zoned, showing high-temperature relic cores (350-
458 450 °C; see white insets in Fig. 4c-d) surrounded by low-temperature rims (150-250 °C).

459 To test Vidal et al. (2005, 2006) and Bourdelle et al. (2013) approaches thermometers, an area
460 of representative chlorites in an S₁ microlithon was selected from each map (see red insets in
461 Fig. 4d). Corresponding chlorite compositions were extracted and introduced in the chlorite-
462 quartz-water equilibria (Fig. 7a, Vidal et al., 2005, 2006; Fig. 7b, Bourdelle et al., 2013). The
463 temperature estimates (Fig. 7) derived from both methods are fairly similar ~~with both~~
464 methods, averaging 120-230 °C in sample PLB-88 and 150-380 °C in sample PLB-84. ~~This~~
465 is These estimates are also in agreement with the estimates derived from temperature maps
466 calculated with the Lanari et al. (2014a) model. Only ~~the~~ sample PLB-93 shows a divergence
467 on temperature averages: mostly 150-250 °C with the thermometer of Bourdelle et al. (2013),
468 and 250-350 °C with ~~the one that~~ of Vidal et al. (2005, 2006). In both cases (Bourdelle et al.
469 (2013) and Vidal et al. (2005, 2006) approaches methods), the higher temperature analyses are
470 obtained from crystal cores and belong mostly to their sample PLB-93.

471

472 4.3. RSCM thermometry

473 The ratio parameters and corresponding maximum temperatures obtained from all the
474 spectra analyzed are shown in the supplementary information. The Raman spectra were
475 decomposed into bands following the appropriate fitting procedure described in Beyssac et
476 al. (2002a) for the lower formations (high-temperature Raman spectra; ratio parameter R2)
477 and Lahfid et al. (2010) for the Santa Iria ~~formation~~ Formation (low-temperature Raman
478 spectra; ratio parameters RA1 and RA2; Fig. 3). The average temperatures are shown in Table
479 1 and Fig. 1c with a relative colour bar from red (higher temperature) to blue (lower
480 temperature). The average temperatures for the lower formations range from 420 to 530 °C,
481 with a mean value of 468 °C (standard deviation of 35). The highest temperatures are found
482 in samples PLB-82 (530 °C) and PLB-93 (495 °C), while and estimates of the remaining
483 ones samples do not exceed 480 °C. As for the Santa Iria ~~formation~~ Formation, temperatures
484 are lower (315-330 °C; Table 1) than in the underlying formations.

485

486 5. Interpretation and discussion

487 5.1. Deformation/metamorphism relationships

488 The obtained analytical results must be are interpreted below in the context of the Variscan
489 evolution of the Pulo do Lobo domain. As described above, two regional deformational

490 events D₁ and D₂ ~~gave way to the development of~~ yielded two foliations (Devonian S₁ and
491 Carboniferous S₂) accompanied by metamorphic phyllosilicate growth (M₁ and M₂). In the
492 cross-sections of Fig. 1c, KI values derived from XRD and average temperature from RSCM
493 suggest that the lowest metamorphic grade (green and blue colours) corresponds to the Santa
494 Iría ~~formation~~ Formation.

495 ~~The textural~~ Textural observations ~~evidencesuggest~~ that in most samples of the lower
496 formations M₁ was the main crystallization event, developing abundant and large-sized white
497 mica and chlorite in S₁ microlithons, while M₂ ~~gave way to~~ generated small-sized white mica
498 (e.g., Fig. 2e and map 1 in Fig. 4). On the other hand, polydeformed rocks commonly show
499 previously grown minerals rotated towards a new foliation developed at lower-grade
500 conditions, without new crystallization. This can be the case of the white micas that define
501 S₂ in some samples (illustrated in Fig. 2d), which, in turn, is ~~not contradictory~~ compatible
502 with the similar chemical composition of S₁ and S₂ micas (Fig. 5a). As shown in these samples, S₁
503 is variably crenulated by D₂, ~~so that~~ implying M₁ minerals are variably rotated towards S₂.
504 Consequently, the metamorphic data obtained from the samples of the lower formations will
505 be ascribed to D₁-M₁. Sample PLB-93 might represent an exception, since its slightly higher
506 RSCM and chlorite-derived temperatures ~~might~~ could be due to nearby intrusions (Fig. 1b
507 and 1c.4). ~~At 2; see below~~. In this respect, it is important to note the Mississippian
508 transtensional event (basins development and abundant mafic magmatism) that took place
509 between D₁ and D₂ (Pérez-Cáceres et al., 2015). The characterization of M₂ can be done by
510 studying the samples from the Santa Iría ~~formation~~ Formation, which are only affected by S₂
511 accompanied by small-sized phyllosilicate growth (Fig. 2f). No crystallization has been
512 observed related to the S₃ disjunctive crenulation cleavage.

