

1 **Raman spectroscopy in thrust-stacked carbonates: an investigation of spectral parameters with**  
2 **implications for temperature calculations in strained samples**

3 **L. Kedar, C. E. Bond, D. Muirhead**

4 **University of Aberdeen**

5 *Note to editors: Since a major restructuring of the manuscript was required as a results of the*  
6 *reviewers' comments, only major additions/omissions are shown in the tracked changes here. For*  
7 *minor changes such as single-word changes, please see the Authors' Response to Reviewers.*

8 **Abstract.** Raman spectroscopy is commonly used to estimate peak temperatures in rocks containing  
9 organic carbon. In geological settings such as fold-thrust belts, temperature constraints are  
10 particularly important as complex burial and exhumation histories cannot easily be modelled. Many  
11 authors have developed equations to determine peak temperatures from Raman spectral  
12 parameters, most recently to temperatures as low as 75°C. However, recent work has shown that  
13 Raman spectra can be affected by strain as well as temperature. Fold-thrust systems are often highly  
14 deformed on multiple scales, with deformation characterised by faults and shear zones, and  
15 therefore temperatures derived from Raman spectra in these settings may be erroneous. In this  
16 study, we investigate how some of the most common Raman spectral parameters (peak width,  
17 Raman band separation) and ratios (intensity and area) change through a thrust-stacked carbonate  
18 sequence. By comparing samples from relatively low-strain localities to those on thrust planes and in  
19 shear zones, we show maximum differences of 0.16 for I[d]/I[g] and 0.11 for R2, while FWHM[d] and  
20 Raman Band Separation show no significant change between low and high strained samples.  
21 Plausible frictional heating temperatures of faulted samples suggest that the observed changes in  
22 Raman spectra are not the result of frictional heating. We also consider the implications of these  
23 results for how temperatures are determined using Raman spectra in strained and unstrained rock  
24 samples. ~~We apply three equations used to derive the peak temperatures from Raman spectra to~~  
25 ~~our data to investigate the implications on predicted temperatures between strained and unstrained~~  
26 ~~samples. All three equations produce different temperature gradients with depth in unstrained~~  
27 ~~samples. We observe that individual equations exhibit apparently varying sensitivities to strain, but~~  
28 ~~calculated temperatures can be upto 140°C different for adjacent strained and unstrained samples~~  
29 ~~using the same temperature equation. These results have implications for how temperatures are~~  
30 ~~determined in strained rock samples from Raman spectra.~~

31 **1 Introduction**

32 Raman spectroscopy can be used to provide information on the nanostructure of organic carbon in  
33 rocks (Tuinstra and Koenig, 1970; Landis, 1971; Nemanich and Solin, 1979; Knight and White, 1989;  
34 Ferrari and Robertson, 2001; Beyssac et al., 2002a; Muirhead et al., 2012; 2017a; Schito et al., 2017;  
35 Kedar et al., 2020; Muirhead et al., 2021). Since this nanostructure changes irreversibly with  
36 increasing temperature (e.g. Beyssac et al., 2002; Rahl et al., 2005; Huang et al., 2010; Muirhead et  
37 al., 2012), Raman is a useful tool for establishing the peak temperature of rocks in a variety of  
38 settings. In turn, peak temperatures can provide information about sedimentary burial conditions  
39 (e.g. Beyssac et al., 2002; Muirhead et al., 2012; 2017a; Schito et al., 2017), low-grade contact  
40 metamorphism (e.g. Aoya et al., 2010; Chen et al., 2017; Muirhead et al., 2017b) and tectonic thrust  
41 stacking. (Nibourel et al., 2018; Muirhead et al., 2019). Much work has been done to develop  
42 temperature equations that are based on Raman spectral parameters and are applicable across a  
43 range of settings and geological processes (e.g. Beyssac et al., 2002; Lahfid et al., 2010; Kouketsu et  
44 al., 2014; Schito and Corrado, 2018; Wilkins et al., 2018). With increasing understanding of organic

45 carbon nanostructure, such equations have been recently applied to a much wider range of  
46 temperatures (down to 75°C; Schito and Corrado, 2018; Muirhead et al., 2019) and geological  
47 settings such as fold-thrust belts (Nibourel et al., 2018, 2021; Muirhead et al., 2019).

48 Fold-thrust systems, driven by deformation, are subject to complex burial and exhumation histories  
49 and so the temperature history for a specific rock within a thrust stacked sequence is often not  
50 straightforward. In addition to this thermal complexity, recent work has shown Raman spectra to be  
51 affected by strain (Kwiecinska et al., 2010; Kitamura et al., 2012; Furuichi et al., 2015; Kedar et al.,  
52 2020). Therefore, if temperatures are to be investigated using Raman spectroscopy in strained  
53 terrains, the individual effects of strain and temperature on Raman spectra need to be isolated from  
54 one another.

55 In this study, we analyse the impact of strain and temperature on organic carbon nanostructure in a  
56 fold-thrust belt. Raman spectral parameters for a suite of samples, taken from a transect across a  
57 thrust-stacked carbonate sequence in the French Alps, are plotted on a cross-section. We use this  
58 visualisation, and the associated Raman data, to investigate how the most commonly used Raman  
59 spectral parameters in published temperature equations – peak intensity ratio ( $I[d]/I[g]$ ), Raman  
60 band separation (RBS), peak width (also known as full width at half maximum, or FWHM), and peak  
61 area ratio (R2) – change through the sequence. We also identify samples which are affected by  
62 locally high strain, such as thrust faults and shear zones, and assess how the Raman spectra of these  
63 samples differ from adjacent samples which have only been exposed to regional ‘background’ strain  
64 levels as opposed to localised deformation. By quantifying how each Raman spectral parameter,  
65 listed above, changes in strained samples, we assess the sensitivity to strain of each parameter. We  
66 compare faulted samples to those from a ductile shear zone to investigate the effect of seismic slip  
67 events vs. aseismic creep on Raman spectra, an important step towards separating the effects of  
68 strain and temperature. We discuss the implications of our findings on the ability of Raman-based  
69 geothermometers to predict geothermal gradients in thrust stacked sequences and to predict  
70 temperatures in locally strained rock samples.

