

1 **Evolution of the Iberian Massif as deduced from its crustal thickness and**
2 **geometry of a mid-crustal (Conrad) discontinuity**

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13 **Abstract**

14 Normal incidence seismic data provide the best images of the crust and lithosphere. When
15 properly designed and continuous, these sections greatly contribute to understanding the
16 geometry of orogens and, together with surface geology, to unravel their evolution. In this
17 paper we present the most complete transect, to date, of the Iberian Massif, the westernmost
18 exposure of the European Variscides. Despite the heterogeneity of the dataset, acquired
19 during the last 30 years, the images resulting from reprocessing with a homogeneous workflow
20 allow us to clearly define the crustal thickness and its internal architecture. The Iberian Massif
21 crust, formed by the amalgamation of continental pieces belonging to Gondwana and
22 Laurussia (Avalonian margin), is well structured in upper and lower crust. A conspicuous mid-
23 crustal discontinuity is clearly defined by the top of the reflective lower crust and by the
24 asymptotic geometry of reflections that merge into it, suggesting that it has often acted as a
25 detachment. The geometry and position of this discontinuity can give us insights on the
26 evolution of the orogen, i.e. of the magnitude of compression and the effects and extent of
27 later Variscan gravitational collapse. Also, the limited thickness of the lower crust below, in
28 central and NW Iberia, might have constrained the response of the Iberian microplate to
29 Alpine shortening. This discontinuity, featuring a Vp increase, is here observed as an orogen-
30 scale boundary with characteristics compatible with those of the worldwide debated, Conrad
31 discontinuity.

32 **Keywords:** Iberian Massif, vertical incidence seismic data, mid-crustal detachment, Conrad
33 discontinuity, geodynamic evolution

34 **Glossary:**

- 35 CIA: Central Iberian Arc
36 CIZ: Central Iberian Zone
37 CU: Central Unit
38 CZ: Cantabrian Zone
39 DB: Duero Basin
40 GTMZ: Galicia Tras-os-Montes Zone
41 IAA: Ibero-Armorican Arc

42 ICS: Iberian Central System
43 NI: Normal Incidence
44 OMA: Ossa-Morena Zone
45 SPZ: South Portuguese Zone
46 TB: Tajo Basin
47 WA: Wide Angle
48 WALZ: West Asturian-Leonese Zone
49

50 1. Introduction

51 In the last 35 years, controlled source seismic data have greatly contributed to the
52 understanding of the European Variscides. National research programs like DEKORP (Bortfeld,
53 1985; DEKORP Research Group, 1987; Franke et al., 1990; Oncken, 1998), BIRPS and ECORS
54 (BIRPS and ECORS, 1986) have sampled this orogen providing a detailed picture of its
55 lithospheric architecture. In the Iberian Massif, normal incidence (NI) seismic reflection profiles
56 often acquired with coincident wide angle (WA) reflection/refraction seismic data have
57 allowed scientists to depict its crustal structure, infer its P and S_v-waves velocity distribution,
58 place constraints on its geodynamic evolution, visualize the accommodation pattern of
59 shortening at different crustal levels and, sometimes, deduce the effect of Alpine reactivation
60 on this Paleozoic orogen.

61 In this regard, seismic datasets acquired in the Iberian Massif (DeFelipe et al., 2020) ~~from~~
62 ~~within the programs~~ ESCIN (Ayarza et al., 1998, 2004; Pérez-Estaún et al., 1991; Pulgar et al.,
63 1996), IBERSEIS (Flecha et al., 2009; Palomeras et al., 2009; Simancas et al., 2003), ALCUDIA
64 (e.g., Ehsan et al., 2014, 2015; Martínez Poyatos et al., 2012) and CIMDEF ~~projects~~ (Andrés et
65 al., 2019) have helped ~~so far~~ to identify several outstanding features such as, i) clear
66 differences in the intensity and geometry of reflectivity at upper and lower crustal levels, ii)
67 contrasting deformation patterns deduced from the good correlation between reflectivity and
68 upper crustal structures as, regardless of the many factors that trigger the concentration of
69 deformation along narrow thrust zones (Butler and Mazzoli, 2006), the latter often follow
70 lithological boundaries in the Iberian Massif (e.g., Alonso, 1987), thus being candidates to
71 appear as outstanding reflections ~~in NI seismic sections~~, and iii) a very reflective and
72 sometimes thick lower crust, even in areas where the upper crust is weakly deformed. In order
73 to explain these features, decoupling of the upper and lower crust, ~~eased by a mid-crustal~~
74 ~~detachment~~, has been invoked along large parts of the sampled area (Simancas et al., 2013). ~~The~~
75 ~~existence of a mid-crustal detachment has been addressed by these authors as the reason~~
76 ~~why different shortening mechanisms exist at different crustal levels.~~ However, coupled crustal
77 deformation has been inferred ~~for in~~ NW Iberia, where large crustal thickening took place.
78 Also, no inference has been made on how this detachment acted during later Alpine
79 deformation in the central and northern Iberian Massif.

80 In this paper, we present a new composite seismic section of the Iberian Massif that integrates
81 results of two new datasets: CIMDEF and ALCUDIA WA. The former fill the gaps ~~of between~~
82 areas previously unexplored, like Central Iberia, ~~altogether thus contributing to providing~~
83 ~~provide~~ an almost complete transect. In addition, we include the N-S₂ ESCIN-2 seismic profile

84 and new time-migrated sections of all datasets, some of them, as yet, unmigrated (e.g., ESCIN-
85 1). Although the detailed interpretation of the upper crustal reflectivity and its correlation with
86 the surface geology does not significantly change, we revisit the deep reflectivity and redefine
87 the extension and implications of a mid-crustal discontinuity that, in our view, exists below the
88 entire Iberian Massif, affecting the two continents involved in the Variscan Orogeny and
89 certainly playing a critical role in the decoupling between upper and lower crustal
90 deformation. Finally, we infer the geometry and nature of this feature, discuss its tectonic
91 significance and its role during the younger Alpine Orogeny, and relate it with the long-
92 debated (e.g., Finlayson et al., 1984; Litak and Brown, 1989; Wever, 1989; Xiaobo and Tae
93 Kyung, 2010) Conrad discontinuity (Conrad, 1925) of the classic continental seismology.

94 2. Geological setting

95 The Iberian Massif represents the westernmost outcrop of the European Variscides, exposing
96 an almost complete section of this Paleozoic orogen. It is divided into six zones (Fig. 1; Arenas
97 et al., 1988; Farias et al., 1987; Julivert et al., 1972) that from N to S and E to W are: Cantabrian
98 (CZ), West Asturian-Leonese (WALZ), Galicia-Trás-os-Montes (GTMZ), Central Iberian (CIZ),
99 Ossa-Morena (OMZ) and South Portuguese (SPZ). The CZ and the SPZ represent ~~the~~ external
100 zones whereas the rest represent the hinterland. The CZ, WALZ and CIZ belong to the northern
101 margin of the Paleozoic Gondwana. The GTMZ represents the remnants of a large nappe stack
102 formed by pieces of the outermost ~~margin of the~~ Gondwana margin, i.e., a pulled-apart peri-
103 Gondwanan terrane and ~~oceanic-ophiolitic~~ units derived from the Rheic oceanic realm
104 separating them. ~~They were emplaced above the autochthonous CIZ in NW Iberia and are~~
105 ~~preserved as a large klippen at the core of late Variscan synforms (Martínez Catalán et al.,~~
106 ~~1997, 2007; Ries and Shackleton, 1971). Its rootless suture (ophiolitic units) is thought to~~
107 ~~represent a branch of the Rheic oceanic realm (Martínez Catalán et al., 1997).~~ The OMZ has
108 been interpreted as a continental fragment that rifted and probably drifted away from
109 Gondwana (Matte, 2001; Robardet, 2002), docking back later to the CIZ giving rise to a suture
110 ~~and thus representing another peri-Gondwanan terrane~~ (Azor et al., 1994; Gómez-Pugnaire et
111 al., 2003; Simancas et al., 2001). Finally, the SPZ is thought to be separated from the OMZ also
112 by the Rheic Ocean suture (Simancas et al., 2003) which was later overprinted by early
113 Carboniferous extension ~~accompanied by mafic intrusions~~ (Azor et al., 2008) and by
114 transpression (Pérez-Cáceres et al., 2015). In this context, the basement of the SPZ represents
115 a fragment of Avalonia (Braid et al., 2011, 2012; Pereira et al., 2014; Rodrigues et al., 2015),
116 and its affinity with the Rheno-Hercynian Zone in Germany (Franke, 2000; Franke et al., 1990)
117 further supports this correlation.

118 From a tectonic point of view, the Iberian Massif shows evidence of pre-Variscan activity. The
119 Cadomian Neoproterozoic event is characterized by continental arc magmatism, deformation
120 and metamorphism (Bandrés et al., 2004; Dallmeyer and Quesada, 1992; Ochsner, 1993). It
121 developed above a previous non-outcropping continental crust, and formed the basement
122 above which the Ediacaran and Paleozoic sedimentation took place, favored by Ediacaran-
123 Cambrian and Cambro-Ordovician rifting which developed a wide continental platform
124 (Linnemann et al., 2008; Sánchez-García et al., 2008, 2010). However, most of the outcropping
125 tectonic features are the result of the Devonian and Carboniferous collision between
126 Gondwana, some peri-Gondwanan terranes and the Avalonian border of Laurentia, which

127 resulted in the Variscan Orogen (Matte, 2001). The deformation associated to the latter was
128 diachronous along the Iberian Massif. Next, we describe the most important tectonic and
129 stratigraphic features of this part of the European Variscides, by tectonic zones and from N to S
130 (in present coordinates), ~~together with its most important tectonic and stratigraphic features.~~

131 2.1. Cantabrian Zone (CZ)

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132 The CZ (Fig. 1) is an external zone located at the core of the Ibero-Armorican Arc (IAA; Dias and
133 Ribeiro, 1995; Lotze, 1929; Matte and Ribeiro, 1975; Stille, 1924). It is a thin-skinned thrust and
134 fold belt with a transport direction towards the foreland ~~in to~~ the E, and overprinted by the
135 oroclinal folding giving rise to the IAA (Alonso et al., 2009; Pérez-Estaún et al., 1988).
136 Stratigraphically, it is characterized by a Precambrian sequence, outcropping at its western
137 part and overlain by a Paleozoic stratigraphic succession that ranges from the Cambrian to a
138 well-developed Carboniferous: pre-orogenic up to early Carboniferous, and syn-orogenic in the
139 Upper Carboniferous (Sánchez de Posada et al., 1990; Truyols et al., 1990). Scarce tholeiitic
140 and alkaline magmatism is related to Cambro-Ordovician rifting (Corretgé and Suárez, 1990).
141 No regional metamorphism accompanied deformation, indicating shallow crustal conditions. In
142 this area, deformation began at ~325-320 Ma, in the Late Mississippian (Dallmeyer et al.,
143 1997), and the emplacement of nappes that characterize the deformation and the formation
144 of folds ~~within each sequence~~ took place between the Westphalian B and the Stephanian (313-
145 300 Ma), in the Upper Carboniferous (Pérez-Estaún et al., 1988). An extensional episode
146 related to the end of the orogeny led to the formation of Permian Basins (Martínez García,
147 1981). Later on, extension related to the opening of the Bay of Biscay triggered the
148 development of deep Cretaceous basins (Quintana et al., 2015; Rat, 1988). Alpine tectonics
149 uplifted the Pyrenean-Cantabrian range from the end of the Late Cretaceous to the Miocene
150 (DeFelipe et al., 2019; Teixell et al., 2018 and references therein) reactivating Variscan thrusts
151 and Mesozoic normal faults (Gallastegui et al., 2016).