513

514 5.2. First tectonothermal event (Devonian M₁)

515 The observed mineral ~~association~~ assemblage (Qz + Ab + Ms + Chl ± Pg), together with the
516 presence of C/S is compatible with low-grade metamorphic conditions (Table 1). White mica
517 “crystallinity” values (0.17-0.22 Δ°2θ; average 0.19) are ~~always~~ consistently in the range of the
518 epizone (low-grade or greenschists facies; >300 °C; Frey, 1987; Kisch, 1987, Warr and
519 Ferreira Mähnlmann, 2015), in accordance with the values reported by Abad et al. (2001) in
520 a more general study of the diagenetic-metamorphic evolution of the South Portuguese Zone
521 metapelites. Nevertheless, both the values of KI, ~~still~~ far from 0.14 Δ°2θ (high epizone
522 conditions according to Abad et al., 2006), and their variability, suggest that temperature was
523 not high enough as to stabilize a highly crystalline white mica. This ~~interpretation~~ is in
524 agreement with the low Na content of K-micas coexisting with paragonite (Fig. 6),
525 ~~meaning~~ implying a very-low temperature position in the muscovite-paragonite solvus for
526 natural quasi-binary Pg-Ms pairs (Guidotti et al., 1994). By contrast, the maximum
527 temperatures obtained ~~with~~ from RSCM geothermometry are surprisingly high (420-530 °C;
528 average 470 °C; corresponding to very high epizone or even medium-grade ~~metamorphic~~
529 conditions; Table 1).

530 The composition of paired chlorite and white mica is normally used to calculate pressure and
531 temperature (e.g., Vidal et al., 2006), but ~~the~~ multi-equilibrium approach was not successful
532 because the P-T equilibrium conditions did not converge. This result is indicative of chemical

533 disequilibrium, precluding their use as a reliable geothermobarometer in this case. The
534 temperatures calculated from chlorite compositions following various approaches (Vidal et
535 al., 2006, Fig. 7a; Bourdelle et al., 2013, Fig. 7b; Lanari et al., 2014a, Fig. 4d) are as follow:
536 100-230 °C for sample PLB-88, 150-375 °C for sample PLB-84, and 150-450 °C for sample
537 PLB-93 (Figs. d and 7, and Table 1). The slightly higher temperature of sample PLB-93 is
538 inferred from its highest white mica “crystallinity” ($0.17 \Delta^{\circ}2\theta$), high RSCM temperature (495
539 °C), high-temperature (amesite-rich) chlorite and higher chlorite thermometry (Table 1), and
540 can be explained by its ~~nearness~~proximity to metric-scale mafic igneous bodies of the
541 Peramora Mélange (located at ≈ 200 m to the south; Pérez-Cáceres et al., 2015) and/or to a
542 granite stock (located at ≈ 5 km to the west) (Fig. 1b).