## 71 **2 Geological setting**

72 The Haut Giffre region of the French Alps (Fig. 1a, 1b) encompasses 3,000 m of Jurassic-Cretaceous  
73 carbonates, split into six broad units (Fig. 1c) of contrasting mechanical properties. Each unit has a  
74 characteristic bed thickness, ranging from 0.005 m in shale-rich layers (Valanginian and Lower  
75 Oxfordian) to 10 m in the massive Tithonian carbonates. The complete sequence is shown in cross-  
76 section in Figure 1d. A regional cleavage pervades the stratigraphy, dipping NNW at a low angle (5-  
77 10°) with an average strike of around 210°. Approaching the Morcles Thrust, the cleavage shallows  
78 to horizontal and then steepens to dip at 20-40° towards the SE (Fig. 1b). Throughout the  
79 stratigraphy, the finer grained, more thinly bedded units, typically with higher organic carbon  
80 content, exhibit stronger cleavage.

81 The Morcles nappe, in which the study area lies, is the lowermost of the Helvetic nappes, and  
82 consists of a ‘normal’ limb and a lower overturned limb (Ramsay, 1980; Dietrich and Durney, 1986;  
83 Dietrich & Casey, 1989; Kirschner et al., 1999; Austin et al., 2008). The normal limb was subject to  
84 around 6 km of burial during the Alpine orogeny (Pfiffner, 1993; Kirschner et al., 1999; Austin et al.,  
85 2008). It is at this point that peak metamorphic (and hence maximum temperature) conditions are  
86 thought to have occurred; these remain ‘sub-greenschist’ (Kirschner et al., 1995). The overturned  
87 limb of the Morcles nappe outcrops in a 600 m-thick band which dips NW, parallel to the Morcles  
88 thrust below (Fig. 1d; note that here the Morcles thrust occupies the geometry of a low-angle  
89 normal fault). This overturned limb is mostly sheared Tithonian limestone in the study area, with

90 small wedges of other units included in thrust splays. Beneath the nappe, Triassic sands cap the  
91 Aiguilles Rouge massif, and these are together treated as basement here.

92 Regional scale thrust faults in the Haut Giffre cut through multiple carbonate units. The carbonate  
93 units themselves have contrasting mechanical properties. Massive or thickly-bedded limestones (e.g.  
94 Tithonian, Urgonian) act as competent beams, folding coherently on 100 m-scale wavelengths in the  
95 hanging walls and footwalls of thrusts. Interspersed between these massive limestones are a series  
96 of thinly-bedded, relatively carbon-rich shales and marls (e.g. Liassic, Lower Oxfordian, Valanginian)  
97 which have undergone internal deformation by means of incoherent folding and the formation of  
98 multiple internal detachment surfaces. The non-uniform distribution of strain in the Haut Giffre  
99 makes it the suitable subject of an investigation into the effect of strain on Raman spectra.

### 100 **3 Sampling strategy**

101 Samples were taken throughout the 3 km thick thrust-stacked sequence. Significant topographic  
102 relief in the form of inaccessible cliff sections necessitated sampling at laterally distributed sites (Fig.  
103 1b); these sites are represented in the cross-section (Fig. 1d) as lateral equivalents. Sample sites that  
104 could not be traced laterally to the section line, e.g. due to faulting or folding, are not included in the  
105 study.

106 Samples were categorised for their level of strain and classified on a simple binary scale, as either (1)  
107 being distal from thrusts or shear zones, where strain fabrics were present but not intense,  
108 indicative of a background level of strain; or (2) where zones of intense strain were present. These  
109 two sample site types are termed (1) “background” and (2) “strained” for the purpose of this study.  
110 It should be noted that although all samples have been subject to regional deformation, the term  
111 “strained” in this context implies that the samples have undergone localised deformation in the form  
112 of thrust faults or shear zones, as opposed to “background” strain levels.

113 In order to differentiate between sites of relatively high and low shear strain, outcrop- and hand  
114 specimen-scale strain indicators were visually analysed in the field. The NNW-oriented regional  
115 cleavage, more pervasive in the marls than the massive limestones, was used as a baseline for signs  
116 of a strain fabric within each lithology, where a more intense fabric (for example, one in which the  
117 cleavage more strongly overprinted the bedding features than in the surrounding rock) or a localised  
118 change in cleavage orientation which might suggest a localised intensification of the strain field.  
119 Usually such changes in cleavage orientation were associated with a fold hinge, a thrust, or a shear  
120 zone. Where these features were not present and the dominant fabric occupied a similar orientation  
121 and intensity to that of the regional cleavage, samples were considered to have undergone  
122 background levels of strain.

123 Localities were defined as ‘strained’ if they were associated with a structure such as a fold hinge,  
124 thrust or shear zone, and had therefore undergone locally high levels of shear strain in comparison  
125 to background levels. As mentioned above, often there was a change in the strain fabric from  
126 background levels – either in intensity or orientation, or both – approaching these structures.  
127 Moving away from the structure, once the fabric had returned to the regional intensity and  
128 orientation, samples were no longer considered ‘strained’.