152 2.2. West Asturias-Leonese Zone (WALZ)

Con formato: Fuente: Negrita

153 The WALZ lies to the W of the CZ (Fig. 1). Stratigraphically, it consists of a Neoproterozoic
154 terrigenous sequence unconformably overlain by a Paleozoic platform succession that ranges
155 from the Cambrian to the Lower Devonian (Pérez-Estaún et al., 1990), being much thicker than
156 that of the CZ. These sediments were actively deformed along three compressional (C1, C2,
157 and C3) and two extensional (E1 and E2) phases during the Variscan Orogeny (Martínez
158 Catalán et al., 2014). Large E vergent folds witness the C1 related compression (360-340 Ma).
159 Those were later affected by E vergence thrusts resulting from ongoing shortening (345-325
160 Ma). Crustal thickening followed by thermal relaxation led to syn-orogenic extension during E1
161 (330-315 Ma). A last compressional episode (C3, 315-305 Ma) produced upright folds
162 associated with wrench shear zones while simultaneous extension (E2, 315-300 Ma)
163 continued, characterizing the latest stages of the orogeny in this area. Crustal melting
164 triggered by compression and thickening (C1-C3) led to extension (E1-E2) and to the intrusion
165 of granitoids in the western part of the WALZ. Thermal models show that the crust could have
166 started to melt within 30 Ma after the onset of crustal thickening, which is then constrained by
167 the ages of Variscan granitoids (Alcock et al., 2009).

168 2.3. The Central Iberian Zone (CIZ)

Con formato: Fuente: Negrita

169 The CIZ is the largest of the Iberian Massif zones. Curvature of magnetic anomalies and that of
170 early (C1) Variscan folds depict the Central Iberian Arc (CIA; Martínez Catalán, 2011a, 2011b),
171 partly explaining the width of this internal zone (Fig. 1). The stratigraphic sequence differs
172 from N to S: Ordovician felsic metavolcanic, subvolcanic and intrusives (Diez-Montes et al.,
173 2010) represent the most ancient lithologies outcropping in the N, defining the 'Ollo de Sapo'
174 domain whereas to the S, Upper-Proterozoic-Lower Cambrian metasediments outcrop (Díez
175 Balda et al., 1995), defining the 'Schist-Greywacke Complex' domain. The pre-orogenic
176 sedimentary sequence continues to the Devonian, followed by a syn-orogenic Carboniferous
177 sequence (Martínez Catalán et al., 2004; Robardet, 2002). This area represents a relatively
178 stable Gondwana margin characterized by the Early Ordovician extension that opened the
179 Rheic Ocean and allowed intrusion of the essentially felsic magmas of the Ollo de Sapo Fm.
180 (Díez Montes et al., 2010). The deformation phases described for the WALZ affected most of
181 the CIZ although C1, C2 and E1 are somewhat older, according to the propagation of
182 deformation from the hinterland. Slight differences in the importance of phases can also be
183 found to the center and S (Martínez Catalán et al., 2019), allowing the CIZ to be divided in two
184 zones. In the NW, intense recumbent C1 folds and important C2 thrusts related to the
185 emplacement of the GTMZ occur. Outcropping rocks show epi- to catazonal metamorphism
186 and ductile detachments. Gneiss domes of both E1 and E2 extensional phases exist, evidencing
187 significant crustal thinning, and Variscan granitoids are abundant. To the S, C1 folds are
188 upright, C2 deformation is limited to the southernmost part, and upright C3 folds are the most
189 important structures (Martínez Catalán et al., 2012; Martínez Poyatos, 2002). Metamorphism
190 is generally weak and the amount of granitoids decreases, except in the Iberian Central System
191 (ICS). Here extension postdates C3 upright folding and thus, it is considered E2. Alpine
192 tectonics in the CIZ reactivated previous Variscan fractures and triggered the development of
193 the Iberian Central System (ICS) mountain belt (de Vicente et al., 1996), allowing the products
194 of syn- and post-E2 crustal melting to outcrop in large areas.

195 2.4. The Galicia-Tras-os-Montes Zone (GTMZ)

196 The GTMZ is represented by five klippen that are ~~the~~ remnants of the emplacement of a thick
197 nappe stack on top of the CIZ. This includes, from bottom to top, a relatively distal part of the
198 northern Gondwana margin (Parautochthon), the outermost edge of that margin (Lower
199 Allochthon), a few oceanic units of Cambro-Ordovician and Lower Devonian age (Middle
200 Allochthon) and a peri-Gondwanan terrane with magmatic evidences of Cambro-Ordovician
201 rifting and a continental arc setting (Upper Allochthon). Several units show high-P
202 metamorphism reflecting subduction of the ocean represented by the Middle Allochthon and
203 involving also the Upper and Lower ones (Arenas et al., 2007; Gómez Barreiro et al., 2007;
204 Martínez Catalán et al., 2007; Sánchez Martínez et al., 2007). Ongoing subduction during most
205 of the Devonian (400-365 Ma) built an accretionary wedge that was subsequently emplaced on
206 top of the CIZ during the early Carboniferous (C1-C2 events, -360-340 Ma).

207 2.5. The Ossa-Morena Zone (OMZ)

208 The boundary between the CIZ and the OMZ (the Badajoz Córdoba Shear Zone) has been
209 largely interpreted as a suture (Gómez-Pugnaire et al., 2003; Simancas et al., 2001), although

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210 no true oceanic units have been identified. It includes amphibolites of oceanic affinity from the
211 early Paleozoic, as well as eclogite relics.

212 In general, SW Iberia features ~~outcropping lithologies ranges~~ from the Upper Precambrian to
213 the Upper Carboniferous, with an angular unconformity at the Lower Carboniferous. In the
214 OMZ, the Serie Negra (Black Series) is a thick Neoproterozoic sequence that includes graphitic
215 quartzites and schists and underwent Cadomian arc-related magmatism and regional
216 metamorphism (Dallmeyer and Quesada, 1992; Ochsner, 1993; Quesada and Dallmeyer, 1994).
217 The pre-orogenic Paleozoic sequence is rather complete and was deposited at the peri-
218 Gondwanan platform, as for the CIZ, although differences in the faunal content and in the
219 Paleozoic facies, generally more pelitic in the OMZ, point to a more distal position (Robardet,
220 2002; Robardet and Gutiérrez Marco, 1990). Ediacaran-Cambrian and Cambro-Ordovician
221 magmatism reflects two rifting events, ~~t-~~ The latter ~~is the most important one, it includes~~
222 alkaline magmatism ~~and is~~ related to the opening of the Rheic Ocean (García Casquero et al.,
223 1985; Ochsner, 1993; Sánchez-García et al., 2008, 2010). The first deformation event, of
224 Devonian age, formed overturned and recumbent folds and thrust faults with SW vergence
225 (Expósito et al., 2002, 2003). Syn-orogenic, early Carboniferous basins developed in an
226 extensional context and are related to calc-alkaline volcanism and magmatism (Casquet et al.,
227 2001). These deposits unconformably overlay ~~the~~ early folds and thrusts. Later, deformation
228 continued with middle and upper Carboniferous sinistral transpression and associated upright
229 NW-SE folds.

230 A salient seismic reflector, the Iberseis Reflective Body (IRB, Carbonell et al., 2004; Simancas et
231 al., 2003), ~~seems to be the result of a mantle derived intrusion~~ located along a mid-crustal
232 detachment, ~~seems to be a mantle-derived intrusion at around 350-340 Ma. It was~~ emplaced
233 ~~in the context of~~ during early Carboniferous extension (350-340 Ma) ~~in the SW of the Iberian~~
234 ~~Massif, while the~~ At this time, the hinterland to the NW was undergoing the first stages of
235 compression (C1-C2). ~~Magmatic activity in the SW triggered a high T/low P metamorphism~~
236 ~~that, otherwise, has a low grade elsewhere in the OMZ (Díaz Azpiroz et al., 2006; Pereira et al.,~~
237 ~~2009).~~

238 2.6. The South-Portugese Zone (SPZ)

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239 The boundary between the OMZ and the SPZ has been long understood as a suture on the
240 basis of geometric assumptions (e.g., Carvalho, 1972). Later evidences have reinforced this
241 point of view suggesting that ~~the above mentioned boundary~~ represents the remnants of the
242 Rheic Ocean, although Carboniferous transtension and transpression have largely obliterated it
243 (Pérez-Cáceres et al., 2015 and references therein).

244 The SPZ is a Variscan foredeep basin strongly deformed by thin-skinned thrust tectonics, and is
245 usually correlated with the Rhenohercynian Zone of Kossmat (1927) ~~in the Bohemian Massif~~.
246 It features wide outcrops of low or very low grade Devonian phyllites, quartzites and
247 sandstones overlain by a lower Carboniferous (Early Mississippian) volcano-sedimentary
248 sequence topped by middle and upper Carboniferous flysch (Oliveira, 1990). From a tectonic
249 point of view, it is characterized by Carboniferous S vergent thrusts and folds, the latter
250 featuring axial traces oblique to the northern boundary of the zone, evidencing transpression

251 (Simancas et al., 2003 and references therein). Deformation propagated towards the S ~~along~~
252 ~~during~~ the lower and upper Carboniferous (Oliveira, 1990).

253 Although the start of the Variscan collision seems to have been frontal or maybe right-lateral
254 in most of Europe (Shelley and Bossière, 2000), surface geology and interpretation of seismic
255 data evidences the existence of relevant left lateral transpression and oblique-slip syn-
256 metamorphic shear zones in the OMZ, SPZ and their boundaries (Pérez-Cáceres et al., 2016;
257 Simancas et al., 2003 and references therein). In the OMZ, folds and thrusts witnessing
258 Devonian and early Carboniferous compression are oblique to the OMZ/SPZ boundary,
259 indicating a transpressional setting. ~~These features are disrupted by later Mississippian~~
260 ~~transtensional tectonics (Expósito et al., 2002) that gave way to the intrusion of the Beja-~~
261 ~~Acebuches mafic and ultramafic rocks (Azor et al., 2008). Convergence resumed soon after,~~
262 ~~leading to the emplacement of the Beja-Acebuches unit onto the OMZ (Pérez-Cáceres et al.,~~
263 ~~2015). Inside the OMZ~~ There, Devonian and Carboniferous left lateral deformation accounts for
264 ~400 km, higher than perpendicular shortening. Likewise, inside the SPZ, left-lateral
265 displacement is estimated to reach 90 km whereas the orthogonal one amounts ~60 km
266 (Pérez-Cáceres et al., 2016).

267 **3. Geophysical setting: Existing datasets, their reprocessing and a brief description**

268 **3.1. Seismic datasets sampling the Iberian Massif**

269 Since the early 1990s, the Iberian Massif has been sampled by different controlled source
270 seismic experiments (DeFelipe et al., 2020): the ESCIN (1991-1992), IBERSEIS (2000 and 2003),
271 and ALCUDIA (2007-2012) experiments acquired normal incidence (NI) and coincident wide-
272 angle (WA) data. The latest project, carried out with the target of understanding the structure
273 and effect of the Alpine reactivation across the central part of the Iberian Massif, is the
274 CIMDEF experiment (2017-2019). It recorded densely spaced controlled source WA reflection
275 and natural source (earthquakes and noise) seismic data. However, the acquisition of
276 coincident NI data along this transect has not currently been planned, regardless of its
277 potential quality and relevance, due to the relatively high costs of this kind of experiment.

278 From N to S, and from E to W, the ESCIN project sampled the northern part of the Iberian
279 Massif (Fig. 1). Profile ESCIN-1 (1991) is an onshore E-W line crossing the CZ from its eastern,
280 most external part to its boundary with the WALZ to the W; Profile ESCIN-2 (1991) is an
281 onshore N-S profile crossing the most external and eastern part of the CZ and reaching the
282 northern end of the Duero Basin (DB) to the S, which represents the Cantabrian Mountains
283 foreland basin. The ESCIN-3 (1992) profiles sampled the WALZ and the CIZ along the northern
284 Iberia shelf. Although it consists of three parts (ESCIN-3.1, 3.2 and 3.3) only the easternmost
285 ~~ones (3.2 and 3.3.) are relevant for this study of the Variscan crust~~ and thus, included here.
286 ESCIN-3.3 crossed the entire WALZ to its western boundary with the CIZ, which in this area was
287 surveyed by the ESCIN-3.2. Geographically, the latter also sampled the allochthonous GTMZ.
288 But as this is an offshore profile, it shows no evidences of the presence of the GTMZ, and most
289 of the imaged crust corresponds to that of its relative autochthon, the CIZ.