543 In our samples there is some evidence of chlorite retrogression: (i) the chemical
544 disequilibrium showed by the white mica/chlorite geothermobarometer, (ii) the presence of
545 C/S mixed layers ~~that are~~ not stable in the epizone (e.g. Potel et al., 2006), (iii) the difference
546 between temperature estimates from crystal rims to cores, and the higher temperature relic
547 cores preserved in large chlorites defining S_1 (Fig. 4c-d), and (iv) the previously reported
548 XRD and TEM data of chlorite retrograded to smectite and corrensite in the Pulo do Lobo
549 domain (see fig. 1 in Nieto et al., 2005). The existence of chlorites with different
550 compositions crystallized at different temperatures is the ~~usual~~typical scenario (e.g., Vidal et
551 al., 2006, 2016; Lanari et al., 2012; 2014a and b; Grosch et al., 2012; 2014; Cantarero et al.,
552 2014). In such ~~situation, the definition~~situations, precise estimates of a single temperature
553 and pressure attributable to peak conditions ~~is~~are really difficult to obtain. The maximum
554 temperature showed by chlorite relic cores is 350-450 °C (Fig. 4d), which is ~~more in~~
555 ~~accordance~~consistent with the conditions estimated for M_1 by means of RSCM data.

556 An issue that deserves some discussion is the difference in temperature estimates between
557 RSCM and other techniques. RSCM thermometry records the peak temperature and is not
558 sensitive to the retrograde path. Alternatively, other methods based on phyllosilicate
559 compositions are prone to record reequilibration during the retrograde path; thus, they rarely
560 record the peak conditions, except perhaps in the core of certain large crystals. Therefore,
561 RSCM and phyllosilicate-based methods do not record the same information on
562 temperature, ~~being~~but are in fact complementary. The analyzed CM grains were carefully
563 checked by microtextural ~~observation~~observations and spectral geometry to ~~make sure~~ensure
564 that these grains ~~are~~were actually derived from in situ organic matter graphitized during
565 metamorphism.

566 In our case study, at the high peak temperature given by the RSCM thermometry, minerals
567 such as biotite or garnet are expected to crystallize in ~~metasediments~~metasedimentary rocks,
568 though they have not been observed in our samples. Biotite has been said to exist in a few
569 previous works (Apalategui et al., 1983; Braid et al., 2010; Rubio Pascual et al., 2013).
570 However, in a few of our samples, biotite-looking crystals have resulted to be oxichlorites
571 under SEM analyses. The absence or exceptional presence of biotite can be due to whole-
572 rock composition, and explained by growth inhibition related to Na-excess, as evidenced by
573 the presence of albite and paragonite in our samples. Another possible explanation could be
574 the higher sensitivity of CM graphitization to ~~fast~~rapid reequilibration during a short-~~time~~
575 duration thermal event. Thus, the Mississippian intrusions subsequent to M_1 in the Pulo do
576 Lobo ~~formation~~Formation (see description in section 2) could have exerted a fast and locally

577 intense thermal imprint that influenced CM but not the crystal chemistry of silicates.
578 Moreover, recrystallization processes are not only a function of temperature, but are also
579 promoted by deformation/stress, time, fluid/rock ratio (Merriman and Frey, 1999).
580 Observations of this kind (differing reaction kinetics between organic and inorganic material
581 (e.g. illite) in a contact metamorphic setting can be found in Olsson (1999) and Abad et al.
582 (2014). Regarding the timetiming of geological processes, Mori et al. (2017) investigated the
583 importance of heating duration for RSCM thermometry by studying graphitization around
584 dykes. They showed that small-scale intrusions generating short thermal events modify the
585 structure of CM in the surrounding rocks, to conclude and concluded that CM crystallinity is
586 clearly related to contact metamorphism. The influence of low-pressure contact aureoles on
587 RSCM temperature patterns is further supported by the results obtained by Hilchie and
588 Jamieson (2014), who concluded that the variation of RSCM temperatures can be controlled
589 by the subsurface geometry of a pluton. Finally, the long-distance thermal influence of
590 plutonic intrusions on low-grade rocks located as far as 10 km has already been
591 evidenced documented (e.g., Merriman and Frey, 1999; Martínez Poyatos et al., 2001) and
592 could also be recorded by RSCM thermometry in our samples.

