

# The effect of low-viscosity sediments on the dynamics and accretionary style of subduction margins

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**Abstract.** Observations of sediments at subduction margins appear to divide them into two classes: accretionary and erosive. Accretionary margins are dominated by accretion of thick piles of sediments ( $>1$  km) from the subducting plate, while tectonic erosion is favored in regions where the sedimentary cover is  $<1$  km. The consequences of the two styles of margins on subduction dynamics remain poorly resolved. In this study, we used 2-D numerical simulations of subduction to investigate how low-viscosity sediments influence subduction dynamics and margin type through plate coupling. We vary the thickness and viscosity of the sediment layer entering subduction, the thickness of the upper plate, and the driving velocity of the subducting plate (i.e., kinematic boundary conditions). Diagnostic parameters are extracted automatically from numerical simulations to analyze the dynamics and differentiate between modes of subduction margin. Our results show three geometry modes of subduction interface: a) tectonic erosion margin, b) low-angle accretionary wedge margin, and c) high-angle accretionary wedge margin. We find that the properties of the sediment layer modulate the extent of viscous coupling at the interface between the subducting and overriding plates. When the viscous coupling is increased (high viscosity, thin layer of sediments), an erosive margin will be favored over an accretionary margin. On the other hand, when the viscous coupling is reduced (low viscosity), sediments are scrapped-off the subducting slab to form an accretionary wedge. Models that develop tectonic erosion margins show small radii of curvature, slow convergence rates and thin subduction interfaces, while models with accretionary margins show large radii of curvature, faster convergence rates and dynamic accretionary wedges. These diagnostics parameters are then linked with observations of present-day subduction zones.

## 1 Introduction

Sediment subduction at convergent plate boundaries has long been recognized to play an important role in the dynamics of our planet as they can provide direct feedbacks between plate tectonics, climate, and life. Quantifying the sediments mass flux through subduction zones is important for understanding: i) generation of arc magmas and the problem of petrogenesis of continental crust (e.g., Plank and Langmuir (1998); Kelemen and Manning (2015)), ii) whether large volumes of existing continental crust are ever recycled back into the mantle over long periods of geologic time, and iii) cycling of volatiles from Earth's crust and atmosphere to its deep interior (e.g., Hawkesworth et al. (1997); Plank and Langmuir (1998); Dasgupta and Hirschmann (2010)). Regarding the latter, carbon and water global cycles in particular depend greatly on the amount of subducted sediments (e.g., Plank and Manning (2019); Dutkiewicz et al. (2018); Merdith et al. (2019)), which in turn

have important implications for climate stability (Kasting, 1989), biogeochemical cycles (Husson and Peters, 2017), and the rheology of the mantle (Hirth and Kohlstedt, 1996).

Sediments are also fundamental to the dynamics of subduction zones and plate tectonics. The lubricating effect of sediments at the plate interface (referred here as the domain between the lower and upper plates, as defined in Agard et al. (2018)) was recognized more than thirty years ago (i.e., Shreve and Cloos (1986)), and is critical for the mechanism of plate tectonics (e.g., Tackley (2000); Bercovici (2003); Bercovici and Ricard (2014); Sobolev and Brown (2019)). On a regional scale, sediments influence patterns of deformation by controlling the morphologies of subduction interfaces, accretionary prisms and forearc basins (Clift and Vannucchi, 2004; von Huene and Scholl, 1991b; Melnick and Echtler, 2006; Simpson, 2010). Moreover, sediments occupying the shallow seismogenic subduction interface, for example, appear to influence seismic coupling and the frequency of megathrust earthquakes (e.g., Moore and Saffer (2001); van Rijsingen et al. (2018); Heuret et al. (2012); Brizzi et al. (2020); Bangs et al. (2020)).

However, the relative role of sediments on subduction dynamics and evolution remains unclear. The force balance during subduction includes the potential energy change of the negative buoyancy of the subducting slab, viscous dissipation in the mantle, bending of the lithosphere, and slab-upper plate interface (e.g., Conrad and Hager (1999)). Considerable effort in subduction dynamics studies focused on quantifying dissipation due to slab bending (e.g., Conrad and Hager (1999); Becker et al. (1999); Capitanio and Morra (2012); Garel et al. (2014a)). That is because slabs were assumed to be strong (3000-5000 times stiffer than the mantle) and sediments weak, implying a low coupling degree (low shear stresses) at the interface (Conrad and Hager, 1999; Duarte et al., 2015; Billen and Hirth, 2007). Laboratory and numerical models, instead, suggest that slabs may be weaker (only 150-500 times stiffer than the mantle) (Funicello et al., 2008; Moresi and Gurnis, 1996; Zhong and Davies, 1999), implying a more prominent role for plate interface dissipation than previously thought. Recent simplified energy balance calculations by Behr and Becker (2018) also suggest that sediment subduction may modulate plate speeds, with sediment-lubricated plates subducting faster than slabs with metabasaltic (i.e., exposed mantle rocks) interfaces.

The lubricating effect of sediments has largely been considered an implicit assumption in previous large-scale subduction dynamics studies. The subduction interface in numerical models is typically implemented by imposing either a fixed interface layer (i.e., subduction channel or weak fault) or a layer at the top of the subducting plate (i.e., weak crust) that is advected with the flow and continuously entrained into the decoupling region. In this way, the weakening effect of sediments, or any other deformation-localizing processes such as damage, grain size reduction, and fabric development, is parameterized by setting a low viscosity or low friction coefficient of the interface material (Gerya et al., 2002). This is a necessary model component for stable asymmetric subduction (Petersen et al., 2017; Cramer and Tackley, 2015; Gerya, 2009; Sandiford and Moresi, 2019) and has become an increasingly common strategy in the last decade (Babeyko and Sobolev, 2008; Capitanio et al., 2010; Magni et al., 2012; Chertova et al., 2012; Cizkova and Bina, 2013; Garel et al., 2014b; Pusok and Kaus, 2015; Agrusta et al., 2017; Pusok and Stegman, 2019).

Indirect observations suggest subduction interfaces are not discrete isosurfaces separating two plates but are rather exemplified by melange zones (Shreve and Cloos, 1986; Vannucchi et al., 2008; Agard et al., 2018), that incorporate material from the subducting plate, the accretionary prism, and the upper plate (Menant et al., 2020; Angiboust et al., 2021). Moreover, when

sediments are considered at the trench, convergent margins appear to fall into one of two classes: accretionary and erosive (i.e., Clift and Vannucchi (2004), Supplementary material, Figure S1). Accretionary margins develop from an accumulation of material from the subducting plate being transferred onto the overriding plate, either by frontal off-scraping at the trench axis or by underplating of the forearc wedge above the decollement at greater depths (Angiboust et al., 2021). Erosive margins develop from a strong coupling between overriding and subducting plates that results in erosion of the underside of the upper plate, as indicated by margin truncation and forearc subsidence (von Huene and Scholl, 1991a; Clift and Vannucchi, 2004; Straub et al., 2020). Accretionary margins are dominated by accretion of thick piles of sediments ( $> 1$  km) from the subducting plate, while tectonic erosion is favored in regions where the sedimentary cover is  $< 1$  km and showing long-term landward retreat of the trench.

Both accretionary and tectonic erosion margins are widely distributed. Clift and Vannucchi (2004) classified the global subduction zones in the two categories based on their dominant mode in the last 10 Myr, and found that 43% of global subduction margin is accretionary, and 57% is tectonic erosion. The implications of this equally-distributed duality in margin type to the global subduction system has not been investigated. Convergence rates for the two types of margins seem to correlate well with sediment thickness, the taper angle and radius of curvature (i.e., Supplementary material Figure S1, replotted data from Clift and Vannucchi (2004) and Wu et al. (2008)). De Franco et al. (2008) also observe a correlation between the margin type and upper plate strain (i.e. proxy for back-arc extension). Lamb and Davis (2003) went further to suggest that the type of margin can affect mountain building, with tectonic erosion producing a higher degree of coupling between the subducting and upper plate. They argue that changes from a sediment-rich to sediment-starved subduction regime during Cenozoic climatic cooling may have been responsible for the rise of the Andean mountain belt.

It is clear that the consequences of the two styles of margins on large-scale subduction dynamics remain poorly resolved and have not been explored extensively with numerical models. Accretionary margins have been investigated in more details, following the theory for critical Coulomb wedges by Dahlen (1984) and Dahlen et al. (1984) (i.e., in analogue models such as Lallemand et al. (1994); Gutscher et al. (1998) and numerical models such as Beaumont et al. (1999); Selzer et al. (2008); Ruh (2017); Menant et al. (2020)). However, accretionary margins have generally not been studied in the same framework as erosion style margins. Thus, we identify a number of outstanding questions regarding the influx of sediments to trenches and the style of margin that could be addressed with numerical models: Why some margins accrete sediments while others do not? What is the feedback between sediment fluxes and subduction dynamics? How much sediment material gets subducted into the mantle?

In this study, we run systematic 2-D numerical simulations of ocean-ocean subduction to investigate how low-viscosity sediments influence subduction dynamics and the plate coupling. We aim to understand what causes convergent margins to either accrete material delivered by the subducting plate or, alternatively, to subduct the trench sediment pile and even erode the basement of the overriding plate. The purpose of the present work is not to model in detail the dynamics of accretionary or erosive margins, but rather, by carrying out numerical experiments on the effect of sediments in geometrically simple configurations of subduction, to further understand the occurrence and evolution of each style of margin.

95 We begin our investigation with a discussion of the numerical setup and diagnostics. We then extract automatic diagnostics  
 from numerical results to evaluate regimes and compare to available observations in natural subduction zones. In particular, we  
 considered a range of dependent and independent variables from statistical analyses of present-day subduction zones (Section  
 2; Clift and Vannucchi, 2004; Lallemand et al., 2005; Wu et al., 2008; De Franco et al., 2008; Heuret et al., 2012) to constrain  
 and validate the results of numerical models. This study intends to consolidate insights from numerical models of subduction  
 100 with an integrated set of global observations.