290 A significant geographical and methodological gap exists between the ESCIN profiles to the N
291 and the location of the CIMDEF experiment (Fig. 1). The latter crosses central Iberia from the N

292 part of the CIZ, then samples the DB down to the ICS, and goes on S across the Tajo Basin (TB)
293 till it reaches again the CIZ metasediments ~~to the S of the ICS.~~

294 In the southern part of the Iberian Massif, the onshore ALCUDIA seismic line (NI and WA),
295 striking NE-SW, was acquired across the CIZ, going from the S of the ICS to the boundary with
296 the OMZ. Finally, the NE-SW IBERSEIS dataset (NI and WA) is also an onshore profile ~~that~~
297 ~~partially~~structurally overlaps ~~the same structures as~~ the SW end of the ALCUDIA line although
298 with some 50 km of offset to the W. This seismic line samples the southern part of the CIZ, the
299 OMZ and the SPZ.

300 Altogether these seismic profiles account for a ~1500 km long seismic transect geared to
301 understand the crustal and, in places, lithospheric structure of the Iberian Massif and to
302 constrain its evolution.

303 3.2. Processing of datasets

304 The data used in this work have been acquired at different times, have different characteristics
305 (onshore and offshore) and accordingly exhibit very heterogeneous quality. Table 1 shows
306 ~~their~~ acquisition parameters ~~of all these datasets.~~ The most outstanding differences are: i) the
307 quality and characteristics of the offshore (ESCIN-3) vs the onshore data, ii) the difference
308 between the low fold (30) ESCIN ~~1, ESCIN 2 and ESCIN 3~~ data, acquired with ~~an~~ explosives
309 sources and airguns, ~~respectively~~ and the high fold (>60) IBERSEIS and ALCUDIA datasets,
310 which used Vibroseis trucks as source of energy, and iii) the fact that the CIMDEF dataset lacks
311 NI data and only provides lower resolution noise and earthquake data, since WA profiles are,
312 as yet, unpublished. Thus, reprocessing all NI data was mandatory, at least at stack and post-
313 stack level. Figure 2 shows the processing flow followed to homogenize the display of datasets
314 while preserving the true amplitude (Martínez García, 2019). The software package used for
315 reprocessing was GLOBE Claritas (www.globeclaritas.com/) and the most important steps were
316 related to frequency filtering, amplitude weighting and equalization, Kirchhoff time migration
317 and coherency filtering (Fig. 2). In addition, up to 20 multi-trace attribute analysis were tested
318 with the goal to enhance structural and lithological impedance contrasts that allowed to
319 improve the interpretation (Chopra and Alexeev, 2005; Taner and Sheriff, 1977). Although this
320 methodology has been mostly used in sedimentary reservoirs, we have seen that the
321 application of these techniques can enhance the continuity of reflections and help to identify
322 different types of crust, thus easing the interpretation. Some of the boundaries resulting from
323 this attribute analysis (e.g., variance and chaos attribute filters, the former estimating the local
324 variance in the signal and the latter measuring the lack of organization in the reflectivity) are
325 included in the interpretations.

326 3.3. Description of the seismic sections

327 The NI datasets included in this paper have already been presented, so the reader will be
328 referred, in every sub-section, to previous publications that include detailed descriptions of
329 pre-stack processing and interpretations. Here we will just focus on those features that are
330 essential to our interpretation.

Código de campo cambiado

331 Geological cross-sections coincident with reprocessed time-migrated sections and their
332 interpretations are presented in figure 3 (ESCIN-1), figure 4 (ESCIN-2), figure 5 (ESCIN-3.3),
333 figure 6 (ESCIN-3.2), figure 7 (ALCUDIA), and figure 8 (IBERSEIS). Migration velocities (Fig. 2) are
334 average crustal values as calculated from coincident or nearby WA models. Depth conversion
335 using migration velocities is also carried out. The description of sections will be done from N to
336 S and from E to W. The CIMDEF dataset will be only described in the discussion (Figs. 9 and 10)
337 as it does not include NI data but is key to understanding the geometry of the mid-crustal
338 discontinuity, its late Variscan reworking and its Alpine reactivation.

339 3.3.1. Cantabrian Zone (ESCIN-1 section)

340 The ESCIN-1 section is a ~130 km long, E-W profile crossing the CZ from its most external part
341 to the Narcea Antiform to the W, in the boundary with the hinterland (WALZ, Figs. 1 and 3). ~~It~~
342 ~~consists of two slightly overlapping parts, 1.1 and 1.2, separated a few kilometers in the N-S~~
343 ~~direction.~~ The complete ESCIN-1 section, migrated at $v=5600$ m/s (Fig. 2) and its interpretation
344 are presented in figure 3.

345 This section-profile was first described and interpreted over an unmigrated image by Pérez-
346 Estaún et al. (1994). Later works revisited the interpretation, adding travel-time modeling to
347 help on the understanding of the unmigrated data (Gallastegui et al., 1997). The reader is
348 referred to these papers for further details than those provided here.

349 In the upper crust, the western part shows W-dipping reflections that represent the Variscan
350 imbrication, through a thrust ramp (t in Fig. 3), of the basement under the Paleozoic
351 succession, indicating the proximity of the hinterland (WALZ). In fact, a Neoproterozoic, non-
352 metamorphic sequence outcrops in this area, which is probably underlain by an older
353 crystalline basement. Another prominent W-dipping reflection roughly parallel to t (t')
354 crosscuts subhorizontal ones, defining a pattern that might be providing an out of the plane
355 image of the above mentioned thrust ramp, as this profile lies in the hinge of an arcuate
356 structure, the IAA (Fig. 1b). To the E, the thin skinned tectonics characteristic of this external
357 zone can be interpreted from shallow subhorizontal to W dipping reflections often coincident
358 with outcropping thrusts (ot), as observed in figure 3a. The main one among these, running at
359 around 5 s (TWT), is interpreted as the sole thrust of the thin-skinned orogenic wedge (st). To
360 the W, it gets involved in the crustal ramp (t in Fig. 3) observed at the Narcea Antiform,
361 suggesting that it ends down rooting into the upper part of the lower crust (c). A low
362 reflectivity wedge of undifferentiated basement (b) located between 4-5 and 8.5 s (TWT) exists
363 underneath the easternmost reflections. This may image some pre-Paleozoic basement that is
364 interpreted as upper crust, since the pattern of reflections changes below, suggesting that a
365 significant boundary occurs underneath.

366 The lower crust shows little reflectivity but seems to be present in the interval between 8.5-14
367 s (TWT) in the E and between 8.5 and 12 s (TWT) in the W. It features subhorizontal (ir) and W
368 dipping internal reflectivity to the E, the latter (ir') crosscutting the former reflections. These
369 might represent the imprint of Alpine tectonics over a previously deformed/reflective lower
370 crust. To the W, reflectivity seems to be subhorizontal or dipping to the E (ir). Some of the
371 dipping and arcuated reflectivity observed at the edges of section ESCIN-1 (mig in Fig.3 and

372 thereafter) might be related to the migration effects over discontinuous features and caution
373 should be taken when interpreting it.

374 The Moho along this section (m) is located at nearly 14.5 s TWT (~45 km) in the eastern part,
375 and shallower (12 s TWT, ~36 km) to the W. The crustal thickening observed to the E is
376 probably related to an out of section image of the crustal Alpine root, better observed in
377 profile ESCIN-2, which is described next.

378 **3.3.2. Cantabrian Zone and Duero Basin (ESCIN-2 section)**

379 The ESCIN-2 seismic line is a 65 km long, N-S section that samples the transition between the
380 CZ and the DB (Fig. 1). Even though this profile was geared to study the Alpine structures, it
381 shows how the Variscan features have been ~~inherited and~~ reactivated during the Cenozoic
382 compression between the Iberian Peninsula and the European plate. The section was first
383 presented by Pulgar et al. (1996). Later on, some authors have used this image to constraint
384 the Alpine structure in the North Iberian Margin (e.g., Fernández-Viejo et al., 2000; Gallastegui
385 et al., 2016). However, only Teixell et al. (2018) used a migrated version (4000 m/s) of this
386 section. Here we present the results of a Kirchhoff time migration at $v=5600$ m/s (Fig. 4).

387 This seismic line shows, in places, a conspicuous reflectivity that allows a straightforward
388 interpretation. To the S end, the upper crust is characterized by high amplitude horizontal
389 reflectivity representing the DB sedimentary sequence (Fig. 4a). It occupies the interval from 0-
390 3.5 s TWT (s in Fig. 4b, c and d) and appears to be offset by N dipping reflections (t). The latter
391 have been interpreted as S vergent Alpine thrusts affecting the CZ basement and partly the DB
392 sediments. The rest of the crust is less reflective although N dipping reflectivity (ot), also
393 interpreted as ~~imaging~~ Alpine thrusts on the basis of the clearer stack image (Pulgar et al.,
394 1996), crosscuts shallow subhorizontal weak reflections that represent the Paleozoic
395 sedimentary sequence of the CZ (ps).

396 The lower crust presents higher amplitude reflectivity. In general, a thick band of horizontal
397 reflections located between 7.5 and 12 s (TWT) at the southern part of the profile, bends and
398 dips to the N in the northern part of the line (lc) in response to Alpine compression. Although
399 the stacked section shows that this N dipping reflectivity reaches 14.5 s TWT (Pulgar et al.,
400 1996) the migrated sections (Teixell et al., 2018 and Fig. 4) indicate that these reflections move
401 southward and upward to less than 14 s (TWT), while losing amplitude and coherence. In fact,
402 the geometry of the bottom of the lowermost crust (Moho, m) is deduced on the basis of the
403 geometry of its uppermost part (c), the lower crust internal reflectivity, the stack image (Pulgar
404 et al., 1996), and the amplitude contrasts observed in the attribute analysis (Fig. 4).
405 Furthermore, its depth is solely established on the basis of the position of the strongest
406 subhorizontal reflections to the S.

407 Even though this profile shows the imprint of recent Alpine shortening, no reflections are
408 observed to crosscut the entire crust. In contrast, reflectivity suggests that deformation is
409 decoupled between the upper and lower crust. However, this section is not long and/or
410 reflective enough as to image where the Alpine thrusts (ot) root. Possibly, they merge into the
411 roof of the underthrust CZ lower crust.

412 A 1D Vp profile (Fig. 4e) extracted from the coincident WA model (Pulgar et al., 1996) shows a
413 conspicuous velocity increase in the lower crust, at a depth roughly coincident with (c). Depth
414 misfits are due to the effect of the low Vp of the DB sediments not being taken into account in
415 the depth conversion.

416 3.3.3. West Asturian-Leonese Zone (ESCIN-3.3 section)

417 The ESCIN-3.3 profile is part of a ~375 km long crooked offshore seismic line consisting in
418 ESCIN-3.1, 3.2 and 3.3. The latter is 137 km long, parallel to the coast and close to it across the
419 WALZ (Fig. 1). It was first presented by Martínez Catalán et al. (1995) and Ayarza et al. (1998,
420 2004). Later on, its image has been used to constrain the structure of the western North
421 Iberian Margin and that of the transition between the WALZ and the CIZ (Martínez Catalán et
422 al., 2012, 2014). The cross-section ~~presented~~ in figure 5a ~~corresponds~~ represents to the
423 equivalent onshore transect of this profile.

424 Reflectivity in the upper crust is characterized by the image of Mesozoic sedimentary basins (s
425 in Fig. 5) related to the extension that led to the opening of the Bay of Biscay. Underneath, W
426 dipping reflections (t) are interpreted as the imprint of the first stages of Variscan
427 compressional deformation in the WALZ (C1 and C2), developing E-vergent thrust faults (Fig.
428 5a). These affect the pre-Paleozoic basement and root in the upper part (c) of a thick reflective
429 band interpreted as the lower crust (lc) or in a sole thrust (st) that also reaches the lower crust.