593

594 **5.3. Second tectonothermal event (middle/upper Carboniferous M₂)**

595 The mineralogy of the Santa Iria samples (Qz + Fsp + Ms + Chl ± C/S) is compatible with
596 very low- to low-grade conditions. The K-white mica “crystallinity” values (0.20-0.26 Δ°2θ;
597 average 0.23) point to lower epizone conditions, very close to the boundary with the
598 anchizone (≈300 °C; Frey, 1987; Kisch, 1987). The temperatures calculated by RSCM in two
599 samples (315 and 330 °C) are compatible with the KI data of XRD analysis.

600 Our metamorphic data corroborate the existence of an unconformity between the lower and
601 upper formations of the Pulo do Lobo domain (Pérez-Cáceres et al., 2015). Table 2
602 summarizes the relationship between deformation and metamorphism of the Pulo do Lobo
603 domain in the context of the Variscan tectonic evolution of SW Iberia (Pérez-Cáceres et al.,
604 2015). The lower formations record a Devonian tectonothermal event that reached epizone
605 or lower greenschist facies conditions (M₁ with generalized phyllosilicate growth at
606 temperatures as high as 450 °C), whilewhereas the overlying upper formation records a
607 middle/upper Carboniferous tectonothermal event close to the anchizone/epizone
608 boundary (M₂ with small-sized phyllosilicate growth at temperatures ≈300-330 °C; Table 1).
609 Obviously, M₂ also affected somehow the lower formations, being, at least in part, the
610 responsible for the observed retrogression of M₁ chlorite and/or crystallization of new
611 chlorites at lower temperature.

612

613 **5.4. Pressure conditions**

614 The measured *b*-cell parameters of K-white mica (in a short range between 8.991-9.002 Å;
615 average 8.996; standard deviation 0.003) are very similar in the lower and upper formations
616 of the Pulo do Lobo domain. Thus, the *b* parameter is consistently homogeneous shows little
617 variation and reflects very low phengite substitution in mica, as expected at low-pressure

618 settings (Potel et al., 2006, 2016), near the intermediate pressure gradient boundary (Guidotti
619 and Sassi, 1986).

620 In agreement with the low *b*-cell parameters, the composition of K-white mica is close to
621 muscovite with very low celadonite and higher pyrophyllite content (Fig. 5a), as expected for
622 illite-rich mica formed at low-pressure gradients. In the case of high- or medium-pressure
623 conditions, a continuous trend in mica compositions would ~~be found reflecting~~ reflect the
624 decompression path after the peak pressure, while the *b*-cell parameter would represent an
625 average value of the range of mica compositions found in the sample (Abad et al., 2003b).
626 On the contrary, at low-pressure settings, the overall range of recorded pressure is very
627 ~~short restricted~~ and micas present similar compositions and *b*-cell parameters among the
628 various samples, as in the case of the Pulo do Lobo samples (Figs. 5a and 6, and Table 1).

629 The Pulo do Lobo domain has been classically interpreted as a pre-collisional subduction-
630 related accretionary prism, based on the MORB geochemistry of their mafic rocks (see
631 section 2.1). According to this classical interpretation, features typical of modern subduction
632 systems should be expected, such as high-pressure metamorphic gradient remnants of partial
633 subduction/exhumation in an accretionary wedge (e.g., Platt, 1986; Ernst, 2005), or slices of
634 oceanic slab-derived lithologies (varied mid-ocean ridge metaigneous lithologies and also
635 deep ocean bottom metasediments). Thus, recent works on the Makran accretionary prism
636 (Omrani et al., 2017) and the subduction system of Japan (Endo and Wallis, 2017) describe
637 an accretionary mélangé complex composed of pelagic sedimentary rocks, ophiolites,
638 greenschists, amphibolites, and blueschists with high-pressure minerals such as lawsonite and
639 glaucophane. ~~On the contrary~~ However, most of the geological data concerning the Pulo do
640 Lobo domain do not ~~back up support~~ such ~~an~~ interpretation (see section 2.1), and our new
641 ~~results about estimates of~~ pressure conditions are also in disagreement. The only ~~suspect~~
642 ~~of evidence supporting~~ high-pressure gradient in the Pulo do Lobo domain is the
643 interpretation of some rhomboidal aggregates of epidote ~~porphyroblasts~~ porphyroblasts as the
644 remnants of supposed lawsonite grown ~~previously prior~~ to S₂ in some samples ~~of in the~~ Pulo
645 do Lobo mafic schists (Rubio Pascual et al., 2013). However, no analytical data have been
646 presented to support the ~~presence of~~ lawsonite pseudomorphs.