129 Within strained localities, fine-scale (0.1 to 10m scale) transects were sampled perpendicular to the  
130 orientation of the structure – for example, in the case of a thrust, a sample would be taken on the  
131 thrust plane, 10-30cm above and below the thrust plane, and 1 to 10m above and below it,  
132 depending on where shear strain appeared to return to background levels. Five samples would  
133 therefore be the minimum number for a transect across a thrust. Within broader shear zones,

134 multiple samples were taken from within the shear zone itself, and then at least two at varying  
135 distances from the shear zone in either direction, provided the outcrop permitted this. Where  
136 complex shear zones contained multiple thrust splays on a metre scale, samples were taken from  
137 individual thrust planes and one or more would be collected from each intervening thrust slice.  
138 Samples were also taken from outwith the fault or shear zone to complete the transect.

139 Background samples were collected to establish the trend in parameters through the complete  
140 stratigraphy. At these sites, transects were not made and background sample sites represent either  
141 one sample or the average of a cluster of 2-3 samples from a single outcrop.

### 142 **3.1 Background sample sites**

143 Samples were deemed to be subject to background strain only, if they conformed to two criteria: (1)  
144 strain fabric at the sample site was parallel to the regional strain fabric in that area, and (2) the  
145 sample strain fabric was visually interpreted to be of similar intensity to the regional fabric in that  
146 unit. This interpretation, for the purposes of initial sample selection, was based on field observations  
147 and confirmed through inspection of micro-scale structures in thin section (note that most organic  
148 material was located between calcite grains and within seams of insoluble material). Practically, this  
149 meant that the sample was not part of a shear zone or a fault. Background samples were collected at  
150 distances of greater than 10 m from such localised high-strain zones. Where a high-strain zone was  
151 diffusely bounded, with a gradual return to background levels, the area was avoided entirely for the  
152 purpose of background sampling. Since the entire field area is part of a fold-thrust system, avoiding  
153 localised zones of high strain significantly limited potential sample sites. In total, 22 background  
154 samples from 15 different sites were included in the study, distributed approximately evenly within  
155 the intervening stratigraphy, between strained sites.

### 156 **3.2 Strained site samples**

157 Four “strained” sample sites were selected (Fig. 2), three of which are centred around thrusts  
158 (Tenneverge, Salvadon, and Finive), and one in the Emaney shear zone. They are described in detail  
159 below. Displacements across thrusts and shear zones are estimated from cross-sections based on  
160 field mapping.

#### 161 **3.2.1 Tenneverge (Fig. 2a)**

162 The Tenneverge thrust forms a discrete fault plane between Valanginian in the footwall and the  
163 Tithonian which overlies it (Fig. 2a(i)). Displacement here is estimated at 1 km (Fig. 1d). The  
164 Tithonian in the hanging wall is overturned at this locality, and the thinner beds at the base of the  
165 Tithonian sequence, which here lie (overturned) directly above the thrust plane, show localised tight  
166 chevron folding.

167 Samples collected from the Valanginian footwall at 12 m, 5 m and 0.5 m from the thrust plane show  
168 a gradual reduction in observable primary bedding towards the thrust, replaced by a deformation  
169 fabric of increasing intensity (Fig. 2a(ii)). This intensification is manifested in a transition from visible  
170 bedding planes, coupled with the regional sub-horizontal deformation fabric, into a dominant fault-  
171 parallel foliation which overrides the other fabrics. In the final 0.5 m below the thrust surface, the  
172 deformation fabric has been further deformed by rotation and small detachments, suggesting highly  
173 localised strain partitioning as part of a complex evolutionary history involving multiple fault  
174 movements and fabric overprinting. Intense veining accompanies this deformed layer (Fig. 2a(ii)). In  
175 the 10 cm adjacent to the thrust plane the fabric appears more coherent, with the foliation  
176 orientated parallel to the fault surface.

177 **3.2.2 Salvadon (Fig. 2b)**

178 The Salvadon thrust is a regional thrust fault with a maximum lateral displacement of 2 km towards  
179 the NW (Fig. 1d). At the sample site, Tithonian limestone is thrust over a thickened wedge of  
180 Valanginian shale. The thrust plane dips approximately 25° SE, with Tithonian bedding sub-parallel to  
181 this (strike 122°, dip 25° SW). The contrasting competencies of the lithologies here give rise to a  
182 discrete fault surface (Fig. 2b(i)).

183 Small (10 cm-scale) undulations in the fault surface (Fig. 2b(ii)), along with small fractures  
184 interrupting bedding at the base of the Tithonian, indicate that some strain was partitioned into the  
185 section of the Tithonian most proximal to the thrust (<0.5 m thick). Above this, the Tithonian loses  
186 evidence of additional horizontal strain, reverting to bedding-parallel stylolites and orthogonal sub-  
187 vertical fracture sets, common to the Tithonian throughout the Haut Giffre. In the footwall,  
188 Valanginian shales show evidence of increased strain several metres below the thrust surface. The  
189 first metre below the fault plane is dominated by a fault-parallel foliation (050/23° SE), which  
190 gradually rotates towards a more bedding-parallel orientation (122/25° SW) with distance from the  
191 fault surface over ~4 m. In a 4-5 m thick zone approaching the fault, en-echelon and orthogonal  
192 fracture sets (Fig. 2b(i)) are present in the footwall, indicative of a high degree of strain.. Foliation-  
193 parallel veins are also present in the upper 1m of the footwall, increasing in frequency towards the  
194 thrust plane (Fig. 2b(i)).