430 The lower crust (lc) is represented by a thick band of subhorizontal reflectivity (8-12 s TWT)
431 that thickens (6-12 s TWT) in the westernmost part of the WALZ (CDP 3000) underneath the
432 Lugo Dome, an extensional structure bounded to the E by the Viveiro normal fault (Fig. 5).
433 Then it thins towards the end of the line, when entering the CIZ, coinciding with an area of E
434 dipping sub-crustal reflections (sc) altogether defining a less clear Moho (m) in this area. The
435 ESCIN-3.3 lower crust seems to feature an internal layer with mantle P-wave velocities when
436 modeled from coincident WA data. Accordingly it was interpreted as consisting of the WALZ
437 lower crust underthrust by the CZ lower crust (Ayarza et al., 1998). This model would
438 compensate the high shortening observed in the upper crust of the CZ, a thin-skinned belt
439 whose sole thrust roots at the contact with the WALZ. The internal reflectivity of the lower
440 crust shows W dipping reflectors (ir), similar to the ones observed in the upper crust and
441 probably imaging Variscan deformation in the lower crust, either compressional or
442 extensional. They crosscut subhorizontal reflectivity, thus postdating it. A 1D P-wave velocity
443 profile (Fig. 5e) derived from coincident WA data (Ayarza et al., 1998) shows again an
444 important increase in relation to the top of the lower crust (c).

445 Even though migration ($v= 5200$ m/s, Fig. 2) over discontinuous reflections blurs the seismic
446 section in the edges (mig in Fig. 5), reflectivity never seems to cross-cut the crust and/or the
447 Moho, indicating that deformation is decoupled at upper and lower crustal level. Subcrustal E
448 dipping reflections (sc) are interpreted as the out-of-the-plane image of the Alpine southward
449 subduction of the Bay of Biscay oceanic crust (Ayarza et al., 1998, 2004), which is out of the
450 scope of this paper.

451 The boundary between the WALZ and the CIZ is the Viveiro Fault, one of the most striking
452 surface expressions of Late Variscan extensional tectonics, featuring a decompression of ~4kb

453 or 14 km (Reche et al., 1998). To its W, gravitational collapse of a thickened crust and
454 associated crustal extension and melting have played a key role in the orogenic evolution of
455 the CIZ. However, to the E, crustal re-equilibration after C1 and C2 thickening was less
456 important and igneous activity decreases. Even though this fault itself is not identified in the
457 seismic section (Fig. 4), the reflectivity in general varies ~~on to~~ both sides of it, featuring a
458 thinner (9 s TWT vs 12 s TWT) and more transparent crust to the W (Fig. 6). In fact the
459 geometry of some reflections (ef) in the boundary between the WALZ and the CIZ, above the
460 thickest lower crust, and the subtractive way the sole thrust (st) merges with the lower crust
461 (st') seem to indicate the effect of extensional tectonics, sometimes reactivating
462 compressional structures (st-st'). Such a reactivation has been described ~~on~~ the basis of
463 structural and metamorphic considerations of the the main WALZ thrust sheet ~~in the WALZ~~
464 ~~based on structural and metamorphic considerations~~ (Alcock et al., 2009). Conversely, further
465 to the E ~~of the section~~, reflectivity probably represents the geometry of preserved
466 compressional deformation.

467 **3.3.4. Northern Central Iberian Zone (ESCIN-3.2 section)**

468 The seismic line ESCIN-3.2 is a 97 km long profile, also parallel and close to the coast, that
469 samples the relative autochthon to the GTMZ, i.e. the CIZ (Figs. 1 and 6). It was first described
470 by Álvarez-Marrón et al. (1996) and later by Ayarza et al. (2004). Here it is presented migrated
471 ~~with a~~ $v=5200$ m/s (Fig. 2). The cross-section ~~presented~~ in figure 6a corresponds to the
472 equivalent onshore transect of this profile and depicts allochthonous sequences not imaged by
473 the NI data.

474 This profile shows, in the upper part, a band of high subhorizontal reflectivity that coincides
475 with the location of Mesozoic basins, as in profile ESCIN-3.3 (s in Fig. 6). The rest of the upper
476 crust is not very reflective although a couple of W-dipping reflections (ef) rooting in a thin
477 band of strong reflectivity are observed. These reflections, located in the E of the section from
478 4.5 s to 8 s TWT, define a sort of duplex, extensional or compressional, but later extended,
479 indicating in any case boudinage and crustal thinning. To the W, the upper crust is very
480 transparent and just a few weak reflections can be observed.

481 The narrow reflective band at 8-9 s TWT represents the lower crust (lc), and is the most
482 striking feature of this profile. Its 1 s TWT thickness contrasts with that observed in the
483 neighboring ESCIN-3.3 and ESCIN-1 sections, which show a much thicker lower crust (4-5 s
484 TWT). Reflectivity in this band is subhorizontal, although somewhat undulated, while the band
485 itself is slightly inclined to the W. In the E, the Moho (m) is located above 9 s TWT, the
486 shallowest identified so far in the Iberian Massif. Subcrustal E dipping reflections (sc) are again
487 associated to the 3D image of the southward subduction of the oceanic crust of the Bay of
488 Biscay during the Alpine convergence (Ayarza et al., 2004) whereas W dipping features might
489 be related with the CZ lower crust (czlc) underthrust also underneath the easternmost part
490 of the CIZ.

491 This profile samples the northern CIZ, where Variscan crustal thickening during C1 and C2, was
492 most important. Consequently, later gravitational collapse triggered extensional tectonics and
493 crustal melting, allowing the intrusion of granites and the development of extensional
494 detachments (with associated metamorphic offsets). The image of line ESCIN-3.2 shows a

495 transparent upper crust to the W suggesting that granites occupy most of it, which is
496 supported by onland geological mapping. Some thrust faults, as those imaged by W-dipping
497 reflections in profile ESCIN-3.3, probably root along this section and are represented by the W-
498 dipping reflections at the base of the upper crust. However, these were later flattened and/or
499 reactivated as extensional detachments by crustal thinning (ef). The narrowness of the highly
500 reflective lower crust here suggests that crustal thinning was largely accommodated at this
501 level, as the upper crust has basically the same thickness as in the ESCIN-3.3 line (up to 6-7 s
502 TWT). In addition, crustal melting might have also affected the top of the lower crust. But even
503 though large parts of the crust were melted, reflectivity exists deep in the upper crust,
504 suggesting that crustal melting was not pervasive and/or reflectivity is linked to syn- or late-
505 tectonic features.

506 **3.3.5. Southern Central Iberian Zone (ALCUDIA section)**

507 The ALCUDIA seismic profile was first presented by Martínez Poyatos et al. (2012) and
508 reprocessed and further interpreted by Ehsan et al. (2014). It is a more than 220 km long, NE-
509 SW seismic profile sampling the CIZ to the S of the ICS down to the boundary with the OMZ
510 (Fig. 1). Here we presented this section migrated and depth converted using a $v=6200$ m/s
511 (Figs. 2 and 7).

512 This profile presents a fairly transparent upper crust when compared to other nearby sections
513 (e.g., IBERSEIS, Fig. 8) although scarce reflectivity exists to the S coinciding with the boundary
514 (suture) between the CIZ and the OMZ, namely, the Central Unit (CU; cu in Fig. 7) and to the
515 presence of vertical folds (vf). Some very transparent zones (g) appear to be in relation with
516 the existence of granitic batholiths. To the N, the intrusion of these granites, associated to the
517 existence of normal faults (ef), is one of the evidences of extensional tectonics affecting the
518 southern part of the CIZ. The rest of the upper crust shows weak and discontinuous reflectivity
519 that responds to the existence of vertical folding affecting lithologies with little impedance
520 contrast. In fact, deformation in the upper crust of this part of the CIZ is weak, with absence of
521 low-dipping structures typical of tangential tectonics.

522 The lower crust shows a very different image to that of the upper crust. It is a thick band, of up
523 to 6 s TWT (from 4 s to 10 s), of mostly subhorizontal high amplitude reflectivity (lc) that at
524 some points appears to be cut across by N-dipping reflectors (cc). S-dipping internal reflectivity
525 is also identified although more scarce (ir). The lower crust thins in the northern end of the
526 profile, near the ICS, where intrusion of granites and other evidences of crustal re-equilibration
527 suggest that extension played a key role. Accordingly, we suggest that the mechanisms that
528 triggered this lower crustal thinning are related to melting and extension and not with
529 compression, as previously proposed (Ehsan et al., 2014; Martínez Poyatos et al., 2012), and
530 that the N dipping reflectivity observed in the top of the lower crust (c) in that area is the
531 expression of extensional tectonics.

532 One of the most striking features of this profile is the crocodile-like structure affecting the
533 lower crust at around CMP 10000 (cc and cc'). This structure, most likely related to Variscan
534 shortening, accommodates an important part of the deformation at lower crustal level and
535 evidences that sub-horizontal reflectivity of the lower crust is pre-Variscan, thus raising the
536 question about its precise age and origin. Despite the presence of this feature in the depth

537 continuation of the suture between the CIZ and the OMZ (Fig. 7a), reflectivity does not
538 crosscut the whole crust, suggesting the existence of a detachment in the top of the lower
539 crust. This contrasts with the presence in the upper crust (CU) of retro-eclogites with peak
540 metamorphic conditions of 19 kbar and ~550°C (López Sánchez-Vizcaíno et al., 2003). Finally,
541 the Moho boundary (m) is located at a fairly constant depth (~10 s TWT, i.e. 30-33 km),
542 although the lower crust seems to be preserved and a local crustal imbrication into the mantle
543 is observed underneath the crocodile-like structure. A 1D Vp profile (inset in Fig. 7d) from a
544 coincident WA-data model (Palomeras et al, in press) shows a conspicuous increase of values
545 along more than 15 km starting in the top of the lower crust (c) thus supporting its important
546 thickness.

547 **3.3.6. Central Iberia, Ossa-Morena and South Portuguese Zones (IBERSEIS section)**

548 The IBERSEIS seismic line was first presented by Simancas et al. (2003). A number of later
549 works added information and details to its interpretation (e.g., Carbonell et al., 2004;
550 Schmelzbach et al., 2007, 2008; Simancas et al., 2006). This section crosses the southernmost
551 CIZ, the whole OMZ, and most of the external SPZ (Fig. 1). It samples two major boundaries
552 interpreted as suture zones: that between the CIZ and the OMZ (CU, Azor et al., 1994) and the
553 one bounding the OMZ and the SPZ, which has been largely affected by younger Carboniferous
554 events (Pérez-Cáceres et al., 2015). The IBERSEIS profile structurally overlaps the ALCUDIA
555 profile along ~30 km, but it is displaced some 50 km to the W. A cross-section along this
556 transect together with its interpretation after migration at v=6000 m/s (Fig. 2) are shown in
557 figure 8.

558 This section is ~300 km long and features an outstanding reflectivity at upper and lower crustal
559 levels. In the upper crust, a wealth of N dipping reflections (t in Fig. 8) image a S verging thrust
560 and fold belt. In the SPZ, these are most reflective and probably related to normal faults
561 derived from the extension that led to the opening of the Rheic Ocean and later reactivated as
562 thrusts during the Late Carboniferous compression. Some authors link the highest reflective
563 features to the middle Carboniferous volcano-sedimentary complex (Schmelzbach et al., 2008),
564 which might have used these fractures as a conduit, thus enhancing reflectivity. In the OMZ, N
565 dipping reflections probably image Variscan thrust faults (ot) as some coincide with such
566 mapped structures. Their lesser reflectivity might indicate the lack of involvement in the
567 thrusts of lithologies that increase the impedance contrast.

568 Upper crustal reflectivity in both, the ZOM and SPZ, does not cross to the lower crust, rooting
569 at a mid-crustal level that, in the SPZ is transparent and does not have any particular
570 expression itself but does coincide with the top of the lower crust (c). However, in the OMZ, a
571 reflective layer exists at this depth (irb): it has been defined as the IBERSEIS reflective body
572 (IRB, Simancas et al., 2003), a 140 km long, high velocity conductive feature (Palomeras et al.,
573 2009) that is supposed to represent an early Carboniferous mantle-derived intrusion. Its origin
574 has been related to mantle plume activity that thinned the lithosphere and extracted mantle-
575 derived melts from the ascending asthenosphere (Carbonell et al., 2004). Its surface expression
576 are intraorogenic transtensional features (Rubio Pascual et al., 2013; Simancas et al., 2006).
577 Alternatively, Pin et al. (2008) have suggested, based on geochemical constraints, a tectonic
578 scenario of slab break-off for this feature. Internal reflectivity along the IRB is mostly

579 subhorizontal, probably due to the effect of the intrusion along a subhorizontal detachment,
580 and evidences little imprint from Variscan deformation. The body is slightly inclined to the S, at
581 odds with the detachment being the sole thrust of the OMZ upper crustal imbricates. Perhaps
582 it was, but its inclination changed during subsequent deformation, as later suggested.