## 2 Methods

Numerical models presented below are purely mechanical. We solve for the slow-creeping motion of solid materials over a  
 timescale of million of years, known as Stokes equations. They are comprised of the equations of conservation of mass and  
 momentum, assuming incompressibility and neglecting thermal diffusion, which are given by:

$$105 \quad \nabla \cdot \mathbf{v} = 0, \quad (1)$$

$$-\nabla P + \nabla \cdot \boldsymbol{\tau} + \rho \mathbf{g} = 0, \quad (2)$$

$$(3)$$

where  $\boldsymbol{\tau}$  is deviatoric stress tensor,  $P$  is pressure,  $\rho$  density,  $\mathbf{g}$  is the gravity vector, and  $\mathbf{v}$  is velocity. Thus, the lithosphere  
 and mantle materials are assumed to behave as a continuous medium deforming by steady state creep over long time intervals  
 110 (Turcotte and Schubert, 2014). We use a variable viscosity constitutive relationship  $\tau_{ij} = 2\eta\dot{\epsilon}_{ij}$ , where  $\eta$  is the Newtonian  
 viscosity, constant for each material phase,  $\dot{\epsilon}_{ij} = \frac{1}{2} \left( \frac{\partial v_i}{\partial x_j} + \frac{\partial v_j}{\partial x_i} \right)$  is the deviatoric strain rate tensor, and  $i, j$  represent spatial  
 directions following the Einstein summation convention.

The above equations are solved using the parallel 3-D finite difference code LaMEM (Lithosphere and Mantle Evolution  
 Model) (Pusok and Kaus, 2015; Kaus et al., 2016). We use a pseudo 2-D Cartesian domain in an approach similar to Pusok  
 115 and Stegman (2019, 2020), meaning we consider infinite extension in the  $y$ -direction. A Lagrangian marker-in-cell method  
 (Harlow and Welch, 1965; Gerya, 2009) is used for accurately tracking distinct material domains (Pusok et al., 2017) as they  
 undergo extensive deformation due to creeping flow. We also employ an internal free surface, using the "sticky-air" approach  
 (Schmeling et al., 2008; Cramer et al., 2012), with a free surface stabilization algorithm (Kaus et al., 2010) that allows for the  
 development of topography. The details of the model setup follow below.

### 120 2.1 Model setup

We performed 2-D numerical simulations of ocean-ocean subduction (Table 1, Figure 1) to investigate the role of sediments  
 on slab dynamics and topography. The model domain is 6000 km across and 1500 km deep. The computational domain  
 has a variable grid spacing, with higher resolution in the upper mantle asthenosphere and close to the subduction trench  
 (minimum and maximum grid spacings:  $\Delta x \in [1.73, 15.62]$  km,  $\Delta z \in [2, 16]$  km). Free-slip boundary conditions are imposed  
 125 on all boundaries and a 60-km layer of "sticky-air" on top of the plates. Previous studies have shown that using a free surface

(i.e., sticky-air method in this case) instead of a free-slip top boundary dramatically changes subduction style (Kaus et al., 2010; Cramer and Tackley, 2015). Here, the rock-sticky-air interface represents an internal free surface formulation, from which topography is calculated.

The initial model setup and material parameters are similar to the ones used in Pusok and Stegman (2019, 2020). These models were used to investigate the dynamics of same-dip double subduction systems, and to explain dynamical processes leading to the fast convergence history between India and Eurasia in the Cretaceous. Here, we model a single subduction setting (Figure 1). The model consists of an oceanic plate subducting beneath another oceanic upper plate (i.e., ocean-ocean subduction). The length of both plates is 2500 km and they are not attached to the side walls, mimicking mid-ocean ridges at the trailing edge of the plate away from the trench. Additional experiments with a fixed upper plate to the wall are presented in the Supplementary material (Figure S8).

The subducting plate lithosphere has a thickness of 80 km with a 20 km thick core and 15 km combined weak crust and sediments. Behr and Becker (2018) estimate that metasediments and serpentinites can be more than two orders magnitude weaker than the reference asthenospheric mantle viscosity, while metabasalts are predicted to be of the same order of magnitude at temperatures between 600-800°C. In our model, we assume that the weak crust is formed by a layer of highly-fractured metabasalts (dark green lithology in Figure 1a,b) overlain by a layer of metasediments (light green lithology in Figure 1a,b). Thus, the viscosity structure for the combined weak crust and sediments represents a parameterization of the strength weakening with depth due to hydration and weak sediment cover. The combined initial thickness of sediments and crust remains the same for all simulations.

Material parameters are the ones used in Enns et al. (2005) and Pusok and Stegman (2019), where the upper mantle asthenosphere has a reference density ( $\rho_0 = 3300 \text{ kg/m}^3$ ) and viscosity ( $\eta_0 = 2.8 \times 10^{20} \text{ Pa.s}$ ), the plates are  $85 \text{ kg/m}^3$  denser, and have a variable viscosity structure (slab:  $500 \times \eta_0$ , strong core:  $5000 \times \eta_0$  and weak crust:  $\eta_0$ ). The transition to the lower mantle is marked by a viscosity jump of 50 in the reference models, consistent with previous estimates of the viscosity jump in the Earth's mantle (Quinteros et al., 2010; Rudolph et al., 2015). For initial conditions, we impose a slab radius of curvature of 150 km and a subduction depth of 200 km, which is enough to initiate subduction. We vary the properties of the sediment layer (viscosity, thickness) and thickness of upper plate as explained below and in Table 1. All other parameters are kept the same among simulations. By not changing the density of the sediments or the slab geometry, the magnitude of initial slab-pull force is the same among simulations. We also consider constant sediment fluxes at the trench.

## 2.2 Input and diagnostics parameters

Previous studies investigated the role of upper plate and subduction plate parameters, such as thickness and strength (i.e., Holt et al. (2015); Brizzi et al. (2020)). Here, we focus on factors acting directly on the subduction interface. In particular, we use a result from Currie et al. (2007) and Cizkova and Bina (2019), which found that the effect of sediment buoyancy and viscous entrainment by the subducting plate are the main factors controlling the behaviour of slab and subducted sediments. Thus, we primarily vary the thickness and viscosity of the sediment layer and those of the upper plate. We also extract automatic

diagnostics from our numerical models that can be compared to parameters available for the natural subduction systems.

160 Parameters discussed in this section are listed in Table 2.

**Input parameters.** For the Newtonian rheology used, the input parameters for each phase (viscosity, thickness, density) are categorized into three groups: 1) subducting plate (slab) parameters, 2) interface (sediments) parameters, and 3) upper plate parameters (Table 2). The following parameters were varied: thickness and viscosity of sediments ( $h_{\text{sed}} = [5, 10]$  km,  $\eta_{\text{sed}} = [0.01\eta_0, 0.1\eta_0, \eta_0]$  Pa.s), and thickness of upper plate ( $h_{\text{UP}} = [50, 80, 100, 150]$  km). The sediment viscosity is consistent  
165 with estimates from Behr and Becker (2018) and the variation in upper plate thickness mimics variable plate ages.

Global compilation studies show that sediment thickness goes from 0-12 km (e.g., Laske et al. (2013); Dutkiewicz et al. (2015); Straume et al. (2019)). Sediment thickness variation in our model setup is, thus, on the upper limit. However, a no-sediment case is represented by high-sediment viscosity (i.e., higher proportion of crust at the interface). The thickness and viscosity of sediments will control the subduction interface shear stresses, while the upper plate thickness will control the  
170 interface length.

**Diagnostics parameters.** The diagnostics (or system-response) parameters calculated from each simulation are also categorized into three groups: 1) subduction (slab) diagnostics, including convergence rate, radius of curvature, trench velocity, 2) interface (sediments) diagnostics, including wedge angle and width of the accretionary wedge, 3) upper plate diagnostics, including trench depth, maximum and mean topographic amplitude in the upper plate. These variables are compatible with  
175 parameters derived from statistical analyses of present-day subduction zones (i.e., Clift and Vannucchi (2004); Lallemand et al. (2005); Wu et al. (2008); De Franco et al. (2008); Heuret et al. (2012)), which will be used to constrain and validate the results of numerical models. Diagnostics parameters are calculated at every time step. Their evolution (steady-state or transient) will constitute the basis of our parameter analysis in Section 3.3.

A schematic of how slab parameters, such as plate convergence ( $u_0$ ), trench retreat ( $u_T$ ), and radius of curvature ( $R_C$ ), are  
180 calculated is shown in Figure 1b. The convergence velocity is calculated as the horizontal motion between the subducting plate and upper plate (i.e., motion between Marker 1 and Marker 2 in Figure 1a). The trench retreat is calculated as  $u_T = dx_T/dt$ , where  $x_T$  is the trench position and  $t$  is time, starting from an initial trench position  $x_T = 0$ . In all simulations, the trench is retreating, specific to ocean-ocean subduction.

The radius of curvature is one of the parameters that requires more careful inspection. The radius of curvature is generally  
185 calculated from a circular fit to the available earthquake data, which for natural subduction systems can be noisy, incomplete or subjective (Buffett and Heuret, 2011; Lallemand et al., 2005; Wu et al., 2008). For example, some analyses fit earthquake data for the first 100 km or less because the plate interface is clearly marked at high resolution. However, there are limits on the length of the arc of a circle that can reasonably fit a unique circle (i.e., see discussion in Val and Willenbring (2020)). For this reason, slab dip angle is often used as a metric for slab orientation. However, radius of curvature is more appropriate to  
190 describe slab deformation with depth, while slab dip represents only the tangent to curvature close to the surface.

We calculated the radius of curvature after Petersen et al. (2017) (Figure 1c), in the following way: we extract the upper surface of the core of the slab (black) and fit a circle to an arc defined by the inflection point where the plate starts bending (red point below A), and the point on the surface corresponding to 150 km depth (red point next to C). The slab core is the

most appropriate feature for the fitting algorithm to calculate the radius of curvature of the slab, as its strength controls the bending of the slab. The upper layers (weak crust and sediments) may deform strongly during subduction and introduce noise into the circle-fitting algorithm. This algorithm remains robust throughout the evolution of a simulation (see movies in data repository). Therefore, our calculations of the radius of curvature are approximately 30 km less than total radius of curvature which includes the crust and sediments.