195 **3.2.3 Finive (Fig. 2c)**

196 The Finive sample site is an intraformational thrust splay (Fig. 2c(i)). The splay branches from a  
197 regional thrust below, which separates the sheared overturned lower limb of the Morcles Nappe  
198 from the normal limb above, and runs parallel to the Morcles thrust 400 m below (Fig. 1d). Above  
199 this regional thrust, Bajocian marls and Liassic shales are tightly folded and thrust. All samples at  
200 the Finive sample site are from the lowermost Bajocian, within which the intraformational thrust  
201 splay sits, and it is likely that this portion of the unit is overturned. However, most sedimentary  
202 features here have been heavily, if not fully, overprinted during deformation. Compositional layers  
203 have been stretched and thinned to 1-5 cm (Fig. 2c(ii)), around 10-30% of the thickness of such  
204 layers outwith this deformation zone. Boudinage and cm-scale folding are both common features in  
205 these compositional bands. Straight, foliation-parallel veins 1-5 mm thick are a pervasive feature.  
206 The thrust splay, around which sampling was concentrated, is parallel to the deformation fabric  
207 which is at a low angle to compositional layering (around 10° separation); as a result, displacement is  
208 difficult to estimate. Samples were taken from 10 m above, 0.1 m above, 0.1 to 1 m below, 2 m  
209 below, and 10 m below the thrust.

210 **3.2.4 Col d'Emaney (Fig. 2d)**

211 Samples taken from Col d'Emaney are from the base of the sheared lower limb of the Morcles  
212 nappe. Here, a wedge of Bajocian material overlain by a shale-rich unit (Fig. 2d(i)) is overthrust by  
213 overturned Tithonian. It is unclear as to whether the Bajocian wedge is overturned or not, as the  
214 overlying shale-rich unit is highly sheared. Previous geological surveys (e.g. those carried out by the  
215 French geological survey, BRGM) have mapped the shale as Oxfordian (suggesting this wedge is the  
216 right way up) but could also be overturned Liassic; distinction based on field observations is  
217 inconclusive, owing to the strong strain fabric that overprints sedimentary characteristics in the  
218 shale. The precise lithological unit is not of great importance here; what matters is the position  
219 within the overall thrust-stacked sequence, and the mechanical properties of the unit.

220 The contact of the shale unit with the Bajocian gives rise to a 4 m-thick shear zone within the shale,  
221 where deformation fabrics are greatly enhanced. S- and C-style fabrics are visible on a cm-scale,  
222 along with low-angle fractures which tend to run parallel to the shear fabrics and are bounded by  
223 rotated compositional bands (Fig. 2d(ii)). The combination of these ductile fabrics and brittle, blocky  
224 fractures suggests a complicated deformation history. Many of the S-C shears form tight clusters  
225 which act to increase the discontinuity between 'blocks' of material. Additionally, deformation-  
226 related undulations in the upper surface of the Bajocian (which resemble 0.5 m-scale normal faults,  
227 with rotation of the Bajocian cleavage to run parallel to the lithological contact) are accompanied by  
228 significant concentrations of ductile strain fabrics and veining in the shale above (Fig. 2d(ii)). The  
229 entire wedge is sheared to an extent, evidenced by strong cleavage and thinning of compositional  
230 bands. However, approaching the contact (i.e. within 1 m), the Bajocian cleavage rotates to be  
231 almost parallel to the perturbations in the contact surface, resulting in metre-scale 'waves' in the  
232 fabric. Within the shale, the highly concentrated deformation zone extends for around 4 m before  
233 the fabric consistently returns to the regional orientation.

## 234 **4 Raman spectroscopy**

### 235 **4.1 Introduction to Raman spectroscopy**

236 Raman spectroscopy measures the wavelengths of radiation produced by inelastic (Raman)  
237 scattering during the de-excitation of electrons in different molecular bonds, in this case focussing  
238 on those involved in different forms of organic carbon. In rocks, organic carbon can take on a range  
239 of nanostructures, depending on many factors during both deposition and burial: these include, but  
240 are not limited to, initial kerogen type, peak temperature, and the strain conditions to which the  
241 rock has been exposed. In the initial stages of burial, the carbon will have a nanostructure  
242 resembling that of kerogen (Thrower, 1989; Beyssac et al, 2002a; Rouzaud et al., 2015). As  
243 temperatures start to increase, the carbon nanostructure breaks down into smaller fragments as  
244 bonds are broken. With the application of strain (Kwiecinska et al., 2010; Kitamura et al., 2012;  
245 Savage et al., 2014; Furuichi et al., 2015; Kitamura et al., 2018) or very high temperatures (Wopenka  
246 and Pasteris, 1993; Oberlin et al., 1999; Schito et al., 2017), these fragments are aligned into parallel  
247 sheets, approaching a graphitic nanostructure. Graphitisation has not occurred in any samples in this  
248 study.

### 249 **4.2 Sample preparation and spectral acquisition**

250 A total of 62 samples were crushed and powdered before being treated with 10% HCl to remove  
251 inorganic carbon and therefore improve the signal to noise ratio when obtaining Raman spectra  
252 (Pasteris, 1989; Salver-Disma et al., 1999; Beyssac et al., 2002b; Mostefaoui et al., 2008; Muirhead et  
253 al., 2012). The residue was then rinsed and dried at room temperature to avoid thermal alteration.  
254 Using a Renishaw InVia Raman Spectrometer at the University of Aberdeen, a 514 nm laser was  
255 targeted at individual grains in the residual powders, where the laser power was <0.3 mW at the  
256 sample, and spot size was 1-2  $\mu\text{m}$ . Each run comprised three co-adds of 5 second acquisitions to  
257 produce a single spectrum for analysis. This process was carried out on 10 individual grains from  
258 each sample.

259 Backscattered radiation was recorded, deconvolved, and analysed using Renishaw WiRE 3.4  
260 software (Fig. 3). Using in-built software functions, noise reduction was first carried out on each  
261 spectrum (Fig. 3a), before baseline removal was performed using a cubic spline interpolation which  
262 was user-guided (Fig. 3b). Finally, a Gaussian curve fit was applied to the two visible peaks in the  
263 spectrum (Fig. 3c), as in Bonal et al. (2006), and the Raman spectra parameters (peak intensity,

264 position, width and area) were recorded (Fig. 3d; Quirico et al., 2009; Olcott Marshall et al., 2012).  
265 This process was carried out 3 times for each spectrum to minimise the error involved in the user-  
266 guided baseline removal process, resulting in 30 analyses per sample. Data points presented in this  
267 study therefore represent an average derived from 30 spectra per sample.