647

648 6. Conclusions

649 Eighteen samples of metapelites from the Pulo do Lobo domain have been studied to
650 ~~characterize determine~~ their Variscan low-grade ~~metamorphism~~ metamorphic conditions.
651 The microstructural analysis of the samples of the lower formations (Devonian Pulo do Lobo
652 and Ribeira de Limas) shows the existence of two superposed low-grade tectonothermal
653 events with associated foliation and phyllosilicate growth (S₁-M₁ and S₂-M₂; Table 2). M₂ was
654 less intense, ~~being and is~~ the only event that affected the overlying Carboniferous Santa Iria
655 ~~formation~~ Formation. The regional geology also shows that a Mississippian thermal
656 (magmatic-~~derived~~) event occurred in-between M₁ and M₂.

657 M₁ and M₂ correspond to ~~the~~ chlorite zone ~~metamorphism~~, but M₁ ~~entered the attained~~
658 epizone ~~conditions~~ (greenschists facies with temperatures up to ≈450 °C), while M₂ did not
659 exceed the anchizone-epizone boundary (≈300 °C).

660 The ~~temperature~~temperature estimates obtained from RSCM are higher ~~compared to~~than
661 the ~~ones derived~~estimates obtained from chlorite geothermometry and white mica data.
662 ~~The~~This discrepancy can be explained by the fact that RSCM records the ~~true~~-maximum
663 temperature, ~~being because of it is~~ not affected by retrogression ~~as~~, in contrast with the other
664 methods ~~do~~. In addition, this difference can be the consequence of the high sensitivity of
665 CM to quickly re-equilibrate at maximum temperatures during short-duration thermal
666 events ~~due to~~such as the magmatic intrusions emplaced during the Mississippian thermal
667 event.

668 Thermodynamic disequilibrium between white mica and chlorite has precluded their use for
669 geothermobarometry, and a variety of data (including the existence of relic high-temperature
670 chlorite cores, the presence of chlorite/smectite mixed layers, or the very-low temperatures
671 calculated with chlorite geothermometers) indicate chlorite retrogression after M₁
672 metamorphic climax and crystallization of new chlorite grains at lower temperature.

673 The low-pressure conditions derived from white mica indicators (very low celadonite content
674 and *b*-cell values) are incompatible with the high-pressure metamorphic gradient expected in
675 a subduction-related accretionary wedge, which has been the classical interpretation of the
676 Pulo do Lobo domain. Instead we interpret that the Pulo do Lobo rocks were deposited in
677 a platform setting located near the northern border of Avalonia, but they were never involved
678 in the subduction-related processes of the SPZ/OMZ suture.

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691

692 **Figure captions**

693 **Figure 1.** a) Location of the studied area in the SW of the Iberian Massif (in grey). CIZ: Central
694 Iberian Zone, OMZ: Ossa-Morena Zone, SPZ: South Portuguese Zone. b) Geological map of the
695 Pulo do Lobo domain and other units related to the OMZ/SPZ boundary with indications of the
696 two cross-sections studied and collected samples. c.1-2) Geological cross-sections of the Pulo do
697 Lobo domain (see b for location) (modified from Martínez Poza et al., 2012 and Pérez-Cáceres et al.,
698 2015). Numbered red circles in b-c locate the samples studied. Big circles show the KI values for 10
699 Å reflection peaks of K-white mica and the average RSCM temperatures, with the relative colour bar
700 according to the results shown in Table 1. BAA: Beja-Acebuches Amphibolites, M: metabasalts, PL:

701 Pulo do Lobo [formationFormation](#), RL: Ribeira de Limas [formationFormation](#), SI: Santa Iría
702 [formationFormation](#).