Sediments reaching the trench may either subduct into the mantle or accumulate into an accretionary wedge. We quantify whether a margin is tectonic erosion (TE) or an accretionary wedge (AW) by calculating two diagnostics for the accretionary wedge at the trench: the angle ( $\alpha_{\text{wedge}}$ ), and the width ( $W_{\text{wedge}}$ ) (Figure 1c). These parameters are not equivalent to the ones calculated in the taper-wedge theory (Dahlen, 1984; Dahlen et al., 1984), which are more difficult to extract as current numerical resolution is too coarse (i.e., the surface topography variations in the wedge are too small).

The algorithm to calculate both the wedge angle and width is the following: 1) isolate the sediment markers (grey material in Figure 1c), 2) determine wedge points (A,B,C): point A is the inflection point of the slab at the surface, point B is the end point to the right of the surface of sediments, and point C is at the base of the upper plate. We then connect ABC into a triangle, and calculate  $\alpha_{\text{wedge}} = \angle ACB$ , and  $W_{\text{wedge}} = \overline{AB}$ . Supplementary material shows that the algorithm works well in the majority of cases (Supplementary material, movies in data repository). It is important to note that the wedge angle is non-zero even in tectonic erosion margins, as there is a finite thickness of the sediments (i.e., points A, B will not overlap).

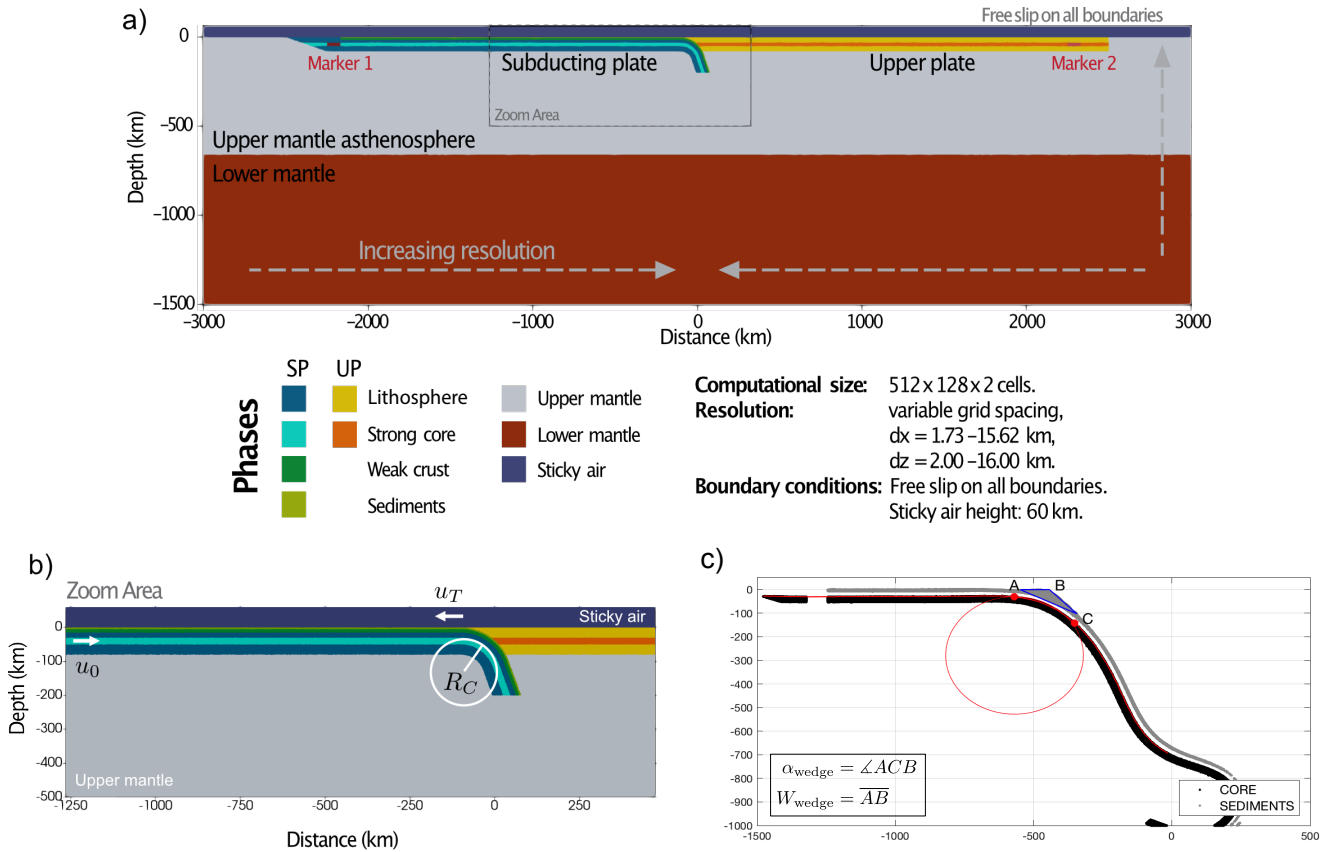
The effect of sediments on topography is also investigated. Lamb and Davis (2003) suggested that sediment-starved subduction may have been responsible for high topography in the Andes. They argue that tectonic erosion favours more coupling with the upper plate, while accretionary wedges favour decoupling, thus lower topographic amplitude. We investigate these hypotheses by extracting three diagnostics related to topography: trench depth ( $h_{\text{trench}}$ ), maximum topographic amplitude in the upper plate ( $h_{\text{max}}$ ), and mean topographic amplitude in the upper plate ( $h_{\text{mean}}$ ). The choice of last two is motivated by the study of Pusok and Kaus (2015), which shows that the two parameters can describe a number of topographic expressions for convergent margins.

### 3 Results

The 2-D numerical experiments below aim to understand what causes convergent margins to either accrete material delivered by the subducting plate or, alternatively, to subduct the trench sediment pile and even erode the basement of the overriding plate. In the first part of results, we describe end-member models of margin styles (accretionary or tectonic erosion) and the corresponding reference model results. In the second part, we analyse results from all numerical models and investigate parameter correlations using the diagnostics presented above.

#### 3.1 Margin styles and reference models

The outcome of each simulation is classified into three regimes: tectonic erosion (TE), low-angle accretionary wedge (low-AW), high-angle accretionary wedge (high-AW) (Figure 2, Table 1). The end-member division was done both qualitatively



**Figure 1.** Model setup. a) The model consists of an oceanic plate (SP) subducting beneath another oceanic upper plate (UP). Both plates are 2500 km long and have an upper mantle lithosphere component with a 20 km strong core. The subducting plate contains a weak crust and sediments (combined 15 km, weak crust - dark green, and sediments - light green lithologies in panel b). Parameter values are listed in Table 1 and are relative to the reference density ( $\rho_0 = 3300 \text{ kg/m}^3$ ) and viscosity ( $\eta_0 = 2.8 \times 10^{20} \text{ Pa}\cdot\text{s}$ ) of the mantle asthenosphere. The red markers (Marker 1 and Marker 2) are located in the strong cores of the subducting and upper plate to track the convergence of the plates. The domain has free-slip boundary conditions on all sides, and a 60 km "sticky-air" layer on top of the plates. Subduction is initiated by prescribing an initial slab depth of 200 km, and an initial radius of curvature of 150 km. The dynamics is entirely controlled by the negative buoyancy (slab-pull) of the subducting slab. b) Zoom area on the subduction interface and illustrating the convergence rate ( $u_0$ ), trench rate ( $u_T$ ), and the radius of curvature ( $R_C$ ). c) Phase markers used to calculate the radius of curvature (core markers - black), and wedge properties such as angle and width (sediment markers - grey).