583 The lower crust shows slightly different patterns in the CIZ and OMZ on one side and the SPZ
584 on the other. In the southernmost part of the CIZ and northern OMZ, N and S dipping
585 reflections define a wedge (cc) that might be the western continuation of the crocodile-like
586 structure observed in the ALCUDIA seismic line in an equivalent structural position (cc in Fig.
587 7). In this section, the limited crustal imbrication into the mantle identified in the ALCUDIA line
588 (cc' in Fig. 7) is not observed, perhaps because it only occurs further to the N or E. This
589 structure may be the reason why the IRB is shallower at this point, indicating that the latter is
590 older than the crocodile compressional feature. The rest of the lower crust shows S dipping
591 and sub-horizontal (lc) reflectivity that does not exhibit clear crosscutting relationships, thus
592 hindering their interpretation. However, near the boundary with the SPZ, this reflectivity
593 seems to be affected by N dipping features (ir) overprinting them. In the SPZ, the lower crust
594 shows a more homogeneous image, with subhorizontal reflectivity (lc) that is often cut across
595 by longer scale S dipping features (lct) that postdate them. The latter probably represent
596 fractures that firstly accommodated the extension linked to the opening of the Rheic Ocean
597 and were then reactivated as thrusts during the late Carboniferous compression and collision
598 of the SPZ basement with the OMZ. The most conspicuous of these reflections (lct') cuts the
599 IRB in its southern part and seems to offset the lower crustal upper boundary between the SPZ
600 and the OMZ. Two 1D Vp profiles derived from coincident WA-data (Palomeras et al., 2009)
601 and shown as insets in figure 8d indicate a velocity increase starting at the top of lower crust
602 (c) and along the IRB.

603 Even though the lower crust in the OMZ and SPZ shows dipping features, none of them crosses
604 to the upper crust, thus rooting at a mid-crustal level as does the upper crustal reflectivity. This
605 implies again the existence of a discontinuity (c) in the mid-crust.

606 Despite of crossing two suture zones and imaging part of a crocodile-like structure, the
607 IBERSEIS profile shows a fairly flat Moho (m) located at ~10 s TWT, the same apparent depth as
608 in the ALCUDIA line (30-33 km). Its signature is very clear underneath the SPZ and a bit blurry
609 below the IRB.

610 **4. Discussion**

611 Simancas et al. (2013) already undertook an integrated interpretation of most of the seismic
612 sections presented here focusing on, i) the accommodation of orogenic shortening at crustal
613 scale, (ii) the relationships between convergence, crustal thickening and collisional granitic
614 magmatism, and (iii) the development of the Iberian Variscan oroclinal. In this paper the same
615 sections are presented, but they have been reprocessed at stack level and time migrated using
616 a Kirchhoff algorithm. In addition, two extra sections that image the alleged mid-crustal
617 discontinuity after the Alpine reactivation are taken into account. The first one is the N-S
618 ESCIN-2 NI dataset (Fig. 4), in the CZ, where this discontinuity has remained untouched during
619 late Variscan evolution but was reactivated during the Alpine Orogeny. The second one results
620 from the CIMDEF experiment, carried out in the CIZ across the ICS, where the mid-crustal

621 discontinuity has probably been affected by crustal melting during the Late-Variscan extension
622 | and by later Alpine reactivation. The latter sections ~~somehow~~ fill ~~the~~ a key gap existing in
623 | Simancas et al. (2013).

624 Figures 3 to 8 represent an effort to show a homogeneous seismic image of the Iberian Massif
625 | crust that eased its integrated interpretation. Next, we discuss the main observed features,
626 | ~~their implications~~ and how they contribute to the understanding of the structure and evolution
627 | of the Iberian Massif, adding constraints to the origin of the elevation of the central Iberian
628 | Peninsula. Figure 9 presents a simplified sketch of the crustal layers observed in the Iberian
629 | Massif. Figure 10 shows a compendium of the position of the mid-crustal discontinuity and the
630 | Moho depth (in TWT) along the entire Iberian Massif as deduced from seismic NI data together
631 | with a map of the ~~entire~~ Iberian Peninsula Moho depth (Palomeras et al., 2017) that includes
632 | the position of the seismic profiles for comparison. We will refer to these figures along most of
633 | the discussion.

634 | A particular feature of the SW Iberian Massif is the ~~great importance of out-of-section,~~
635 | ~~mainly~~ relevance of left-lateral shear zones associated to its suture boundaries. They displaced
636 | central and northern Iberia to the NW with respect to southern Iberia (Simancas et al., 2013).
637 | The seismic sections do not provide constraints about this movement, as it is perpendicular to
638 | their layout. Thus, interpretations in these areas must be taken with caution.

639 | **4.1. The upper crust in the Iberian Massif: a depth image of outcropping geology**

640 | Most of the seismic sections display a moderate to thick upper crust (4 to 8 s TWT, Fig. 9), with
641 | very variable reflectivity. Reflections coincide with outcropping Variscan structures and thus, a
642 | link has been established. As such, N dipping reflectivity in the SPZ and the OMZ is related to S
643 | | vergent folds and thrust faults mapped ~~in-at~~ the surface. W dipping reflections in the CZ are
644 | related to mapped thin-skinned thrusts. The same type of reflectivity observed in the WALZ,
645 | albeit reaching deeper levels ~~in the crust~~ and rooting in the lower crust, has been addressed as
646 | evidence of thick-skinned thrust tectonics, which in the hinterland affects the pre-Paleozoic
647 | basement. Particularly interesting is the upper crustal SPZ seismic image in contrast with that
648 | of the CZ, both representing external zones. While in the latter thrusts are observed to root in
649 | a shallow sole detachment, in the former one reflections/thrusts root in the lower crust. This
650 | feature will be discussed in the next section.

651 | Only a few seismic profiles feature a transparent upper crust. Lack of reflectivity has been
652 | addressed to low fold data (ESCIN-1 and ESCIN-2, Figs 3 and 4), and most importantly to the
653 | | existence of a re-equilibrated upper crust ~~having recorded~~ featuring large amounts of partial
654 | melting, as shown by voluminous outcropping granitoids (ESCIN-3.2 and N of ALCUDIA, Figs. 6
655 | and 7). The existence of vertical folds affecting little reflective monotonous lithologies also
656 | results in a fairly transparent upper crust in most of the ALCUDIA section (Fig. 7).

657 | None of the upper crustal reflections observed ~~and interpreted~~ in the presented Iberian Massif
658 | | NI seismic sections ~~seems to~~ cuts across the whole crust, always rooting in a sole thrust (parts
659 | of ESCIN-1 and ESCIN-3.3, Figs. 3 and 5) or in the lower crust (the rest of them).

660 **4.2. The lower crust in the Iberian Massif: accommodation of shortening, extension and its**
661 **nature**

662 The Iberian Massif dataset presented here shows a very coherent image of the lower crust. Its
663 reflectivity is high and usually subhorizontal. However, cross cutting relationships with later
664 features of opposite dips evidence a multi-phase origin for this reflectivity.

665 The SPZ, OMZ, WALZ, CZ and the southern CIZ show that this part of the crust is also thick (4 to
666 6 s TWT). However, ~~in NW Iberia~~ the ESCIN-3.2 and the northern part of the ALCUDIA sections
667 (Figs. 1, 6 and 7); ~~the few existing NI profiles~~ indicate that, ~~the~~ in the northern CIZ, the lower
668 crust is much thinner (1 to 2 s TWT) and irregular (ESCIN-3.2, Figs. 6 and 9). This thin lower
669 crust has been observed in ~~the areas~~ characterized by outcropping syn-collisional granitoids
670 (zone II of Simancas et al., 2013). These witness ~~the onset of~~ crustal re-equilibration processes
671 triggered by gravitational collapse, extension and crustal melting during the Late
672 Carboniferous. The straightforward conclusion is to attribute ~~the architecture of this thin~~ lower
673 crust to late Variscan orogenic extension, which features at the surface high metamorphic
674 offsets (chlorite to sillimanite zone, Díez Balda et al., 1995) and melting, implying that crustal
675 thinning has been ~~mostly accommodated by its lowermost part at upper and lower crustal~~
676 levels.

677 **4.2.1. Effects and outreach of late Variscan extension in the CIZ lower crust: CIMDEF,**
678 **ALCUDIA and ESCIN-3.2 sections**

Con formato: Fuente: Negrita

679 ~~Nevertheless~~ The CIZ is a conflicting area. Although, a NI data gap ~~of crustal scale NI data~~ exists
680 in most of the northern CIZ. ~~The~~ CIMDEF ~~noise~~ autocorrelation profiles (Figs. 1 and 9 and ~~),~~
681 based on passive source seismic data, show a thick (~5 s TWT) lower crust in most of this area,
682 ~~which essentially corresponds to the CIZ~~ (Andrés et al., 2019, 2020). This is in conflict with the
683 NI sections ESCIN-3.2 and northernmost ALCUDIA, where the highly reflective lower crust is
684 less than half as thick (1-2 s TWT). However, granitoids are probably scarce in the Variscan
685 basement sampled by CIMDEF and hidden under the DB (Fig. 1), ~~which can then present~~
686 explaining the prevalence of a thick lower crust. But in and near the ICS batholith, a rather
687 continuous internal reflection in the lower crust could be interpreted as its top part (Figs. 9
688 and 10), thus indicating that crustal thinning and melting, observed in the surface, has also
689 affected the lower crust (Andrés et al., 2020).

690 Extension in the northern CIZ occurred simultaneously with shortening in the SW Iberian
691 Massif. According to Simancas et al. (2013) this suggests that the tectonic stresses would be
692 dominantly compressional, still induced by ongoing collision. In fact, gravitational instabilities
693 in a thickened crust should mostly be affecting the upper crust. In this context, theoretical
694 models (Royden, 1996; Seyferth and Henk, 2004) indicate that beneath ~~the~~ areas of extension
695 in the upper crust, shortening may prevail in the lower crust. This mechanism is an efficient
696 way for syn-convergent exhumation of deep rocks.

697 Indeed, from a regional tectonic perspective, compression was active till the end of the
698 Variscan orogeny, and at times, clearly simultaneous with extension (C3 and E2 overlapped in
699 the interval 315-305 Ma; Martínez Catalán et al., 2014). But it is clear that extension affected
700 the lower crust, as it ~~appears is~~ thinned in areas of transparent, extended molten upper crust

701 | (ESCIN-3.2 and ALCUDIA sections, Figs. 6 and 7). However, the irregular ~~pattern observed in~~
702 | ~~the geometry of the~~ ESCIN-3.2 lower crust might indicate the existence of folds in this re-
703 | equilibrated layer, witnessing the simultaneity of extension and compression ~~even~~ at lower
704 | crustal level (Fig. 6). ~~In addition~~ However, we cannot rule out that these undulations represent
705 | boudinage (i.e. extension) or Alpine folding, although we consider the latter less likely.

706 | In the ALCUDIA section, the imaged part of the CIZ underwent only moderate upper crustal
707 | shortening (Martínez Poyatos et al., 2012). ~~According to Simancas et al. (2013)~~ However, it
708 | ~~exhibits a the~~ thick laminated lower crust, representing pre-Pennsylvanian (most probably pre-
709 | Variscan) ductile deformation (Simancas et al., 2013). ~~In addition, the latter,~~ appears deformed
710 | in two sectors near both ends of the profile, concentrating ~~shortening deformation~~ in discrete
711 | structures ~~that compensate the upper crustal deformation~~. The first of them is the very
712 | conspicuous crocodile-like structure observed in the southern end, and also imaged in the
713 | northern part of the IBERSEIS line (Fig. 9b). This structure mimics localized crustal indentation
714 | of the OMZ into the CIZ, producing a local underthrusting of the latter to the S that is still
715 | (partly?) preserved. Indentation generated tectonic inversion of the Los Pedroches early
716 | Carboniferous basin (Simancas et al., 2013) and bending of the overlying upper crust, as seen
717 | by the uplift of the IRB, both of which predate the imbrication. The Los Pedroches batholith
718 | intruded above at 314-304 Ma in an extensional setting (Carracedo et al., 2009), postdating
719 | the age of the wedge as no further deformation affected the batholith. ~~Indeed, t~~ This
720 | crocodile-like feature ~~must probably~~ represents early Carboniferous Variscan compressional
721 | deformation and must account for part of the shortening observed at upper crustal level.