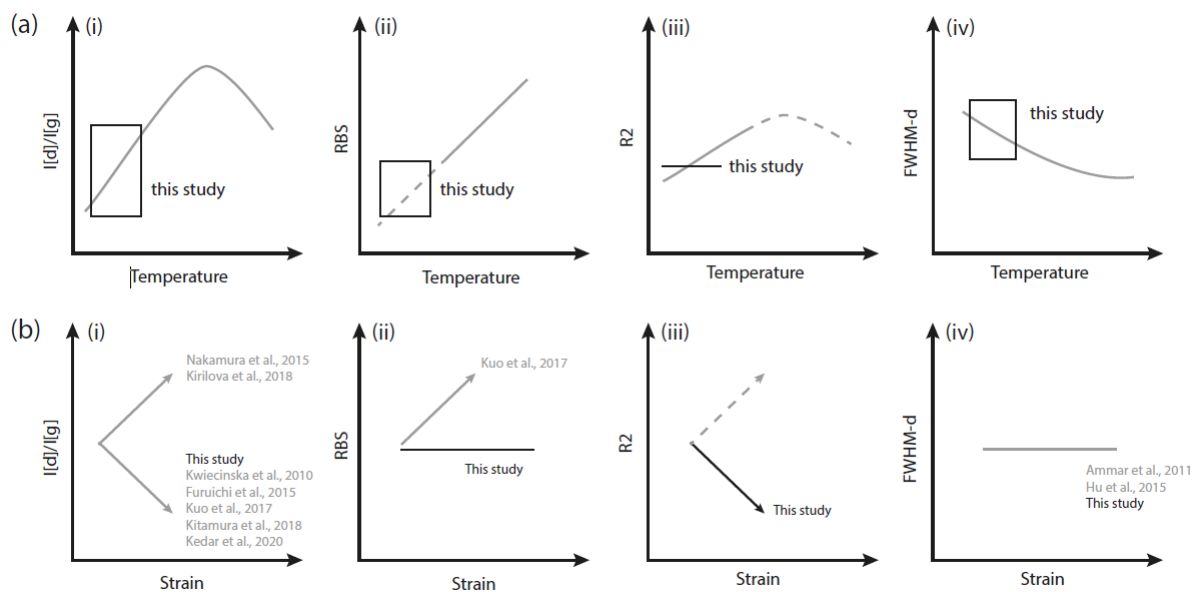
### 268 4.3 Raman spectral parameters

269 Different Raman spectral parameters are used in combination to determine the carbon  
270 nanostructure in a sample. Figure 3d highlights five key spectral parameters that can be calculated  
271 from the two curves fitted to the D- and G-peaks:

- 272 1. Peak area (A). The height and therefore area of a single peak is affected by signal  
273 strength, but comparing the areas beneath the D- and G-peaks negates this issue.  
274 The most common ratio comparison is known as 'R2', which is calculated as  
275  $A[d]/(A[g]+A[d])$ .
- 276 2. Peak position (W). This is the wavenumber of the peak. In this study we consider a  
277 broad D-peak around 1350 cm<sup>-1</sup>, and a sharp G-peak in the range of 1585-1610 cm<sup>-1</sup>.  
278 1.
- 279 3. Raman Band Separation (RBS). This is the difference between the two peak positions  
280 ( $W[g]-W[d]$ ).
- 281 4. Peak width (FWHM). Calculated as the 'Full Width at Half Maximum', FWHM is  
282 measured parallel to the horizontal axis.
- 283 5. Peak intensity (I). The intensity of a single peak is a direct product of signal strength,  
284 i.e., how many Raman-scattered photons come into contact with the detector. This  
285 can be affected by several factors including the amount of carbon present within the  
286 laser spot, or the strength of the laser. It is therefore more common to use the ratio  
287 between the D- and G-peaks ( $I[d]/I[g]$ ), which will be characteristic of the  
288 nanostructural features regardless of signal strength.

289 The G-peak defined here can be considered a composite of up to three spectral bands (D2, G, and  
290 D3) depending on metamorphic grade, but at low maturities such as those in this study they are  
291 difficult to distinguish and can be collectively referred to as a single peak (Beysac et al., 2002;  
292 Muirhead et al., 2021).

293 Figure 4 shows a schematic summary of the changes in Raman spectral parameters with increasing  
294 temperature and strain. At low maturities, as amorphous carbon degrades with increasing  
295 temperature, the D-peak increases in intensity (Tuinstra and Koenig, 1970; Levine, 1993; Oberlin et  
296 al., 1999 ). This increases the  $I[d]/I[g]$  ratio (Fig. 4a(i)). Subsequently, as higher maturities are  
297 reached, carbonaceous fragments align into sheets and the G-peak becomes more intense,  
298 decreasing  $I[d]/I[g]$  (Muirhead et al., 2012, 2017; Buseck and Beysac, 2014). Generally, a decrease in  
299  $I[d]/I[g]$  is observed when strain is applied to relatively low-maturity organic carbon (Fig. 4a(ii));  
300 Kwiecinska et al., 2010; Kitamura et al., 2012; Furuichi et al., 2015), but brittle fragmentation has  
301 been reported when mature, near-graphitic carbon is subject to low temperature strain, resulting in  
302 an  $I[d]/I[g]$  increase (Nakamura et al., 2015; Kirilova et al., 2018).