No.	Sim Name	Margin Type	Sediment thickness (km)	Sediment viscosity (Pa.s)	Upper plate thickness (km)	$u_0$ (cm/yr)	$R_C$ (km)	$u_T$ (cm/yr)	$\alpha_{\text{wedge}}$ ( $^\circ$ )	$W_{\text{wedge}}$ (km)	$h_{\text{trench}}$ (km)	$h_{\text{max}}$ (km)	$h_{\text{mean}}$ (km)
1	SubdSed01	low-AW	5	$0.01 \times \eta_0$	80	6.727	514.6484	-2.9119	23.4255	171.4541	-2.8382	0.74299	-0.78694
2	SubdSed02	low-AW	5	$0.1 \times \eta_0$	80	6.7718	408.8925	-2.7225	20.4801	124.9362	-3.3691	0.8073	-0.76812
3	SubdSed03	TE	5	$\eta_0$	80	5.4276	319.7157	-1.525	17.3787	94.3482	-4.1917	0.80901	-0.76912
4	SubdSed04	high-u-AW	10	$0.01 \times \eta_0$	80	6.131	708.9844	-1.7564	31.2584	268.5403	-4.67	0.92618	-0.79155
5	SubdSed05	high-AW	10	$0.1 \times \eta_0$	80	9.332	519.104	-4.376	28.157	205.5639	-2.6026	0.7522	-0.76238
6	SubdSed06	TE	10	$\eta_0$	80	5.4405	308.8108	-1.5226	21.244	128.9402	-4.2016	0.79123	-0.7723
7	SubdSed01_50	high-AW	5	$0.01 \times \eta_0$	50	6.5998	708.3037	-2.2749	26.8088	209.3449	-3.9974	0.62997	-0.4324
8	SubdSed02_50	high-AW	5	$0.1 \times \eta_0$	50	7.4274	594.9156	-3.233	25.5747	149.9525	-3.8102	0.64673	-0.43965
9	SubdSed03_50	low-AW	5	$\eta_0$	50	5.9242	397.1737	-1.9264	20.6625	105.5102	-3.7775	0.59663	-0.44411
10	SubdSed04_50	high-AW	10	$0.01 \times \eta_0$	50	13.4398	617.7619	-7.2085	25.2127	208.2968	-1.5817	0.51963	-0.42698
11	SubdSed05_50	high-AW	10	$0.1 \times \eta_0$	50	10.3734	629.6387	-5.0662	27.2371	204.1125	-2.6618	0.59171	-0.44608
12	SubdSed06_50	low-AW	10	$\eta_0$	50	5.918	392.3118	-1.9238	24.0504	148.7363	-3.7849	0.64602	-0.43274
13	SubdSed01_100	low-AW	5	$0.01 \times \eta_0$	100	6.8774	467.8223	-2.8441	21.8272	155.5387	-2.8897	1.0432	-1.0002
14	SubdSed02_100	TE	5	$0.1 \times \eta_0$	100	6.3013	351.5625	-2.4277	16.6464	103.3699	-3.5028	1.0929	-0.99431
15	SubdSed03_100	TE	5	$\eta_0$	100	5.0905	294.5964	-1.2357	14.8858	83.386	-4.1857	1.1723	-0.99182
16	SubdSed04_100	u-AW	10	$0.01 \times \eta_0$	100	7.2866	596.5169	-2.3679	30.1553	247.4326	-4.0447	1.1466	-1.0876
17	SubdSed05_100	high-AW	10	$0.1 \times \eta_0$	100	8.2337	494.5845	-3.6275	27.0269	198.0693	-2.7064	0.97735	-0.97861
18	SubdSed06_100	TE	10	$\eta_0$	100	5.0838	284.3424	-1.2265	18.5907	115.977	-4.1134	1.1568	-0.9942
19	SubdSed01_150	TE/low-AW	5	$0.01 \times \eta_0$	150	5.5265	273.763	-1.4631	13.9752	71.4896	-2.4593	1.9874	-1.512
20	SubdSed02_150	TE	5	$0.1 \times \eta_0$	150	5.067	279.248	-1.018	11.3664	67.3686	-3.5097	2.1631	-1.6142
21	SubdSed03_150	TE	5	$\eta_0$	150	4.3101	260.0369	-0.60011	11.9206	66.8439	-3.8979	2.2071	-1.631
22	SubdSed04_150	u-AW	10	$0.01 \times \eta_0$	150	6.1352	349.5206	-2.3916	20.9114	136.2829	-2.5751	1.8707	-1.551
23	SubdSed05_150	low-AW	10	$0.1 \times \eta_0$	150	6.5457	333.933	-2.1902	20.4162	138.7542	-2.7802	1.8619	-1.5804
24	SubdSed06_150	TE	10	$\eta_0$	150	4.2999	247.9384	-0.59526	15.2628	93.0433	-3.8739	2.2166	-1.6283

**Table 1.** Simulations performed in this study. Margin type: TE - tectonic erosion, low-AW - low-angle accretionary margin, high-AW - high-angle accretionary margin, u-AW - unstable accretionary margin. Reference viscosity is  $\eta_0 = 2.8 \times 10^{20}$  Pa.s (upper mantle viscosity). Diagnostics (mean values):  $u_0$  - convergence rate,  $R_C$  - radius of curvature,  $u_T$  - trench retreat,  $\alpha_{\text{wedge}}$  - wedge angle,  $W_{\text{wedge}}$  - wedge width,  $h_{\text{trench}}$  - trench depth,  $h_{\text{max}}$  - maximum topographic amplitude in the upper plate,  $h_{\text{mean}}$  - mean topographic amplitude in the upper plate. Evolution of each diagnostic parameter is shown in Supplementary material.

	Slab (Subduction)	Sediment (Plate interface)	Upper plate (Topography)
Fixed Input Parameters	$\rho_{SP}, \eta_{SP}, h_{SP}$	$\rho_{sed}$	$\rho_{UP}, \eta_{UP}$
Variable Input Parameters	—	$\eta_{sed}, h_{sed}$	$h_{UP}$
Diagnostics	$u_0, R_C, u_T$	$\alpha_{wedge}, W_{wedge}$	$h_{trench}, h_{max}, h_{mean}$

**Table 2.** Input and diagnostics parameters. Input parameters represent parameters that are prescribed at the beginning of each simulation and stay the same throughout the evolution, while diagnostics are parameters that are the result of the dynamics of the system, and are calculated during model evolution. Input parameters:  $\rho$  - density,  $\eta$  - viscosity,  $h$  - thickness, and subscripts represent: SP - subducting plate, sed - sediments, UP - upper plate. Diagnostics:  $u_0$  - convergence velocity between subducting plate and upper plate (horizontal motion between Marker 1 and Marker 3 in Figure 1),  $u_T$  - trench motion,  $R_C$  - radius of curvature,  $\alpha_{wedge}$  - angle of accretionary wedge,  $W_{wedge}$  - width of accretionary wedge,  $h_{trench}$  - trench depth,  $h_{max}$  - maximum topography in the upper plate,  $h_{mean}$  - mean topography in the upper plate.

(i.e., formation of the accretionary wedge as seen in Figure 2, left column) and quantitatively using the evolution of diagnostics parameters (transient versus steady state) which is shown in Figure 3. We explain the individual classification below. Details of each model evolution are included in the Supplementary material, as they have been extensively investigated and described in previous studies (i.e., adjustment of the model to initial conditions and development of slab curvature, formation of accretionary  
230 wedge, interaction of slab with the lower mantle).

We focus here on characterizing overall model outcomes from incipient subduction to slab consumption. The time taken for subduction to consume the slab varies for each model because the velocity field (i.e., subduction velocity, convergence rate) of the system is controlled by interface processes, in turn controlled by material parameters (see dependence of end time and average time step of a simulations on convergence rate in Supplementary material, Figure S7). For this reason, we use  
235 a characteristic time ( $t_{char} = t/t_{final}$ ) to highlight the entire model evolution in several figures (i.e., Fig. 3). Initial model time corresponds to  $t_{char} = 0$ , while final time  $t_{final}$ , corresponding to  $\sim 2500$  km slab consumption, becomes  $t_{char} = 1$ . Every model simulation has an initial adjustment period of  $t_{char} \sim 0.1$  in which the subduction system acquires a natural curvature, and a final stage  $t_{char} \sim 0.9$  in which the slab is consumed (grey intervals in Figure 3). These initial and final condition stages are excluded in our calculation of diagnostics parameters.

240 For the rest of the study, we adopt a color code for each regime: purple for tectonic erosion, yellow for low-angle accretionary wedge, and orange for high-angle accretionary wedge.

**Tectonic erosion margin (TE).** Figure 2a shows a typical simulation outcome of a tectonic erosion margin (model SubdSed03, with thin cover  $h_{sed} = 5$  km and high viscosity of sediments  $\eta_{sed} = \eta_0$ ). Evolution of diagnostics parameters is shown in Figure 3, and additional model evolution snapshots are shown in Supplementary material, Figure S3.

245 We classified simulations as tectonic erosion when the evolution of the radius of curvature, convergence rate, wedge angle and width, and trench rate are in steady-state (constant) throughout the simulation (purple lines, Figure 3 and Supplementary material Figures S13-S15). The means (dotted purple lines) follow closely the evolution diagnostics.

The low convergence velocity and trench retreat rate maintained throughout the simulation (Figure 3b,e) suggest a high degree of coupling between the subducting slab and upper plate. Figure 2a-B shows that the motion between subducting and upper plate in tectonic erosion margin is accommodated in the middle of the sediment layer, a region of high strain-rates. 250 Instead, very low strain-rates just below the slab suggest a rigid-body rotation of the slab in order to maintain a constant radius of curvature. The radius of curvature remains small as seen in Figure 2a-C, with a steeply-dipping slab.

Entrainment of sediments within upper plate material at the interface (Figure 2a-A) is indicative of some erosion of the upper plate. All simulations with high viscosity sediments show this behaviour (Supplementary material, Figures S13-S15, cases with 255  $\eta_{sed} = \eta_0$ ), which could be regarded as having a stronger interface (i.e., a more mafic cover and/or lack of weak unconsolidated sediments).

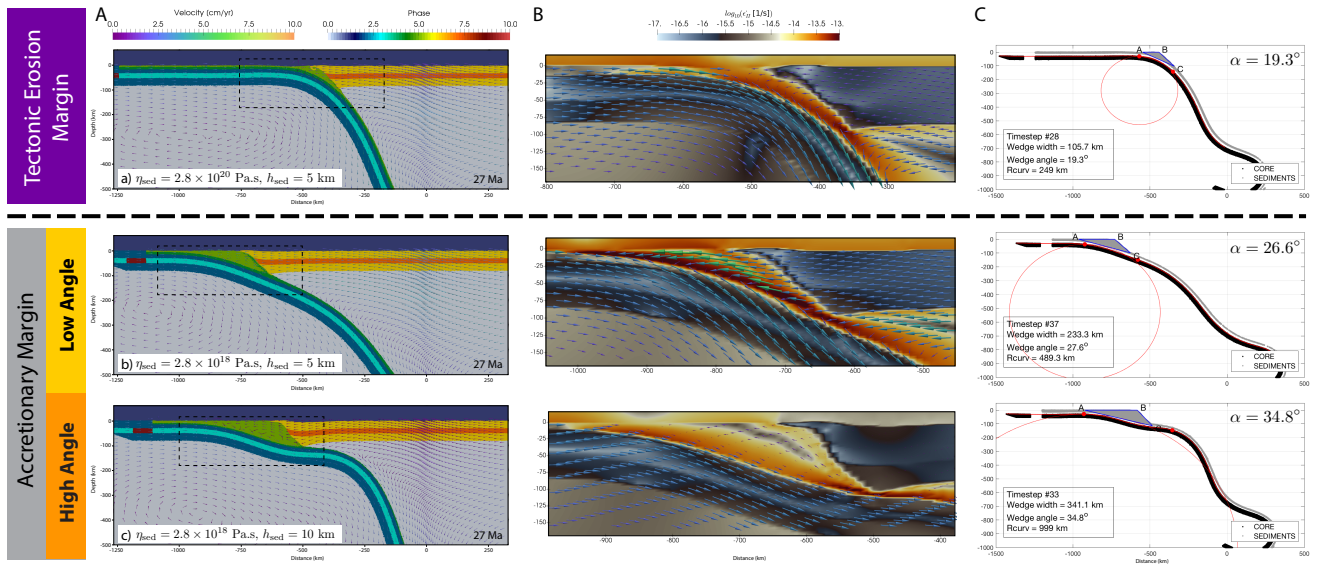
Topographic signals in tectonic erosion (trench depth, maximum and minimum topography in the upper plate) show more variability in Figure 3, which will be discussed later.