722 | ~~However~~ Contrarily, to the NE of this section, a ramp-and-flat geometry has been interpreted
723 | as a major lower crustal thrust (Martínez Poyatos et al., 2012; Simancas et al., 2013) that
724 | helps to compensate upper and lower crustal shortening. However, the highly reflective
725 | lower crust is not repeated in the hanging wall to the structure, so that a subtractive character
726 | is a reasonable alternative. As stated above, the thin lower crust to the N of the ramp seems to
727 | ~~be clear~~ evidence of lower crustal thinning (Fig. 9b), ~~supported by the fact that~~ as it underlies
728 | an area of upper crustal extension, the Toledo gneiss dome, characterized by normal faulting
729 | and pervasive partial melting (Barbero, 1995; Hernández Enrile, 1991). Regardless of how
730 | much shortening that area accommodated during crustal thickening and even though the
731 | observed ramp could be a ~~former~~ thrust fault reactivated ~~during by~~ later extension, the
732 | present image of the lower crust does not suggest compensation of upper crustal shortening.
733 | ~~In fact,~~ But the lower crust in the ALCUDIA section is anomalously thick elsewhere (up to 6 s
734 | TWT, 18 km) suggesting the possibility of important ductile thickening previous to the
735 | extension that triggered thinning at its northern part.

736 | 4.2.2. Lower crust signature in SW Iberia: The IBERSEIS and southern ALCUDIA sections

Con formato: Fuente: Negrita

737 | S of the CIZ, ~~In the IBERSEIS profile,~~ lower crustal ~~dominant~~ reflectivity is also subhorizontal but
738 | disrupted by N and S dipping features (Fig. 9b). Whereas in the OMZ these features usually dip
739 | to the N, as do the upper crustal reflections representing Variscan thrusts, in the SPZ they
740 | surprisingly mirror the S dipping upper crustal Variscan thrusts, ~~dipping to the S~~. Furthermore,
741 | one of these features, placed close to the boundary with the OMZ, affects almost the entire
742 | lower crust.

743 Orogenic orthogonal shortening in the ~~OMZ~~ upper crust has been estimated in 120 km (~57%)
744 for the OMZ and ~~in the SPZ around~~ ~80 km (~45%) or even less for the SPZ (Pérez-Cáceres et
745 al., 2016). According to Simancas et al. (2013), the crocodile structure and a ~~not-observed~~
746 ~~associated~~ hypothetical northward subduction of the OMZ might account for this shortening in
747 the OMZ. Similarly, in the SPZ, the lower crustal imbricated structures represent only ~ 20 km
748 of shortening so that ~~according to these authors,~~ a detached lower crustal subduction along the
749 OMZ/SPZ might have accommodated the other 60 km.

750 In this regard, we suggest that the present day SPZ crustal image represents its decoupled
751 early Carboniferous extension and later compression. This evolution would have erased any
752 evidences of previous (pre-Carboniferous) subduction, and forced the SPZ to thin during
753 extension. i.e., the lower crust had to decrease its thickness ductilely, perhaps first in a more
754 or less distributed way and later through localized shear zones (brittle or not depending on the
755 depth) as it became shallower. However, the upper crust could have preserved most of its
756 original thickness, as the developing basins associated to extension would have been
757 constantly fed by sediments and igneous extrusions and intrusions (like the IRB in the OMZ).
758 Later compression would have folded and thrustured the upper crust, and also thickened the
759 lower crust. A few lower crustal normal shear zones might have developed during extension
760 and then be reactivated as ductile thrusts during compression. Those are today observed as S
761 dipping reflections that disrupt the previous subhorizontal ~~previous~~ reflectivity in the lower
762 crust and mirror ~~thrusts in~~ the upper crustal thrusts. Accordingly, distributed ductile
763 deformation and thrusting might have thickened the lower crust back to its original (or simply
764 stable) thickness in the SPZ ~~and elsewhere~~, something that cannot be measured-quantified but
765 would need to be accounted for when comparing shortening at upper and lower crustal levels.
766 The resulting seismic image of the SPZ would then be that of an extended and then inverted
767 margin, with mirroring reflectors in the upper and lower crust merging in a mid-crustal
768 discontinuity and providing a seismic image different to that of a typical foreland thrust and
769 fold belt (e.g., CZ; Fig. 3). This evolution differs from that of a hyperextended magma-rich
770 margin as stretching of the upper and lower crust is not coupled and faults do not cut across
771 the crust and penetrate down into the mantle. In any case, the S dipping lower crustal
772 reflections, active during the Late Carboniferous, postdate ~~the sub-~~horizontal reflectivity of at
773 the lower crust. It is worth mentioning here that the SPZ seismic image is identical to that of
774 the Rhenohercynian Massif in Germany (Franke et al., 1990; Oncken, 1998) suggesting a similar
775 evolution.

776 The discussion above shows that the lower crust in the Iberian Massif is thick, except when it is
777 affected by late orogenic extension. The mechanisms that produced lower crustal thickening
778 are probably related to compressional deformation, mostly ductile. Continental underthrusting
779 of the CZ underneath the WALZ (Ayarza et al., 1998, 2004) and part of the CIZ (Fig. 6, 9 and 10),
780 indentation of the OMZ in the CIZ (Figs. 9 and 10) and Variscan thrust-like structures probably
781 played an important role. In addition, the latter help to constrain the age of the subhorizontal
782 reflectivity. Frequent disruption of subhorizontal lower crustal lamination by Variscan (late
783 Carboniferous) dipping features indicates that the lamination developed prior to Variscan
784 compressional deformation. What this lamination represents is still an open question.

785 4.2.3. Origin of the lower crust lamination

Con formato: Fuente: Negrita

786 Many vertical incidence seismic reflection profiles worldwide have shown reflective ~~and~~
787 ~~laminated~~ lower crusts (e.g., Meissner et al., 2006; Wever, 1989) ~~that~~ ~~Lower crust seismic~~
788 ~~lamination~~ ~~has~~ ~~often~~ been ~~often~~ related to late orogenic extensional events (Meissner,
789 1989). In the Iberian Massif, surface geology shows that late orogenic extension affects the
790 upper crust, mainly in areas of large previous thickening. But in contrast to the latter author's
791 models, important thinning of the lower crust takes place in those areas (ESCIN-3.2 and
792 northern ALCUDIA, Figs. 6, 7 and 9). Certainly, lower crustal lamination might come from
793 underplating eased by extension in magma rich margins (Klemperer et al., 1986). But also,
794 ductile deformation is a very likely source of lower crustal lamination. Dipping events observed
795 in the lower crust crosscutting a strong banded reflectivity represent the latest orogeny-
796 related shortening, which will be further flattened and horizontalized in the next orogeny.
797 Continuous superposition of deformational events at lower crustal level ~~managed to~~
798 decrease the dip of structural/lithological markers and define a subhorizontal fabric. These
799 deformation mechanisms can generate structures with a strong ~~ly-defined~~ anisotropy, which
800 results in ~~a strongly~~ laminated lower crustal fabrics (Carbonell and Smithson, 1991; Okaya et
801 al., 2004). Accordingly, a laminated lower crust may represent an overly reworked lower crust
802 that has been ductilely deformed over several orogenies. Opposite to the model by Meissner
803 (1989), such a horizontal reflectivity is observed along the Iberian Massif in areas where late
804 orogenic extension is absent or weak: the SPZ (Avalonia), the OMZ (peri-Gondwana), the not
805 extended CIZ, the WALZ and the CZ (Gondwana). Thus, we suggest that strong lamination in
806 the deep crust is probably a global characteristic of reworked lower crusts not affected by late
807 orogenic extension in the latest orogeny.

808 4.3. The Moho and crustal thickness in the Iberian Massif

809 The crust-mantle boundary, i.e., the Moho, is basically flat in the Iberian Massif except where
810 affected by ~~the~~ Alpine tectonics (Fig. 10). This is rather surprising as the lower crust seems to
811 be quite well preserved, suggesting that the Moho geometry has been flattened out through
812 slow, not invasive, readjustments.

813 Flat Mohos imply the existence of either isostatic and/or thermal, late to post-orogenic
814 processes that have managed to eliminate crustal roots (Cook, 2002). NW Iberia was affected
815 by late Carboniferous extension that heated and reworked the CIZ, possibly without significant
816 mantle involvement (Alcock et al., 2009, 2011), but producing crustal thinning (Palomeras et
817 al., 2017: see Moho depth map in Fig. 10). Thick and thermally mature crusts might experience
818 lateral flow of its low-viscosity deeper part that ~~contributed~~ ~~contributes~~ to reduce crustal roots
819 (Seyferth and Henk, 2004). This process might have partly occurred in the CIZ sampled by the
820 ESCIN-3.2 section (Fig. 6) where an outstanding change in lower crustal thickness and signature
821 exists, manifested by a thinner and very reflective lower crust in contrast to that to the E, in
822 the WALZ (1 vs 3 s TWT) or to the S, in the ALCUDIA profile (1 vs 5-6 s TWT). In the former, the
823 Variscan crust is still thick even though it experienced late-Variscan extension in its western
824 part and the whole area was slightly affected ~~offshore~~ by ~~Permian and Mesozoic the~~ extension,
825 ~~the latter~~ linked to the opening of the Bay of Biscay. In fact, underthrusting of the CZ lower
826 crust is still preserved in the eastern CIZ (Figs. 6 and 9).

827 In the SW Iberian Massif, a thick laminated lower crust is still observable while the Moho depth
828 is fairly constant (~10s TWT). Carboniferous-to-Permian isostatic rebound in response to
829 tectonic thickening, erosion and localized Permian thermal readjustments must have
830 contributed to flatten the Moho. However, seismic reflections show that crustal imbrication
831 into the mantle has locally survived post-orogenic Moho resetting. This indicates that isostatic
832 equilibrium has been reached in a long wavelength scale, but that local features can still
833 remain if they are supported by the crustal strength and do not pose an isostatic constraint.

834 4.4. The (missing) middle crust in the Iberian Massif (and elsewhere?)

835 One of the highlights of this work is the lack of a ~~seismic~~ layer that can be identified with ~~the a~~
836 ~~seismic~~ middle crust. But, what is the middle crust?

837 From a metamorphic point of view, the middle crust could be ascribed to the mesozone, which
838 may be correlated to the amphibolite facies, whose temperature ranges between 400-500 and
839 600-800°C, the precise limits depending on the pressure (Spear, 1993). In addition, the
840 epizone, between 200-250 and 400-500°C and typically represented by the greenschist facies,
841 is also a metamorphic entity which develops during metamorphism under several kilometers
842 of anchi- and no metamorphic rocks. The depths corresponding to these temperature intervals
843 vary with the geothermal gradient. For a Barrowian gradient, typical of a continental crust
844 undergoing collision, the depths for epizone and mesozone can be estimated around 10-20
845 and 20-30 (\pm 5) km respectively. Although the boundaries of these metamorphic zones might
846 have a gravity, i.e. density signature, they lack a seismic one. Furthermore, epi-, meso- and
847 catazonal rocks outcrop everywhere in any eroded orogenic belt defining a very complex
848 pattern that contrasts with the simplicity of seismic images. This implies that the metamorphic
849 middle-crust does not need to coincide with a hypothetical seismic middle crust, the former
850 often being part the upper crust in ancient orogens.

851 Seismic data are sensible to velocity and density contrasts and not to the absolute value of
852 velocity and density. If a sharp contrast exists, a reflection appears, but metamorphic zones
853 usually lack sharp boundaries. So from a seismic point of view, a middle crust should be a
854 crustal level bounded in its upper and lower parts by reflections indicative of the existence of
855 important impedance contrasts at its top and bottom. In this regard, only the IRB, intruded
856 between the upper and the lower crust (Carbonell et al., 2004; Simancas et al., 2003), and
857 providing conspicuous velocity contrasts (Palomeras et al., 2009, 2011) fulfill that
858 requirement. However, it is most probably an intrusion emplaced at a mid-crustal discontinuity
859 and does not represent the middle crust.