303

304 In addition to changing intensity, both the D- and G-peaks shift towards lower wavenumbers as the  
 305 material approaches complete graphitisation, but the D-peak shifts more significantly (Wopenka and  
 306 Pasteris, 1993; Beyssac et al., 2002; Quirico et al., 2009). This causes the RBS to change (e.g. Zhou et  
 307 al., 2014; Schmidt et al., 2017; Schito and Corrado, 2018; Henry et al., 2019). Figure 4b(i) shows an  
 308 increase in RBS with increasing temperature at higher maturities, but little is known about how the  
 309 parameter changes at low maturities such as in this study. The effect of strain has not been studied  
 310 extensively, although Kuo et al. (2017) reported an increase in RBS with the application of strain (Fig.  
 311 4b(ii)).

312 Figure 4c(i) illustrates how the R2 (curve area) ratio has been shown to correlate with temperature  
 313 (Beyssac et al., 2002; Aoya et al., 2010; Nakamura et al., 2015; Chen et al., 2017; Kirilova et al., 2018;  
 314 Henry et al., 2019). It follows a similar pattern to that of  $I[d]/I[g]$  with temperature, but little work  
 315 has been done to establish whether this similarity extends to strain (Fig. 4c(ii)).

316 It is generally accepted that the width (measured as full width at half-maximum, FWHM) of the D-  
 317 and G-peaks change with temperature (Zeng and Wu, 2007; Aoya et al., 2010; Kouketsu et al., 2014;  
 318 Zhou et al., 2014; Hu et al., 2015; Bonoldi et al., 2016; Chen et al., 2017). However, the nature of the  
 319 change varies depending on the thermal and barometric conditions, along with the nature of the  
 320 organic starting material. At relatively low temperatures (<300°C), these studies report a decrease in  
 321 FWHM-D with increasing temperature (Fig. 4d(i)), which also correlates with an increase in  $I[d]/I[g]$ .  
 322 FWHM-D is thought to undergo very little change, if any, when exposed to differential strain (Ammar  
 323 et al., 2011; Hu et al., 2015; Fig. 4d(ii)).

324 Therefore, comparison of the relative intensities, positions, widths, and areas of the D- and G-peaks  
 325 is a common method of assessing the extent to which a rock has been heated, which can be  
 326 correlated to geological processes such as burial depth, contact metamorphism, and exposure to hot  
 327 fluids. By comparison, the investigation of the effect of strain on these parameters is in its infancy  
 328 (Kwiecinska et al., 2010; Kitamura et al., 2012; Furuichi et al., 2015; Kuo et al., 2017; Kedar et al.,  
 329 2020).

#### 330 4.4 Temperature calculations

331 Using the parameters described above, three different methods of calculating temperature from  
 332 Raman data were employed. Choice of method was based on the assumption of regional peak



333 temperatures of less than  $\sim 300^{\circ}\text{C}$  in the field area, due to evidence of low-temperature  
334 metamorphism not exceeding sub-greenschist facies (Kirschner et al., 1995). This excludes certain  
335 Raman-based geothermometric equations which are only applicable above  $350^{\circ}\text{C}$  (e.g. Wopenka and  
336 Pasteris, 1993; Beyssac et al., 2002; Aoya et al., 2010).

337 The first equation used was that of Schito and Corrado (2018), which uses (in order of decreasing  
338 significance) the following Raman spectral parameters:  $I[d]/I[g]$  (intensity ratio), RBS (Raman Band  
339 Separation), FWHM[d] (D-peak width), FWHM[g] (G-peak width),  $A[d]$  (D-peak area) and  $A[g]$  (G-  
340 peak area) to give a % $R_0$  equivalent between 0.3 and 1.0. This is then converted to an approximate  
341 temperature value using the equation proposed by Barker and Pawlewicz (1994):

$$342 \quad T_1 = \frac{\ln(\%R_{0_{eq}}) + 1.68}{0.0124}$$

343 The second temperature calculation used is that of Kouketsu et al. (2014). This equation is based  
344 purely on FWHM[d] (D-peak width), and is reported as being effective in the range  $150^{\circ}\text{C} < T <$   
345  $400^{\circ}\text{C}$ :

$$346 \quad T_2 = -2.15(\text{FWHM}_D) + 478$$

347 The final temperature calculation used was that proposed by Lahfid et al. (2010), which converts  $R_2$   
348 (area-ratio) into temperature using the following equation:

$$349 \quad T_3 = \frac{R_2 - 0.3758}{0.0008}$$

350 where

$$351 \quad R_2 = \frac{A_D}{A_G + A_D}$$

352 Lahfid et al. (2010) report that their equation is most accurate within the range  $200^{\circ}\text{C} < T < 320^{\circ}\text{C}$ .

353

## 354 5 Raman spectral parameters across the fold-thrust belt

355 Figure 5 shows four individual Raman spectral parameters ( $I[d]/I[g]$ ,  $R_2$ , FWHM[d] and RBS) plotted  
356 on separate cross-sections, with each point corresponding to a sample site. Background samples  
357 taken distal from thrusts and shear zones are marked as single points. We first consider the  
358 relationship between the thrust plane and shear zone samples and their proximal neighbours, and  
359 how these relate to the regional trends observed across the area and visualised on the cross-section.  
360 Later we examine in more detail the 10 cm to 1 m-scale sampling in the interceding 10 m above and  
361 below the thrust planes and across the shear zones. Results are grouped by parameter, in each case  
362 with the “background” samples described first, followed by the “strained” samples.

363 Figure 6 shows detailed transects for each strained locality, demonstrating how the parameters  
364 change with distance from the fault or shear zone. Lithological unit is indicated by the colour of the  
365 chart area, whilst thrusts are represented by a thick dashed line and the Emaney shear zone by a  
366 series of thin, grey dashes. For context, the furthest points above and below the thrusts/shear zones  
367 in each panel of Figure 6 are the values labelled above and below strained samples in Figure 5.