**Accretionary Wedge margin (AW).** When the viscous coupling is reduced, sediments are scrapped off the subducting slab 260 to form an accretionary wedge (Figure 2b-c). We identify two types of accretionary margins: low- and high-angle accretionary wedges, primarily controlled by the thickness of sediments in cases of low-viscosity sediments. Evolution of diagnostics parameters is shown in Figure 3 (yellow and orange lines), and additional model evolution snapshots are shown in Supplementary material, Figures S4-S5.

The distinction between the two cases comes from the behaviour of the slab: low-angle accretionary margins have increasing 265 radii of curvature, wedge properties, but fairly constant convergence rate (Figure 3b), while high-angle accretionary margins result in flat slab subduction with large radii of curvature and irregular behavior of the convergence rate. In high-angle accretionary margin simulations, in a first stage of evolution, sediments accumulation lubricates the interface and promotes fast convergence rates, but once the wedge reaches a critical size and slab curvature is too large (i.e., subduction needs to accommodate horizontal slab motion), subduction rate is inhibited. When this stage is reached, plate convergence may happen at slower 270 rates than in tectonic erosion simulations.

Accretionary margins models are favoured by lower viscosities ( $\eta_{sed} = 0.01 \times \eta_0$ ). By increasing the thickness of the sediments, more sediment is available to create a thicker wedge (high angle and width). The larger the wedge angle, the larger the radius of curvature, suggesting that wedge geometry has a control on slab bending (Figure 3a-d).

The interface dynamics is also different with increasing availability of weak sediments. The velocity field at the subduction 275 interface suggests internal counter-clockwise flow inside accretionary wedges, detached from both corner flows in the mantle (Figure 2B). Plate motion is also accommodated at the base of the sediment wedge, indicated by high strain-rates. Despite the unrealistically large geometry of the wedge, Figure 2c-B shows that the accretionary wedge can impede motion within the upper plate. However, we are not concerned here with further details of internal wedge dynamics compared to numerous previous studies because we also lack the numerical resolution required (i.e., Ruh (2017); Menant et al. (2020)).

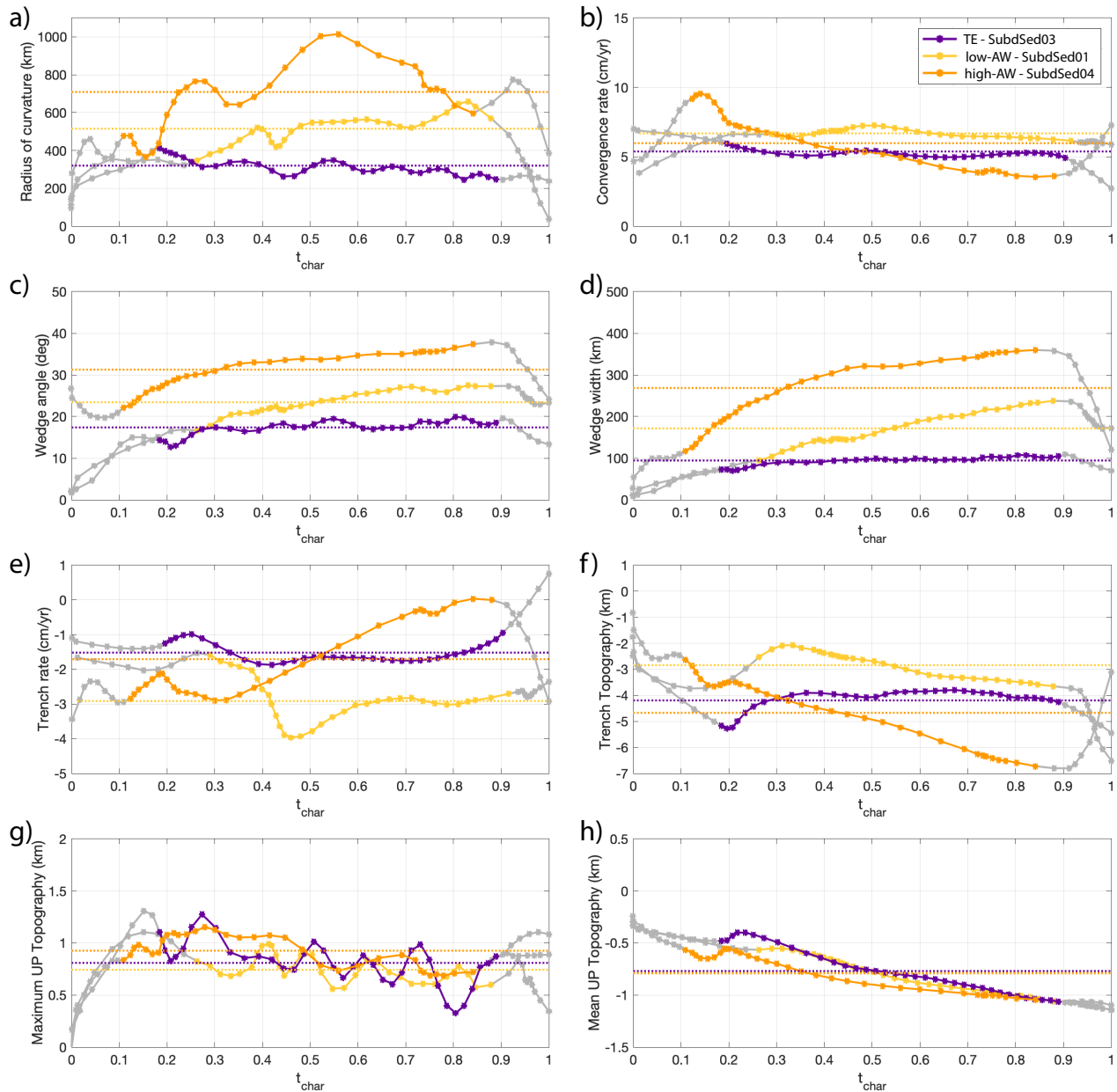


**Figure 2.** Margin style end-member results. a) Tectonic erosion margin (TE, reference model SubdSed03), b) Low-angle accretionary margin (AW, reference model SubdSed01), c) High-angle accretionary margin (AW, reference model SubdSed04). A. Left column shows model snapshots at 27 Ma, B. Middle column shows second invariant of strain rate,  $\dot{\epsilon}_{II} = (\dot{\epsilon}_{ij}\dot{\epsilon}_{ij})^{1/2}$  and velocity field as arrows in an enlarged area of the subduction interface, and C. Right column shows core and sediment markers, with calculated radii of curvature, wedge angles and widths. Tectonic erosion margin shows low-angle wedges and small radius of curvature. When an accretionary wedge forms, in cases of lower sediment viscosity, the wedge angle and width increases over time, together with the radius of curvature. In high-angle accretionary margins (c), episodes of flat slab may occur, and strongly influence plate bending. Time evolution of diagnostic parameters for the reference cases are shown in Figure 3.

280 A third margin style of a highly-unstable accretionary wedge is shown in the Supplementary material (e.g., results of SubdSed04\_100 in Figures S6, S13-S15). In these cases, the accretionary wedge reaches a critical angle, and instead of moving laterally, material is being expelled down the subduction channel. The wedge will continue to deform and grow again until it reaches a new critical angle. This unstable mode occurs in simulations with thick upper plate thickness, which acts as a deformable backstop, in combination with accumulating weak and thick sediments. This margin style is a consequence of  
 285 the density model chosen here, in which sediments have the same density as the rest of the lithosphere. Clearly, this is an overestimated effect.

### 3.2 Major Impacts on Subduction Dynamics

Although material within the weak layer at the plate interface (i.e., between the subducting and upper plates) is a volumetrically insignificant component of the larger plate-mantle coupled system, we observe this small feature can exert a profound influence  
 290 on the emergent regional-scale subduction dynamics. Figures S10 and S11 show models with the strongest sediment layers



**Figure 3.** Evolution of diagnostic parameters for end-member reference models: Tectonic erosion - SubdSed03 (purple), low-angle accretionary margin - SubdSed01 (yellow), high-angle accretionary margin - SubdSed04 (orange). a) Radius of curvature, b) convergence rate, c) wedge angle, d) wedge width, e) trench rate, with positive values indicating trench advance and negative values indicating trench retreat, f) trench depth, g) Maximum topographic amplitude in the upper plate ( $h_{max}$ ), h) Mean topographic amplitude in the upper plate ( $h_{mean}$ ). Continuous lines represent model data, while dotted lines represent the mean over entire simulation time. Characteristic time  $t_{char}$  is defined in the main text. Grey portions represent initial and final conditions, corresponding to the system forming its natural slab curvature, and to the last stage of slab consumption. The two stages are excluded from calculating the means.

( $\eta_{sed} = \eta_0$ ) have sub-vertical slab morphologies, smaller values of  $R_C$ , and slower trench retreat rates than similar models with weaker sediment layers ( $\eta_{sed} = 0.01 \times \eta_0$ ). Stronger sediment layers also stabilize the subduction system as seen in Figure S12, in which much larger variations in  $R_C$  occur for models in column A than column C, where values remain approximately constant. The steady-state values of  $R_C$  can vary by more than a factor of two due to the viscosity of the sediments, with model  
295 SubdSed06\_50 having 400 km while models SubdSed04\_50 and SubdSed05\_50 both evolve such that  $R_C$  exceeds 800 km. For these models, a similar increase of more than a factor of 2 can also be observed in convergence rate (Figure S14) and trench motion (Figure S15).