860 WA reflection seismic data from the northern Iberian Massif have often resulted in
861 multilayered models despite weak evidences of continuous reflectivity at these levels (Ayarza
862 et al., 1998; Fernández-Viejo et al., 1998, 2000; Pedreira et al., 2003). Even though local
863 velocity contrasts capable of providing weak and patchy reflectivity ~~contrasts~~ exist at different
864 crustal depths (e.g., thrust faults and normal detachments may represent lithological
865 boundaries with a noticeable velocity contrast), these are not orogen-scale features but local
866 reflectors. However, many of these reflections have been extrapolated and interpreted as
867 middle crust in seismic WA datasets, despite of being part of upper crust when compared with
868 NI data, i.e., lie above the mid-crustal discontinuity (e.g., Vp increase at ~10 km in Fig. 5e). In

869 this regard, the short wavelength heterogeneities of the crust have been often considered by
870 low resolution WA datasets as laterally continuous features (Levander and Holliger, 1992),
871 something that has led us to wrong models.

872 According to the above we argue that, in the Iberian Massif, no seismic middle crust can be
873 identified. In the hinterland, reflectors imaging deformation in the upper crust root in the top
874 of the lower crust. Only in ESCIN-1, which depicts the thin-skinned deformation of the CZ,
875 thrust faults root in a sole thrust and one could argue that the basement underneath these
876 shallow reflections represents the middle crust. But in the shallower part, to the E, early
877 Paleozoic and Neoproterozoic sediments occur on both sides of the sole thrust. Also, in the
878 deeper parts, to the W, the previous crystalline basement is probably involved in imbrications
879 affecting the upper crust. Thus, in our opinion, that non-reflective basement represents the
880 seismic upper crust.

881 In the Iberian Massif, the Paleozoic was deposited unconformably above Neoproterozoic
882 sediments which could be considered as its basement, but these were not metamorphic then.
883 Only in the OMZ, greenschist to amphibolite facies Neoproterozoic represents the Cadomian
884 basement, but it cannot be distinguished from the overlying Paleozoic metasediments in the NI
885 profiles. An even older crystalline basement of felsic composition exists, as indicated by
886 inherited zircons of 830-2000 Ma found in Ediacaran orthogneisses, Lower Ordovician
887 volcanics and Variscan granitoids that resulted from partial melting of such a basement
888 (Fernández-Suárez et al., 1998; Montero et al., 2007; Villaseca et al., 2012). Again, its upper
889 boundary is not imaged on NI profiles. These data also suggest that, in the Iberian Massif,
890 there are no crustal intervals that can be related to a seismic middle crust. Decoupling of
891 reflectivity, i.e. deformation, at a mid-crustal level led us to define just an upper and a lower
892 crust.

893 **4.5. Significance of a mid-crustal discontinuity: ~~the Conrad discontinuity?~~Geometry and** 894 **lateral extent**

895 Inspection of the Iberian Massif NI seismic dataset leads us to conclude that an orogenic-scale
896 mid-crustal discontinuity exists. This surface does not always provide a clear reflection, as in the
897 SPZ, but it is clearly defined by the geometry of the upper and lower crustal reflections,
898 asymptotically merging into it. The discontinuity coincides with the top of the lower crust,
899 which is often much more reflective than the upper crust and features a Vp increase.
900 Furthermore, this discontinuity has probably acted as a detachment for Variscan deformation
901 in the hinterland of the orogen and in the SPZ. However, in the CZ, the transition between
902 upper and lower crust is poorly defined, as most of its basement was not affected by Variscan
903 tectonics.

904 Simancas et al. (2013) already described this discontinuity on the basis of the asymptotic
905 geometry of the SPZ faults towards the middle of the crust. These authors concluded that its
906 depth greatly varies when reaching suture boundaries, where the discontinuity roots. Although
907 we do not observe a subduction zone in the reworked elusive suture between the SPZ and the
908 OMZ (Pérez-Cáceres et al., 2015), and interpret the OMZ/CIZ suture as an indentation between
909 two continental crusts, triggering imbrication into the mantle of the latter (crocodile
910 structure), we agree that this discontinuity would have eased the decoupling of the Iberian

911 crust, allowing subduction of its lower part while the upper part was deformed by folds and
912 thrust faults. In fact, this is clearly observed in the Alpine northward subduction of the Iberian
913 Massif lower crust underneath the CZ (ESCIN-2, Fig. 4) and also, in the Pyrenees, where a
914 detached Iberian lower crust subducts to the N (Teixell et al., 2018). But in the Iberian Massif,
915 the complexity of Variscan tectonics and late-Variscan crustal re-equilibration has mostly
916 removed evidences of a such mechanisms, although a comparable example has been
917 preserved in the NW: the thick lower crust imaged by ESCIN-3.3 is interpreted as
918 underthrusting of the CZ lower crust under that of the WALZ (Ayarza et al., 1998; Martínez
919 Catalán et al., 2003, 2012, 2014), ~~and~~ even reaching the CIZ, as shown in profile ESCIN-3.2.

920 Some authors have interpreted the Iberian Massif mid crustal discontinuity as the brittle-
921 ductile transition (e.g. Ehsan et al., 2014; Simancas et al., 2013). Indeed, it bounds a lower
922 crust, highly reflective and ductilely deformed from the upper crust. However, Variscan ductile
923 deformation occurs also above the discontinuity in the entire Iberian Massif, with the
924 exception of the CZ. On the other side, if we deal with present deformation mechanisms, it is
925 unlikely that the brittle-ductile transition, which depends on ~~P and T the values of P and T,~~
926 ~~coincided~~ with the described discontinuity, because, i) it does not necessarily imply an
927 impedance contrast (Litak and Brown, 1989), and ii) according to figure 10, the depth of the
928 discontinuity varies from 4 s TWT (~12 km, ALCUDIA section) to 8 s TWT (~21 km, ESCIN-2
929 section) which would imply unrealistic ~~crustal P and T variations on P and T in the present~~
930 ~~crust~~. Accordingly, a different interpretation must be sought.

931 The Iberian laminated lower crust is probably very old. Granulites dredged in Mesozoic
932 sediments of the Cantabrian margin have yielded ages of up to 1400 Ma (Capdevila et al., 1980
933 and references therein). Even older ~~Ar-Ar values (~2500 Ma)~~ have been obtained for the
934 Galicia Bank granulites (Gardien et al., 2000), ~~which featured Ar-Ar ages of up to 2500 Ma.~~
935 These granulites have been deformed ductilely during several orogenies. But rocks lying above
936 the lower crust, whatever their nature, are separated from it by a discontinuity that fosters
937 decoupled deformation between both crustal layers. Accordingly, the observed mid-crustal
938 discontinuity probably represents a rheological boundary that separates rocks that have been
939 deformed differently. This boundary, located at the top of the lower crust, represents a
940 velocity increase as the latter is probably composed of dense granulites and includes relatively
941 abundant basic rocks, which makes it easily identifiable in NI and WA seismic sections.

942 The geometry of this discontinuity and its depth, together with that of the Moho (Fig. 10),
943 provide insights of the evolution of the Iberian Massif. Along the SW Iberian Massif, the mid-
944 crustal discontinuity is sub-horizontal and lies at a depth between 4-6 s TWT. In the OMZ, the
945 intrusion of the IRB allows to establish its depth in the top or the bottom of this feature, but in
946 average, its location would fit the above given values. However, in the center and NW, the
947 position of the discontinuity varies, deepening down to 8 s TWT (Figs. 9 and 10).

948 The low resolution ~~noise~~-autocorrelation models obtained along the CIMDEF profile show
949 confusing results along the central Iberian Massif. There, the mid-crustal discontinuity might
950 lie at 5-6 s TWT, deepening around the ICS to 8 s TWT as it has been affected by pervasive
951 extension and melting, thus defining a thin lower crust (2 s TWT, ~6 km, Figs. 9 and 10).
952 Accordingly, this feature appears redefined in this area, and now follows the geometry of the

953 ICS batholith. The ~~change in the depth and geometry of this~~ deepening of this discontinuity and
954 ~~the associated~~ thinning of the lower crust might have allowed coupled deformation, letting
955 part of the upper crust to the S of the ICS to underthrust it (Andrés et al., 2020). This would
956 foster the 400-500 m topographic change between the N and S foreland basins of this Alpine
957 mountain range (Fig. 9). In fact, Simancas et al. (2013) argues that coupled crustal deformation
958 takes place when a relatively weak lower crust exists something that might well represent the
959 context of the ICS. The resulting geometry of this Alpine reactivation and its topographic
960 imprint is different to that observed to the N, in the CZ, where late orogenic extension and
961 melting does not exist and the mid-crustal discontinuity has been preserved.

962 ~~On the other hand, the lower crust imaged along the CIMDEF transect presents a conspicuous~~
963 ~~internal reflection that could also be interpreted as the top of the lower crust (Andrés et al.,~~
964 ~~2020). If this were the case, the lower crust would be even thinner along the entire section,~~
965 ~~matching the characteristics observed to the N of the ALCUDIA section and in the ESCIN-3.2~~
966 ~~line. In any case, we argue that the mid-crustal discontinuity and the lower crust seen in the~~
967 ~~CIMDEF profile are both probably reworked by extension but not totally re-equilibrated and~~
968 ~~thus, its seismic image is confusing to the N and S of the ICS. Moho depth models (Fig. 10)~~
969 ~~derived from shear wave tomography (Palomeras et al., 2017) indicate that along the CIMDEF~~
970 ~~profile the crust is thin (except in the Alpine root) but not as much as in NW Iberia, so that~~
971 ~~lower crustal extension and re-equilibration may have not been as intense as in the GTMZ and~~
972 ~~CIZ of the NW Iberian Massif.~~

973 The most outstanding change in the mid-crustal discontinuity architecture appears in NW
974 Iberia, along the ESCIN-3.2 profile. This section features the thinnest crust (9 s TWT)
975 accompanied by the thinnest lower crust (~1 s TWT). The mid-crustal discontinuity lies at 8 s
976 TWT in contrast to the depth where it appears in the neighbouring ESCIN-3-3 and ESCIN-1 lines
977 (6 ~~and~~ 8 s TWT), suggesting that it has been redefined. Nevertheless, clear reflections root in
978 its upper part indicating that it still acted as a discontinuity/detachment. ~~The depth of this~~
979 ~~feature in NW Iberia is similar to that of the high amplitude lower crustal internal reflection~~
980 ~~near the ICS.~~ Accordingly, we suggest that in NW Iberia, gravitational collapse followed by
981 crustal melting and extension has thinned the crust (Fig. 10), and specially the lower crust,
982 relocating the mid-crustal discontinuity.

983 NW Iberia was importantly thickened (up to 50-70 km) due to the emplacement of the GTMZ
984 allochthonous complexes. Thermal models by Alcock et al. (2009, 2015) show that as a result,
985 the upper mantle continued increasing its temperature 60-65 Ma after the start of
986 compressional deformation at 360 Ma. This implies large thinning of the mantle lithosphere,
987 from 70 to 25-30 km, due to the ascent of the 1300 °C isotherm. It is not surprising that the
988 lower crust there became highly extended as a consequence of the heat increase, as in the
989 models it reached 800 °C after 45 Ma and 900 °C after 55 Ma (315-305 Ma).

990 4.5.1. The Iberian Massif Conrad discontinuity

991 The idea of a mid-crustal velocity discontinuity was put forward in the 1920's (Conrad, 1925).
992 Early analysis of natural source earthquake recordings and later images from controlled source
993 seismic reflection data provided further evidences that supported a clear distinction between
994 upper and lower crust. These evidences led to consider ~~ing~~ the Conrad discontinuity as a global

995 scale feature present in the continental crust. However, this was later challenged as some
996 results of deep seismic reflection profiling did not show a clear distinction between upper and
997 lower crust (Litak and Brown, 1989).