703 **Figure 2.** Pictures of the Pulo do Lobo rocks illustrating deformation at outcrop scale: a) Pulo do
704 Lobo formation, b) Ribeira de Limas [formationFormation](#), c) Santa Iría [formationFormation](#).
705 Microphotographs from thin-sections: d) Cross-polarized light image of sample PLB-84 (Pulo do
706 Lobo [formationFormation](#)), e) SEM-BSE image of sample PLB-88 (Ribeira de Limas
707 [formationFormation](#)), f) Cross-polarized light images of sample PLB-71 (Santa Iría
708 [formationFormation](#)).

709 **Figure 3.** Representative Raman spectra of CM across the Pulo do Lobo domain from low
710 temperature (bottom; Santa Iría [formationFormation](#)) to high temperature (top; lower formations)
711 including the average maximum temperatures (°C) for each sample. Vertical scale for spectrum
712 intensity is arbitrary. See Fig. 1 for sample location and Table 1 and supplementary information for
713 RSCM data.

714 **Figure 4.** X-Ray maps of the three selected samples analyzed by EPMA and processed with
715 XMapTools. The samples belong to the lower formations of the Pulo do Lobo domain (sample PLB-
716 88: Ribeira de Limas [formationFormation](#); samples PLB-84 and PLB-93: Pulo do Lobo
717 [formationFormation](#); the latter (PLB-93) is close to Early Carboniferous igneous intrusions). a)
718 EPMA BSE photographs. b) Mineral maps. c) Al^{IV} content map in chlorites, which increases with
719 temperature. The white square highlights the zonation of a chlorite grain from core to rim. d)
720 Temperature maps of chlorite using the Lanari et al. (2014a) geothermometer assuming all iron as
721 ferrous. White squares show selected areas illustrating higher-temperature chlorite cores. Red squares
722 show the selected areas (representative of S₁ foliation) used for chlorite-quartz-water
723 geothermometric calculations shown in Fig. 7.

724 **Figure 5.** Ternary plots of all the analyzed white micas (a) (Cel: celadonite, Mus: muscovite, Prl:
725 pyrophyllite) and chlorite (b) (Cli+Daph: clinocllore + daphnite, Am: amesite, Sud: sudoite) plotted
726 with the XmapTools TriPlot3D module. Colour bars refer to the number of mica/chlorite pixels
727 analyzed.

728 **Figure 6.** Compositional diagram of white micas showing Na/(Na+K) vs Si/Al (atomic ratios) for 31
729 EPMA point analyses from seven samples of the lower formations of the Pulo do Lobo domain
730 (different symbology, for each sample). Point analyses were obtained on the microprobe at the
731 University of Huelva (Spain). Qualitative information about temperature and pressure conditions are
732 respectively according to Guidotti et al. (1994), Coggon and Holland (2002), Parra et al. (2002),
733 Massonne and Schreyer (1987) and Massonne and Szpurka (1997).

734 **Figure 7.** Histograms of temperatures obtained using the chlorite-quartz-water geothermometer of
735 Vidal et al. (2006) (a) and Bourdelle et al. (2013) (b) on selected representative S₁ chlorites (see red
736 squares in Fig. 4d for location). *n* represents the number of chlorites that could be used for each
737 calibration. The number of analyses is lower in those with Vidal et al. (2006) approach because the
738 assumption that the Si content of chlorite is lower than 3 apfu.

739

740 Table captions

741 **Table 1.** Samples and results obtained by XRD (<2 μm fraction), white mica and chlorite
742 compositions, temperature ranges from chlorite thermometry, and average RSCM thermometry.
743 Basel KI values and average RSCM temperatures show a relative colour-bar scale. Mineral
744 abbreviations according to Whitney & Evans (2010). Qz: Quartz, Ms: Muscovite, Fsp: Feldspar, Chl:

745 Chlorite, Pg: paragonite, C/S: chlorite-smectite mixed layers, Cel: celadonite, Prl: pyrophyllite,
746 Cli+Daph: clinocllore + daphnite, Am: amesite, Sud: sudoite, Std Dv: standard deviation.

747 **Table 2.** Summary of the tectonometamorphic Variscan evolution of the Pulo do Lobo domain.

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