It is not just the low strength of the sediment layer that influences the system, but also the thickness of the upper plate. The dynamics of the plate interface depend on the total length of contact area between the two plates as well as the thickness  
300 and strength of the material between them (Beall et al., 2021). Comparisons of models that are otherwise identical except for having different upper plate thicknesses (Figure S11, columns a-A and b-A) exhibit more shallow-angled and variable slab morphologies for those models with thinner upper plates.

### 3.3 Parameter analysis

In this section, we investigate correlations between the means of the diagnostic parameters for all simulations (Figures 4 and  
305 5). The diagnostics parameters can then be compared to similar parameters observed in the global subduction system.

Figure 3 shows that the evolution means characterize the margin style of each simulation. For each diagnostic parameter, we calculate the mean value (Figure 3, dotted lines) and the variability during evolution (minimum and maximum values). The means in TE models remain close to the evolution curves (i.e., steady-state with the mean close to the min/max values). In AW models, the means differ significantly from the evolution curves. Therefore, the min/max values during time evolution reflect  
310 this larger variability (grey bars in Fig. 4 and 5). Diagnostic means are given in Table 1 and evolution curves are shown in Supplementary material (Figures S13-S18).

Figure 4 shows correlations between subduction and sediment diagnostic parameters, while Figure 5 shows correlations between sediment and topography diagnostic parameters. In both figures, each colored point represents the mean value in a simulation, and the grey bars represent variability intervals (min/max values). The colorscale represents the margin type,  
315 identified as in the previous section. TE simulations (purple) immediately have smaller variability bars, also emphasizing steady-state. On the other hand, AW models have larger variability bars (largest for high-angle AW) suggesting transient evolution for that diagnostic.

We find strong correlations between sediment parameters and subduction parameters in 2-D numerical models (Figure 4). Convergence rate correlates well with radius of curvature, wedge angle and width (panels a-c). TE models have low conver-  
320 gence rates, radii of curvature, but also small accretionary wedge properties (angle and width). As seen in previous section, with increasing sediment availability (thickness) and/or decreasing sediment viscosity, wedges form more readily. Figure 4 shows that AW models register higher convergence rate, higher radii of curvature and larger wedge properties. This indicates that sediments lubricate the subduction interface, promoting faster convergence rates. The rate of trench motion is also influenced by the dynamics of subducting sediments, with sediment-rich AW trenches retreating faster (panel d).

325 Figure 5 suggests weaker correlations between sediment and topography parameters. In panels b-j, TE models tend to  
produce higher maximum topography and deeper trench depths (i.e., extreme amplitudes), while AW models produce lower  
extreme signals, but higher mean values. The weaker correlation of sediments diagnostics with topographic signals is likely due  
to the nature of the upper plate (i.e., oceanic/continental, free/attached plate to side walls). In our set of simulations with ocean-  
ocean subduction and unattached upper plate, stresses accumulated at the plate interface are accommodated in trench motion,  
330 instead of upper plate deformation. Topographic signals for simulations with a fixed upper plate are shown in Figure S9, which  
are higher and more distinct for end-member models, as subduction is now accommodated more in deformation of the upper  
plate, and less in trench motion. Topography also builds faster in continental lithosphere compared to oceanic lithosphere due  
to less resistance against gravity (Pusok and Kaus, 2015). Investigation of the effect of sediments on topography signals with a  
continental upper plate is reserved for a future study.

335 To summarize, tectonic erosion models suggest a stronger coupling at the plate interface, yielding lower radii of curvature,  
slower convergence rates and higher topographic signatures. TE margins also retreat slower compared to AW margins. These  
correlations and simulation snapshots suggest that dynamics of the wedge controls the bending of the slab and the radius of the  
curvature.

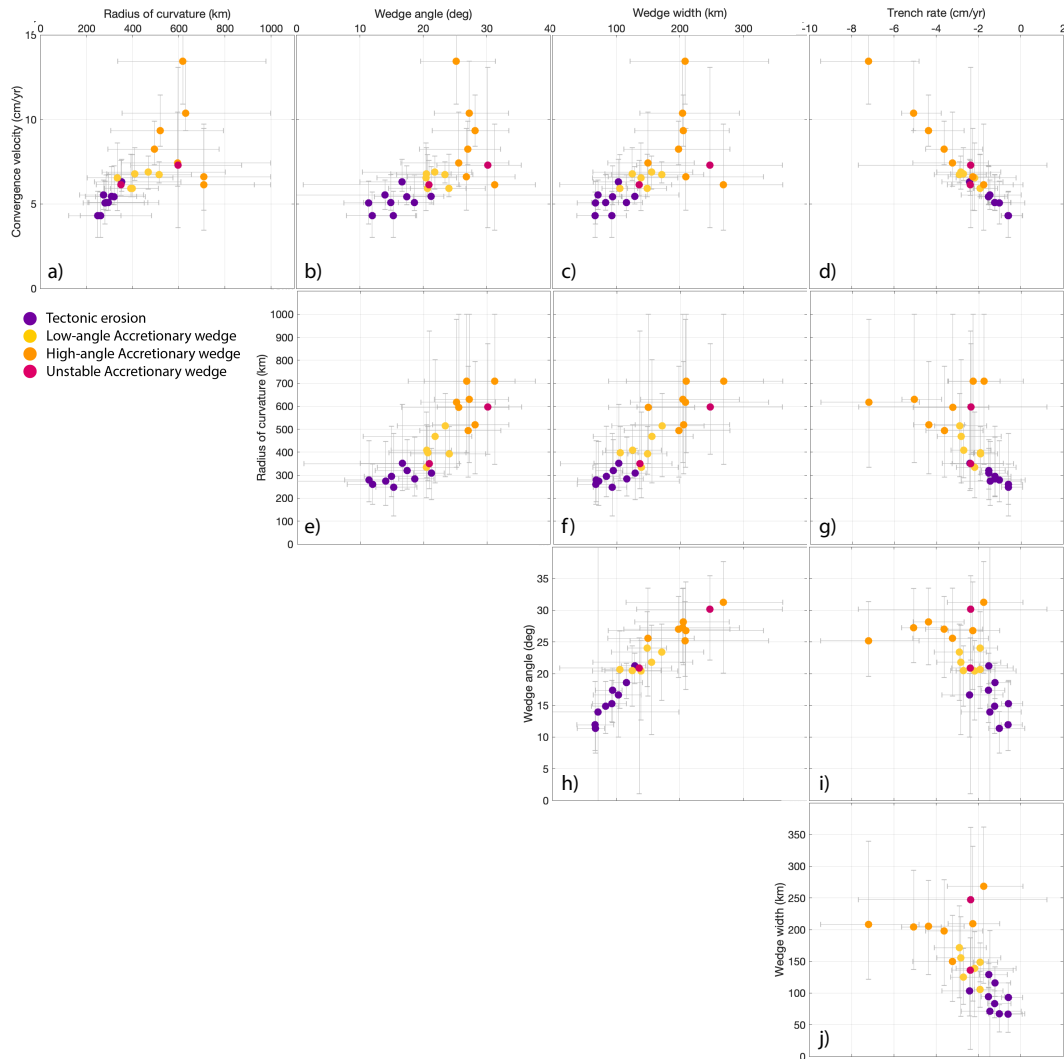
## 4 Discussion

340 Our results show three modes of subduction interface: tectonic erosion margin, low-angle accretionary wedge margin, and  
high-angle accretionary wedge margin. We obtain a diverse response in subduction geometry to just a few varied input pa-  
rameters: sediment viscosity and thickness, and upper plate thickness. The focus of our analysis is solely on the effect of  
low-viscosity sediments on large-scale subduction dynamics. Other parameters have also been shown to be important (i.e.,  
density of sediments, age of slab, thermal structure, upper plate structure) that will be discussed below.

345 The viscosity of sediments represents the critical parameter, and thickness as a secondary parameter in our simulations. High-  
viscosity sediments lead to tectonic erosion margin, while low-viscosity sediments lead to accretionary wedge margins (Table  
1). Sediment thickness controls the availability of sediments to be accumulated in accretionary wedges. A detailed analysis of  
Figures S10-S18 shows that the thickness of the upper plate plays an important role in determining the subduction interface  
length and the depth at which sediments can be locked into an accretionary wedge. For the same thickness of sediments, when  
350 the upper plate is thinner, the accretionary wedge is volumetrically smaller, but wider and with high wedge angle, while for a  
thick upper plate, sediments are distributed across the entire interface into a thinner layer with a small wedge angle, leading  
to tectonic erosion (i.e., compare results of SubdSed02\_50 for 50 km thick upper plate, and SubdSed02\_150 for 150 km thick  
upper plate in Figure S11).