998 Mid-crustal discontinuities have, however, been observed often in different types of seismic
999 data worldwide (e.g., Fianco et al., 2019; Hobbs et al., 2004; Melekhova et al., 2019; Oncken,
1000 1998; Ross et al., 2004; Snelson et al., 2013). Important changes in the rheology of the crust
1001 have also been reported at those depths (Maggini and Caputo, 2020; Wever, 1989) supporting
1002 the idea that a mechanical boundary must exist. Thus, we suggest that, even though it is not
1003 observed everywhere (Litak and Brown, 1989), this feature is an orogen-scale, world class
1004 continental crustal discontinuity (Artemieva, 2009), often coinciding with the top of the highly
1005 laminated lower crust (when there is one). Its existence might determine the way the crust
1006 deforms, easing decoupled deformation. Orogenic evolution, i.e. rifting, extension, melting,
1007 etc. may modify it or even erase it, thus its existence and geometry might help us to
1008 understand the geologic history of continents. In this regard, and coming back to the long-
1009 forgotten discussion of the nature of the Conrad discontinuity (Conrad, 1925) and its position
1010 on top of the laminated lower crust (Wever, 1989), we suggest that, in the Iberian Massif, the
1011 observed mid-crustal feature fulfills the characteristics of this debated discontinuity. Its clear
1012 signature and regional extension contributes to unravel its nature and significance.

1013 **5. Conclusions**

1014 Normal incidence seismic data acquired across the Iberian Massif in the last 30 years have
1015 provided an entire [section-image](#) of a well exposed and almost complete part of the European
1016 Variscides. Existing gaps in the central part have been recently sampled by passive source
1017 seismic recordings (noise and earthquakes) that provide fairly good constraints on the crustal
1018 structure.

1019 Results show that crustal thickness varies from ~ 9 s TWT in late-Variscan extended areas (NW
1020 of the Central Iberian Zone), to ~ 10 s TWT (30-33 km) in the external South Portuguese Zone to
1021 ~ 12 s TWT (36-38 km) in the internal West Asturian-Leonese Zone. Alpine reactivation has
1022 managed to further thicken the crust to at least ~ 14 s TWT (42-45 km) in the external
1023 Cantabrian Zone and to 35-38 km in the Iberian Central System, a Tertiary orogenic belt
1024 developed in Central Spain. The top of an often thick (up to 6 s TWT) and very reflective lower
1025 crust helps to define a mid-crustal discontinuity across the entire Iberian Massif. This boundary
1026 represents a level where reflections from the upper and lower crust merge asymptotically,
1027 thus suggesting that it has often acted as a detachment or a decoupling level. Its position and
1028 geometry varies mostly in relation to the late Variscan evolution. Accordingly, it is deeper in
1029 NW and central Iberia (~ 8 s TWT), where Variscan crustal thickening was important and
1030 gravitational collapse melted and extended the crust, thus defining a very thin lower crust.
1031 However, it appears between 4-6 s TWT to the SW, where the crust did not thicken as much
1032 and its original structure is better preserved, being later re-equilibrated through slow isostasy
1033 and erosion.

1034 This discontinuity exists in all the Iberian Massif tectonic zones, regardless of their Gondwana
1035 or Avalonia affinity, thus suggesting it is an orogenic-scale discontinuity. We interpret it as the
1036 rheological boundary between an overly ductilely deformed old lower crust and a

1037 heterogeneous variably (often also ductilely) deformed upper crust that mostly (but not only)
1038 shows evidences of the latest orogenic event. Its geometry, position and extent match the
1039 characteristics defined for the long-forgotten Conrad discontinuity. The identification of similar
1040 features in normal incidence profiles worldwide supports its inclusion as a major crustal
1041 discontinuity.

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Código de campo cambiado

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1505 Table 1: Acquisition parameters of the NI seismic profiles shown from figures 3 to 8 and
 1506 former processing flows. These are grouped according to their similarities.

1507	Acquisition parameters	ESCIN-1 (onshore)	ESCIN-2 (onshore)
1508	Source	Dynamite-single-hole	Dynamite-single-hole
1509	Charge	20 kg at 24 m depth	20 kg at 24 m depth
1510	Trace Interval	60 m	60 m
1511	# of traces	240	240
1512	Spread configuration	Symmetrical Split-spread	Symmetrical Split-spread
1513	Fold	30	30
1514	Geophones per group	18	18
1515	Spread length	14.5 km	14.5 km
1516	Sample rate	4 ms	4 ms
1517	Kirchoff time migration	5600 m/s	5600 m/s
1518			
1519		ESCIN-3.2 and ESCIN-3.3 (offshore)	
1520			
1521	Source	Airgun (5490 cu.in)	
1522	Shot spacing	75 m	
1523	Receiver interval	12.5 m	
1524	Spread length	4500 m	
1525	Fold	30	
1526	Internal offset	240 m	
1527	Sample rate	4 ms	
1528	Record length	20 s	
1529	Kirchoff time migration	5200 m/s	
1530			
1531		IBERSEIS (onshore)	ALCUDIA (onshore)
1532			
1533	Source	4, 22T vibrators	4 (+1), 22T vibrators
1534	Recording instrument	SERCEL 388, 10 Hz	SERCEL 388, 10 Hz
1535	# active channels	240 minimum	240 minimum
1536	Station spacing	35 m	35 m
1537	Station configuration	12 geophones	12 geophones
1538	Source spacing	70 m	70 m
1539	Sweep frequencies	non-linear 8-80 Hz	non-linear 8-80 Hz
1540	Sweep length	20 s	20 s
1541	Listening time	40 s	40 s
1542	Sample rate	2 ms	4 ms
1543	Spread type	Asymmetric split-spread	Asymmetric split-spread
1544	Nominal fold	60 (minimum)	60 (minimum)
1545	Kirchoff time migration	6000 m/s	6200 m/s
1546			
1547			
1548	Figure captions		

1549

1550 Figure 1: (a) Map of the Iberian Peninsula showing the outcrops of the Variscan basement and
1551 the subdivision in zones of the Iberian Massif. The main strike-slip shear zones, traces of
1552 Variscan folds and gneiss domes are also included. Blue lines show the position of normal
1553 incidence seismic reflection profiles and that of the CIMDEF transect. See legend for
1554 abbreviations. (b) Map of the outcropping granitoids in the Iberian Peninsula together with the
1555 main structures.

1556 Figure 2: Processing flow carried over the SEG-Y original stack sections. This task was
1557 geared to improve and homogenize the resolution of the seismic images while creating
1558 new migrated sections. See Martínez García, (2019) for further details.

1559 Figure 3: Cross-section (a) and depth converted time migrated section ($v=5600$ m/s) along
1560 the NI seismic profile ESCIN-1 (Fig. 1), without (b) and with interpretation (c). A sketch of
1561 the most important features is presented in (d). CDP: Common Depth Point. TWT: Two-way
1562 travel time. WALZ: West Asturian-Leonese Zone. CZ: Cantabrian Zone. The position of the
1563 Narcea Antiform is indicated. Red dashed lines represent the boundaries provided by chaos
1564 and variance attribute analyses. Nomenclature for reflections goes as follows: (ot),
1565 outcropping thrusts; (t), thrusts affecting the basement; (b), indifferiated basement;
1566 (st), sole thrust; (c), top of the lower crust; (ir), lower crust internal reflectivity; (m) is the
1567 Moho; (mig) are curved features resulting from the migration of discontinuous reflections.
1568 Depth conversion is based on migration velocities.

1569 Figure 4: Cross-section ((a) modified from Gallastegui et al., 2016) and depth converted
1570 time migrated section ($v=5600$ m/s) of the NI seismic profile ESCIN-2 (Fig. 1), without (b)
1571 and with interpretation (c). A sketch of the most important features is presented in (d). A
1572 1D velocity profile as modeled from wide-angle data (Pulgar et al., 1996) appears in (e).
1573 CDP: Common Depth Point. TWT: Two-way travel time. CZ: Cantabrian Zone. Some
1574 discontinuous reflections are traced on the basis of their geometry on the stack image
1575 (Pulgar et al., 1996). Red dashed lines represent the boundaries provided by chaos and
1576 variance attribute analyses. (s), sediments; (t), thrusts; (ot), outcropping thrusts; (ps),
1577 Paleozoic sediments; (c), top of the lower crust; (lc), lower crust; (m), Moho. Depth
1578 conversion is based on migration velocities.

1579 Figure 5: Cross-section (a) and depth converted time migrated section ($v=5200$ m/s) of the
1580 NI seismic profile ESCIN-3.3 (Fig. 1), without (b) and with interpretation (c). A sketch of the
1581 most important features is presented in (d). A 1D velocity profile as modeled from wide-
1582 angle data (Ayarza et al., 1998) is presented in (e). CDP: Common Depth Point. TWT: Two-
1583 way travel time. CZ: Cantabrian Zone. WALZ: West Asturian-Leonese Zone. CIZ: Central
1584 Iberian Zone. The offshore projection of the Viveiro Fault is indicated. Red dashed lines
1585 represent the boundaries provided by chaos and variance attribute analyses. (s),
1586 sediments; (t), thrusts; (st), sole thrust; (ef), extensional features; (c), top of the lower
1587 crust; (lc), lower crust; (ir) lower crust internal reflectivity; (m), Moho; (sc), subcrustal
1588 reflections; (mig) curved features resulting from the migration of discontinuous reflections.
1589 Depth conversion is based on migration velocities.

1590 Figure 6: Cross-section (a) and depth converted time migrated section ($v=5200$ m/s) of the
1591 NI seismic profile ESCIN-3.2 (Fig. 1), without (b) and with interpretation (c). A sketch of the

1592 most important features is presented in (d). CDP: Common Depth Point. TWT: Two-way
1593 travel time. GTMZ: Galicia-Trás-os-Montes Zone. CIZ: Central Iberian Zone. (s), sediments;
1594 (ef), extensional features; (lc), lower crust; (m), Moho; (sc), subcrustal reflections; (czlc) CZ
1595 underthrust lower crust. Depth conversion is based on migration velocities.

1596 Figure 7: Cross-section ((a) modified from Martínez Poyatos et al., 2012) and depth
1597 converted time migrated section ($v=6200$ m/s) of the NI seismic profile ALCUDIA (Fig. 1),
1598 without (b) and with interpretation (c). A sketch of the most important features is
1599 presented in (d). A 1D velocity profile as modeled from wide-angle data (Palomeras et al.,
1600 in press) is overlapped in (d). CDP: Common Depth Point. TWT: Two-way travel time. CIZ:
1601 Central Iberian Zone. OMZ: Ossa-Morena Zone. (cu): Central Unit; (vf): vertical folds; (g):
1602 granites; (ef): extensional features; (c): top of the lower crust; (lc): lower crust; (cc) and
1603 (cc'): crocodile structure; (m): Moho. Red dashed lines represent the boundaries provided
1604 by chaos and variance attribute analyses. Depth conversion is based on migration
1605 velocities.

1606 Figure 8: Cross-section ((a), modified from Simancas et al., 2003) and depth converted time
1607 migrated section ($v=6000$ m/s) of the NI seismic profile IBERSEIS (Fig. 1), without (b) and
1608 with interpretation (c). A sketch of the most important features is presented in (d). Two 1D
1609 velocity profiles as modeled from wide-angle data (Palomeras et al., 2009) are overlapped
1610 in (d). CDP: Common Depth Point. TWT: Two-way travel time. CIZ: Central Iberian Zone.
1611 OMZ: Ossa-Morena Zone. SPZ: South Portuguese Zone. (t), thrusts; (ot), outcropping
1612 thrusts; (irb), Iberseis Reflective Body; (c), top of the lower crust; (lc), lower crust; (ir) lower
1613 crust internal reflectivity; (lct), lower crust thrusts; (m), Moho. Depth conversion is based
1614 on migration velocities.

1615 Figure 9: Joint geological interpretation of all the seismic sections (normal incidence and
1616 seismic noise) whose location is shown in figure 1. (a): ESCIN-1, ESCIN3-3 and ESCIN-3.2.
1617 (b): ALCUDIA and IBERSEIS. (c): CIMDEF (Andrés et al., 2020). Special attention should be
1618 paid to the depth and geometry of the Moho and mid-crustal discontinuity. Alpine
1619 structures (i.e. crustal thickening) appear in ESCIN-1 and in CIMDEF. The rest are Variscan
1620 features.

1621 Figure 10: Map of the Moho depth as derived from tomography of shear waves (seismic
1622 noise and earthquakes, Palomeras et al., 2017) with the projection of the seismic profiles
1623 already shown in figure 1 and described along the text. A sketch of the geometry of the
1624 main discontinuities (Moho and Conrad) is also shown.

Código de campo cambiado