### 368 5.1 $I[d]/I[g]$

369 I[d]/I[g] shows a general trend from lower values (0.3-0.4) in the upper stratigraphy to higher values  
370 (0.7-0.8) in the lowest stratigraphy (Fig. 5a) in background samples. This trend correlates with the  
371 depth through the thrust stack, and the graph of I[d]/I[g] with depth (Fig. 5a(ii)) highlights the  
372 gradual increase towards higher values approaching the basal thrust. The Morcles thrust flattens  
373 towards the NW end of the cross-section, coinciding with a more vertical trend in I[d]/I[g] values.  
374 This trend appears to be disrupted across thrust planes and shear zones in the strained samples. In  
375 the case of the Salvadon thrust, the samples taken 10 m above and below the thrust plane are  
376 consistent with the regional trend, with values of 0.377 and 0.417 respectively. However, I[d]/I[g]  
377 values on the thrust plane are 25-30% lower, with an average value of 0.249. There is a similar  
378 though slightly less significant drop on the Tenneverge thrust plane, from 0.552 and 0.541 to 0.469,  
379 a decrease of around 18%. In the case of both the Salvadon and Tenneverge thrusts, there is a  
380 gradual reduction in I[d]/I[g] approaching the thrust plane, which occurs more abruptly and closer to  
381 the thrust plane in the hanging wall than the footwall where the change is more gradual. On the  
382 Finive thrust plane there is a drop to 0.632 from surrounding values of 0.739 and 0.772. The Emaney  
383 shear zone, which is not immediately associated with a major fault, exhibits values of 0.624 and  
384 0.627 within the shear zone itself, with higher values (0.772 and 0.888) outwith the shear zone, a  
385 similar change as observed at the thrust localities.

## 386 5.2 RBS

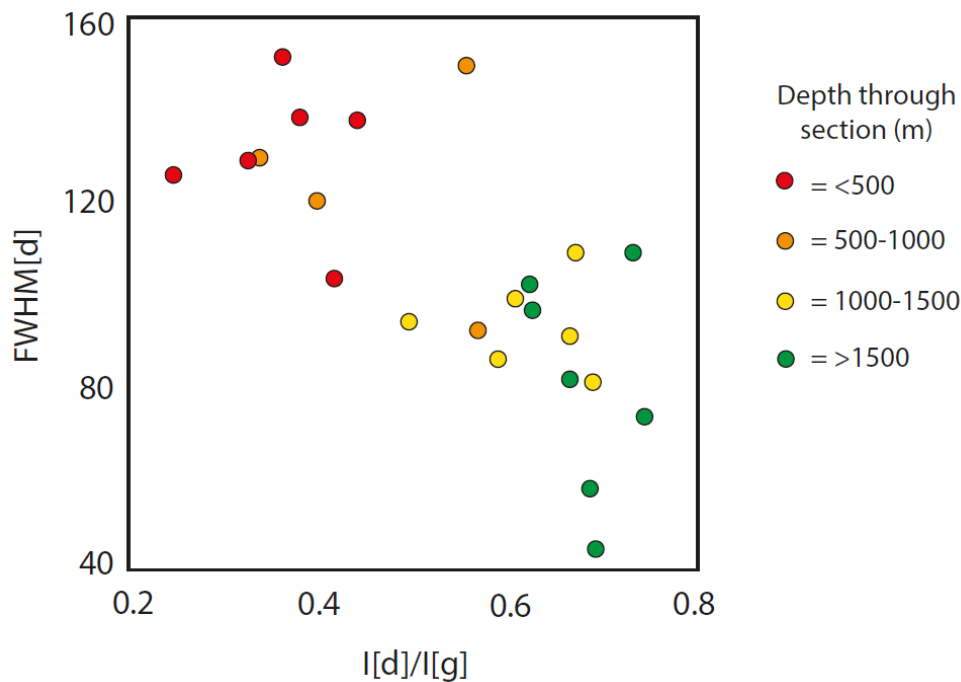
387 Raman band separation (RBS) varies through the stratigraphy (Fig. 5b), with what appears to be a  
388 prevalence of values >265 in the upper stratigraphy and <265 in the lower sequence. The average  
389 error associated with each sample is +/- 4, suggesting that the change through the stratigraphic  
390 sequence is not significant. Unlike I[d]/I[g], there is no distinct shift on the thrust planes (Figure 6);  
391 instead the RBS value on the Tenneverge and Salvadon thrust planes sits between the values of  
392 samples taken immediately above and below the thrust. There is a small shift to slightly higher  
393 values in the Emaney shear zone (263/261 from 255/259 below the shear zone), but these values are  
394 accompanied by an approximate error of ±2, which is comparable to magnitude of the difference  
395 between samples.

## 396 5.3 R2

397 R2 (the area ratio; Fig. 5c) shows little overall change with depth towards the Morcles thrust,  
398 although there is a weak trend towards higher values with increasing depth within individual  
399 thrust packages. Outwith the strained samples, there is a prevalence of values <0.6 towards the  
400 top of the stratigraphy, and >0.6 in the lower half, but many values are accompanied by errors as  
401 high as ±0.1. There is a marked drop on each thrust plane and within the shear zone, similar to the  
402 behaviour seen in I[d]/I[g]. On the Salvadon thrust, there is a decrease from 0.572 (20 m below the  
403 thrust plane) and 0.568 (10 m above the thrust) to 0.460 on the thrust plane. Here, the pattern  
404 resembles that of the I[d]/I[g] change for this thrust, with a more gradual change approaching the  
405 thrust from the footwall than in the hanging wall. On the Tenneverge thrust plane the R2 value is  
406 0.486, compared to 0.503 at a distance 0.5m above and 0.551 at 5m below (note the different  
407 distances are due to the extent of the shear zone bounding the thrust, which is thicker in the  
408 footwall). In the case of the Finive thrust, R2 is 0.512 on the thrust plane, whilst above and below it,  
409 R2 sits at 0.580 and 0.555 respectively. Finally, there is also a decrease in R2 moving from 10 m  
410 below the Emaney shear zone (0.560 and 0.564) to within it (0.549 and 0.529). The average error for  
411 each sample was +/- 0.078, so only the Salvadon thrust samples exhibit a change with a greater  
412 magnitude than this, but importantly (at least within the immediate vicinity of the strained samples)  
413 the direction of change is consistent.