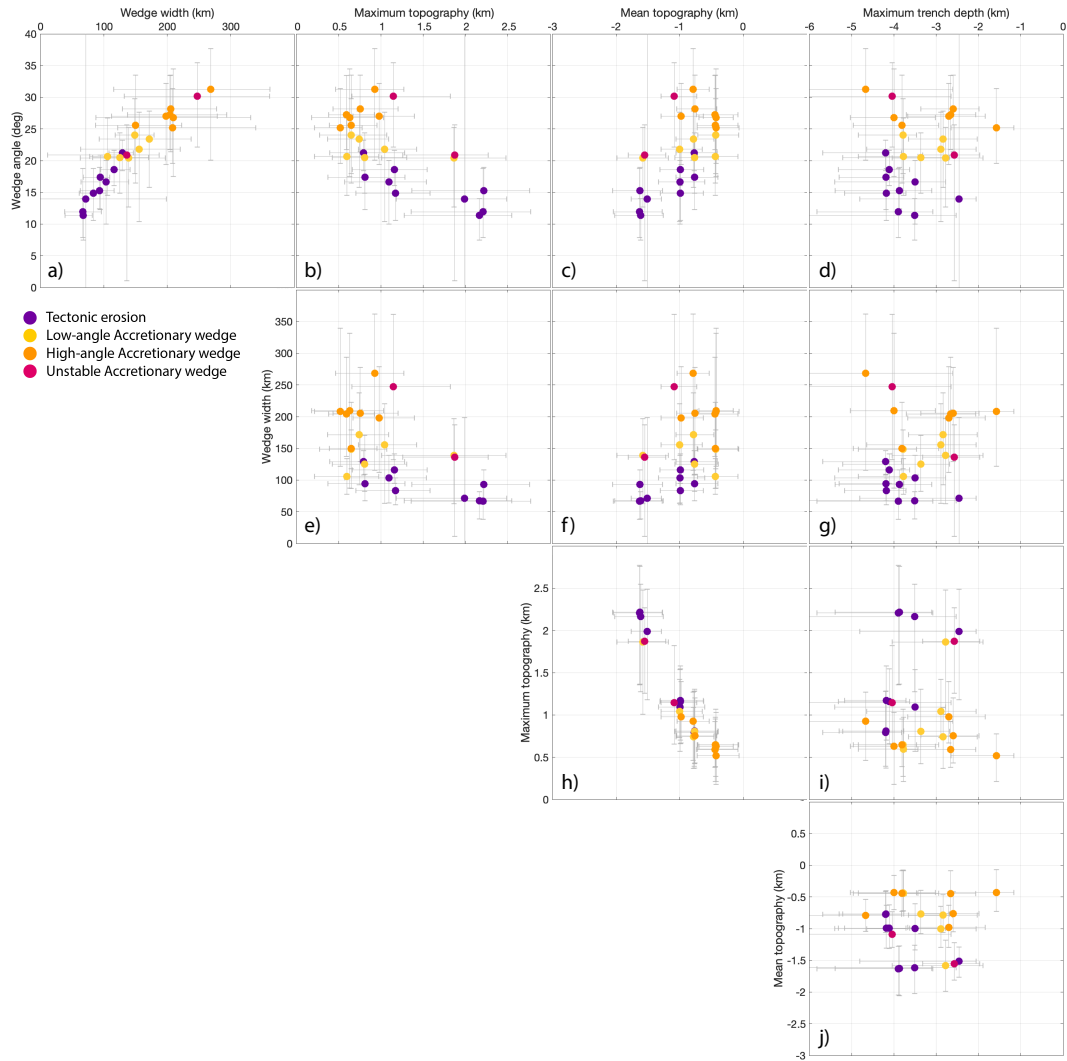
### 4.1 Parameter correlation and observations

355 A common approach to constrain the effect of different parameters that control subduction dynamics has been done considering  
the statistical analysis of present-day subduction zones (Supplementary material, Figure S2) (Clift and Vannucchi, 2004; Lalle-



**Figure 4.** Subduction versus sediment diagnostic parameters. Each point represents the evolution mean of the given parameter in a simulation, and the grey bars represent the variability intervals (see Figure 3 on how the mean is calculated). Initial and final stages of the evolution are removed from calculations. Tectonic erosion models (purple) have lower variability in diagnostics than accretionary wedge models (yellow/orange). a) Radius of curvature and convergence velocity. Small radii of curvature are correlated with small convergence velocities. b-c) Small wedge properties (angle and width) are also correlated with low convergence rates and radii of curvature. We observe clear correlations between margin types and diagnostics. Tectonic erosion margins have low convergence velocities, small radii of curvature, and wedge properties. Accretionary wedge models have higher radii of curvature, faster convergence velocities and larger wedge properties.





**Figure 5.** Sediment versus topography diagnostic parameters. Each point represents the evolution mean of the given parameter in a simulation, and the grey bars represent the variability intervals (see Figures 3 and 4). Tectonic erosion models yield smaller trench depths and mean topography, and higher maximum topography in the upper plate. The weaker correlations are due to the nature of the upper plate, which is considered oceanic and unattached to the right boundary, so any topographic signal is smaller than it would be for continental upper plate lithosphere. Topographic signals for simulations with fixed upper plate are shown in Figure S9, which are higher and more distinct for end-member models.

mand et al., 2005; Wu et al., 2008; De Franco et al., 2008; Heuret et al., 2012). Our diagnostic parameter analysis in Section 3.3 attempts to create a bridge between these studies and numerical models, as diagnostics from numerical models are often not comparable with those from statistical analyses. There remain a number of fundamental differences between our results and  
360 these studies. Present-day subduction parameter correlation represent only current snapshots and do not always account for the evolution stage of a given subduction system. Our analysis considers the entire evolution of a subduction system and evolution averages. Moreover, natural subduction zones have variable sediment influxes, changing boundary conditions and multi-phase physics. On the other hand, the numerical model setup is ideal and simplified, which is further discussed below.

Despite these differences, we succeed in obtaining subduction margins that accrete sediments, and ones that are erosive.  
365 Our results agree with findings in Clift and Vannucchi (2004). Accretionary margins form in simulations with thicker and weaker sediment covers entering the trench, while tectonic erosion margins form in simulations with less sediment cover (high viscosity sediments are representative of a stronger crust and mantle component at the slab surface). Replotted data from Wu et al. (2008) and using the margin classification from Clift and Vannucchi (2004) (Supplementary material, Figure S2), shows that accretionary margins also tend to have larger radii of curvature, which is consistent with our numerical results. A recent  
370 study obtained similar correlations between margin style and convergence rate and radii of curvature (Brizzi et al., 2021).

Our findings using slab curvature are consistent with observations on slab dip. Diversity of subduction zones is generally investigated through the perspective of slab orientation (i.e., Beall et al. (2021); Riel et al. (2018)). Many analogue and numerical studies used slab dip as the preferred diagnostic for slab orientation instead of slab curvature (i.e., Heuret et al. (2007), Section 2.2). They suggest that slab dip depends on properties of the subducting plate, mantle, overriding plate, and the coupling between the subducting and overriding plate (e.g., Bellahsen et al. (2005); Heuret et al. (2007); Billen and Hirth (2007);  
375 Schellart et al. (2007); Funicello et al. (2008); Babeyko and Sobolev (2008); Duarte et al. (2013); Riel et al. (2018)). We find that coupling at the plate interface due to sediment influxes can strongly influence bending of the slab (radius of curvature). TE models have small  $R_C$  (large dip), while AW models have large  $R_C$  (small dip). The nature of upper plate (oceanic or continental), however, will further influence these correlations by changing the load on subduction interface, which should be  
380 investigated further.

Topographic signals have been incorporated less in both statistical and numerical studies of subduction zones, despite the fact that topography is a direct and easily acquirable observable (i.e., Pusok and Kaus (2015); Riel et al. (2018)). In higher resolution models, Menant et al. (2020) found that deep accretion processes influence forearc topography over time, but did not compare the effect in both accretionary and erosion margins. In our ocean-ocean subduction, we see a lesser control of  
385 sediments on topographic diagnostics, with tectonic erosion margins yielding deeper trenches. However, the effect increases when the upper plate is attached to the right wall (Supplementary material, Fig. S9). This is because the subduction interface stresses are transferred in the upper plate (i.e., topography build-up) rather than accommodated in trench retreat.

## 4.2 Convergence rate and margin type

Results from numerical models suggest strong correlations between margin type and convergence rate (Figure 4). Convergence  
390 rates in AW models are faster than in tectonic erosion models, as sediments help lubricate the interface and reduce coupling

between subducting and upper plate. A special case could be the high-angle AW model results (Figure 3b), where the convergence rate is faster in the growing-stage of the accretionary wedge, but once it reaches a critical value, the convergence rate becomes slower than in the case of tectonic erosion. Behr and Becker (2018) and Brizzi et al. (2021) obtain similar predictions, in which weaker (lower viscosity) sediments promote faster convergence rates. However, observations suggest an inverse correlation between convergence rate and margin type (see Supplementary material, Figure S2), with tectonic erosion margins subducting faster than accretionary wedge margins. Most likely, this inverse correlation between convergence rate and margin type suggests that we do not capture all complex processes happening at the subduction interface (von Huene et al., 2004).

The dominant mechanisms governing accretionary and tectonic erosion margins are different. The style of deformation within accretionary wedges is thin-skinned, that is, associated thrusts and folds are uncoupled from the underlying oceanic crust by a basal decollement with a large amount of displacement (Weiss et al., 2018; Angiboust et al., 2021). The dynamics of accretionary wedges can become more complex if one considers multiple types of sediments, fluid pressure and deformation mechanisms (Ruh, 2017; Menant et al., 2020). In accretionary margins, ocean sediments are available to lubricate the interface, while little sediments enter tectonic erosion margins. However, in tectonic erosion margins, the subducting plate may erode the basement of the upper plate, and create further debris along the plate interface (von Huene et al., 2004). In any given system, both processes may be occurring simultaneously, either in time and space or at the same time in different parts of the subduction zone (Clarke et al., 2018; Comte et al., 2019; Ducea and Chapman, 2018; Straub et al., 2020). We do not account for these detailed processes in our numerical models.

Moreover, it is unclear whether sediment influxes affect convergence rate, or convergence rate affects sediment accumulation at trench, or both. Accretion is generally favored by slow convergence ( $<7.6$  cm/yr) (Silver et al., 1985; von Huene and Scholl, 1991a; Clift and Vannucchi, 2004; Syracuse et al., 2010), while fast convergence favors larger volumes of sediment to be dragged down at the interface, thus lubricating the interface. Erosive margins tend to occur in regions where the convergence rate exceeds 6 cm/yr and where the sedimentary cover is thin (Straub et al., 2020). Testing this hypothesis requires different experimental setups: one in which convergence rate evolves dynamically due to sediments input, and second, in which the convergence rate is prescribed and sediments deform accordingly. Here, we tested the first scenario, where the convergence rate is the dynamic response of the system due to incoming sediments at the trench. We briefly tested the second scenario in Supplementary material, Figures S8-S9, with kinematic boundary conditions. Imposed convergence with a fixed upper plate reduces the amount of trench retreat, and the variation of diagnostic parameters is smaller among end-member cases, but the correlations remain valid. Menant et al. (2020) investigated the dynamics of accretionary margins as a function of convergence rate to find that more sediment and crustal material is subducted at higher convergence rates.

### 4.3 Sediment fluxes to trench and to depth

Sediment fluxes to depths below the lithosphere influence the amount of volatiles recycled into the mantle (Plank and Langmuir, 1998; van Keken et al., 2011; Plank and Manning, 2019). The type of margin may also affect how much sediment gets subducted into the mantle. It has been proposed that tectonic erosion margins can subduct higher percentages of sediment influx, however, large volumes of continental crust are subducted at both erosive and accretionary margins (von Huene and Scholl, 1991a; Clift

425 and Vannucchi, 2004). Clift and Vannucchi (2004) calculated that accretion is a relatively inefficient process for cleaning sediment off the oceanic basement and that 70% of the sediment column is likely subducted to great depths below the forearc.

Here, we calculate the volume fraction of sediments accreted and subducted in the mantle below the lithosphere in our reference models (Figure 6). We find that tectonic erosion margin subducts higher percentage of influx sediments than accretionary margins. The total percentage of sediments subducted right before the slab was consumed in tectonic erosion margin was  
430 60%, while in accretionary margins the percentage remains at  $\sim 10\%$ . However, considering that tectonic erosion margins have a smaller sediment cover (and influx), accretionary margins may in total subduct a larger volume of sediments. For example, both 50% of a 1 km column of influx sediments and 10% of a 5 km column of influx sediments give 500 m column of subducted sediments. We conclude that both tectonic erosion margins and accretionary wedges can subduct a high volume of sediments, but at different rates relative to influx material.

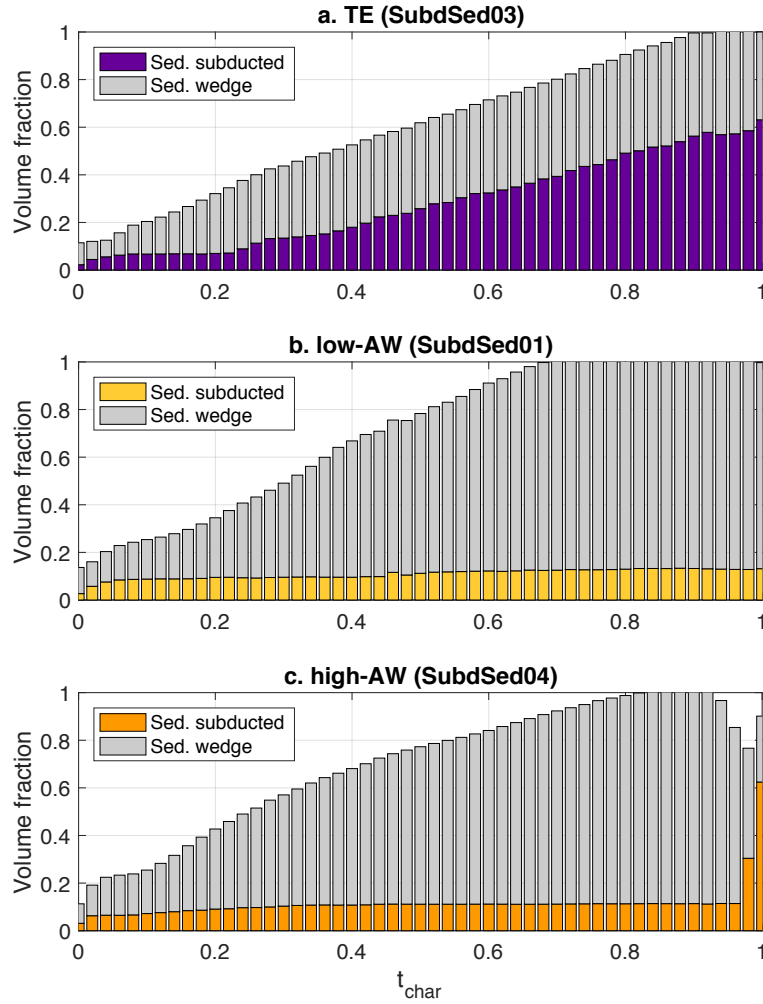
435 This is, however, a simplistic view of sediment subduction to depth. As sediments approach the trench, they can be accreted in the trench, subducted into the mantle, accreted structurally to the bottom of the upper plate after initial downward transport, and returned to the upper plate either via magmatism with partial melting of the downgoing sediment or some other form of diapirism or partially-molten upward transport (Gerya and Yuen, 2003; Currie et al., 2007; Tian et al., 2019). To model all these processes requires future development of multi-physics models that is beyond the scope of this study.

440 In this study, we considered a steady-state (constant) sediment flux to the trench. Sediment subduction is neither a steady-state nor a globally averaged process (Plank and Manning, 2019) and can have major implications for subduction dynamics. Modern oceanic sediments cover 70% of the planet's surface, but sediment distribution and lithology occur in drastically different proportions globally (Dutkiewicz et al., 2015, 2018). Moreover, the oceanic lithosphere is covered by various sediment types depending on the depth, proximity to continental margins, and interactions with the oceanic currents and biosphere.  
445 For example, an abundant carbonate cover is subducted at the Central American margin, while little sedimentary carbonate is subducted along the Tonga, Central Aleutian and Kuriles–Kamchatka trenches (Plank and Langmuir, 1998; Plank and Manning, 2019). The global sedimentary cover also varies in both space and geological time, with greatest volume in the geologically recent and decreasing exponentially with increasing age (Peters and Husson, 2017).

We expect the abundance and lithology of sediments at trenches to influence the mode of occurrence of margin styles in  
450 space and geologic time. The subduction interface structure and properties are sensitive to the composition of the material that is being subducted (Behr and Becker, 2018). Our results cover broadly the end-member scenarios. The degree to which variability in these influxes impacts long-term subduction dynamics remains debated (Cloos and Shreve, 1988; Duarte et al., 2015; Behr and Becker, 2018) and should be studied with further numerical modelling.

#### 4.4 Model limitations

455 In order to be able to address the points above, a number of model improvements are needed. The models shown here provide an initial experiment on the viscosity of sediments, where we considered constant sediment fluxes at the trench. Future work should explore the effect of other material parameters (i.e., density, lithology of sediments, as in Currie et al. (2007)), and active surface processes such as erosion and sedimentation for potential delivery of continental sediments to the trench. The



**Figure 6.** Evolution of total volume fraction of sediments accumulated in accretionary wedge and subducted to depth in the mantle for the three reference cases: a) Tectonic erosion (SubdSed03), b) low-angle Accretionary wedge (SubdSed01), c) high-angle Accretionary wedge (SubdSed04).

effect of sediments is overestimated in some of our model outcomes, especially in models with unrealistically large wedges.  
460 Thinner and lighter sediment covers require higher resolutions at the trench, which could also help connect results with those from high-resolution accretionary wedge models (Menant et al., 2020).

Despite our simplified model, mechanical coupling between plates should not just be investigated only in variations in subduction velocity or dynamics, but also in the type of margin: accretionary or erosion. The definition of subduction interface in numerical models could also be relaxed. A recent methodological analysis (Sandiford and Moresi, 2019) investigated the  
465 emergence of significant spatial and temporal thickness variations within the interface zone, with the sole focus of maintaining a constant thickness through time. Instead, Beall et al. (2021); Riel et al. (2018) also investigated variations in interface thickness as potential controlling factor of dynamics. All these studies, in fact, highlight the tendency for the subduction interface to develop spontaneous thickness variation as the models evolve. The interface widens near the trench, building a prism-like complex, and thins at depths beyond the brittle–ductile transition. This pattern was also noted in the boundary element models  
470 of Gerardi et al. (2019), who attributed a down-dip thickness variation to lubrication layer dynamics. However, natural observations suggest that variable subduction fluxes enter the subduction trench, questioning this approach of constant thickness interface layer.

## 5 Conclusions

Systematic 2-D numerical simulations of ocean-ocean subduction are run to investigate how low-sediment sediment influence  
475 subduction dynamics and plate coupling. The aim is to understand what causes convergent margins to either accrete material delivered by the subducting plate or, alternatively, to subduct the trench sediment pile and even erode the basement of the overriding plate. We obtain end-member cases that are governed primarily by sediment viscosity and thickness: accretionary and tectonic erosion margins. We find that the properties of the sediment layer modulate the extent of viscous coupling at the interface between the subducting and upper plate. When the viscous coupling is increased, an erosive style of margin is  
480 favoured. On the other hand, when the viscous coupling is reduced, sediments are scrapped off the subducting slab to form an accretionary wedge. The geometry of the wedge controls the bending of the slab and the radius of the curvature. We perform an automated analysis of diagnostic parameters to differentiate between the two end-member modes of margin type and to better understand fundamental differences between them. Strong correlations between sediment, subduction diagnostics and margin type are observed. Tectonic erosion margins have smaller radii of curvature, wedge parameters and slower convergence rate,  
485 while accretionary margins are dominated by larger sediment wedges that can strongly influence subduction dynamics.

However, a more detailed study on the effect of sediments is needed, especially on buoyancy of sediments. The margin type, accretionary or tectonic erosion, is intimately linked to earthquakes. The amount of sediments filling the trench was proposed to facilitate seismic rupture (Heuret et al., 2012; van Rijsingen et al., 2018; Brizzi et al., 2020). Subduction zones with large amounts of trench sediments positively correlate with the occurrence of great interplate earthquakes.

490 *Code and data availability.* The Bitbucket version of the numerical code (LaMEM) used can be found here: [https://bitbucket.org/bkaus/lamem/branch/cvi\\_test](https://bitbucket.org/bkaus/lamem/branch/cvi_test), and the repository containing the input parameters files to reproduce the data can be found here: <https://adina@bitbucket.org/adina/rep-msubdsd.git>. The full simulation data (> 100Gb) presented in this study can be provided on request from AP.

*Author contributions.* AP and DS designed the study, AP built the numerical model and performed the analysis, MK helped run and analyze part of the simulations during REU internship. AP wrote first draft of manuscript, all authors contributed equally to the revision of this  
495 manuscript.

*Competing interests.* The authors have no competing interests.

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