414 **5.4 FWHM[d]**

415 Only FWHM[d] is shown in Fig. 5d, as FWHM[g] varies significantly and shows no discernible trend in  
416 our data (the reader is referred to Supplementary Material). There appears to only be a small shift, if  
417 any, on thrust planes and in the shear zone. There is, however, a slight general decrease in FWHM[d]  
418 with depth towards the Morcles thrust (Fig. 5c) and therefore a reverse correlation with  $I[d]/I[g]$ , as  
419 illustrated in Figure 7.



420

421 **6-Temperature calculations across the fold thrust belt**

422 Using the above parameters, the results of Equations 1, 2 and 3 (from Schito and Corrado (2018),  
423 Kouketsu et al. (2014), and Lahfid et al. (2010), respectively) were plotted on the cross-section for  
424 each sample site.

425 **6.1 Temperatures based on Schito and Corrado (2018)**

426 Applying the Schito and Corrado (2018) equation — based on  $I[d]/I[g]$ , RBS, FWHM[d], FWHM[g],  $A[d]$   
427 and  $A[g]$  — to our data gives a calculated temperature range across all the samples of 79°C to 104°C  
428 (Fig. 7a). In background samples within intra-fault stratigraphic packages, there is a slight trend  
429 towards higher temperatures with increased depth towards the Morcles thrust. The calculated  
430 temperatures appear to show a 10–20°C increase as thrusts are approached. There is little change on  
431 thrust planes compared to surrounding values, with the only significant change being that of the  
432 Salvadon thrust, where the thrust plane ‘temperature’ is calculated at 93°C compared to 100°C and  
433 103°C above and below it. In contrast, there is a slight increase in calculated temperature in the  
434 Emaney shear zone (85 and 88°C) compared to below it (79 and 81°C).

435 **6.2 Temperatures based on Kouketsu et al. (2014)**

436 Since the Kouketsu et al. (2014) equation relies solely on FWHM[d], the calculated temperatures  
437 show a similar (although reversed) pattern to that produced by the FWHM[d] parameter itself. There  
438 is a slight overall trend towards higher temperatures towards the basal thrust (290–320°C compared  
439 to 250–280°C in the upper section), but the pattern is irregular and a clear trend is difficult to

440 identify. There is a drop of around 30°C on the Tenneverge thrust plane, and 5–10°C in the Emaney  
441 shear zone (Fig. 7b). However, the Salvadon and Finive thrust planes do not show significant change,  
442 and fall within the variation of temperatures calculated for the samples above and below those  
443 thrusts.

### 444 **6.3 Temperatures based on Lahfid et al. (2010)**

445 The Lahfid et al. (2010) equation calculates temperatures as a linear derivative of R<sub>2</sub>, and so the  
446 trend is similar to that of R<sub>2</sub> but easier to discern due to the larger range of values (138–312°C).  
447 There is a clear increase in temperature both with depth and with proximity to the basal Morcles  
448 thrust (Fig. 7c). There is disruption in the vicinity of regional-scale folds and thrusts; notably,  
449 temperatures above and below the Salvadon thrust are significantly higher (240°C) than those at  
450 similar stratigraphic levels (150–190°C). On all three thrust planes and within the shear zone, there is  
451 a marked decrease in calculated temperature, ranging from a 20°C difference (Tenneverge) to 140°C  
452 (Salvadon) below the temperatures calculated for the surrounding samples. As with the Schito-  
453 Corrado temperature calculation, there is a slight increase (10–30°C) in apparent temperature  
454 approaching the thrusts compared with ‘background’ values.

## 455 **6 Discussion**

### 456 **6.1 I[d]/I[g]**

457 At low thermal maturities, increasing temperatures cause a rise in I[d]/I[g] (Fig. 4a; Dietrich and  
458 Casey, 1989; Muirhead et al., 2012; Sauerer et al., 2017). In the Haut Giffre, we estimate peak burial  
459 temperatures to be 150–250°C, based on a standard geothermal gradient of 25°C/km and an  
460 estimated burial depth of 6km at the top of the exposed stratigraphic pile (Pffifner, 1993; Kirschner  
461 et al., 1999; Austin et al., 2008). Therefore, it is reasonable to suggest that the increase in the  
462 I[d]/I[g] ratio towards the basal Morcles thrust is associated with increasing peak temperature, and  
463 hence maximum burial. This fits with the observations, for example, of Schito et al. (2017), where a  
464 similar trend of increasing I[d]/I[g] with depth is seen through some 4 km of core containing  
465 siliciclastics from the Lower Congo Basin, Angola, at temperatures up to 170°C.

466 There is a significant drop in I[d]/I[g] on thrust planes and in shear zones, with as much as a 40%  
467 decrease in the ratio values on the Salvadon thrust plane compared to the surrounding stratigraphy  
468 (0.417 to 0.279, a difference of 0.168) and 30% difference between samples in the Emaney shear  
469 zone to those adjacent (0.888 to 0.624, a difference of 0.264). There are several possible causes of  
470 this drop:

- 471 1. A lower peak temperature on the fault plane than the surrounding rock would  
472 reduce the I[d]/I[g] value on the fault plane. However, there is no plausible  
473 mechanism to explain how this would occur.
- 474 2. A very large temperature increase on the fault plane (>500°C) could cause such a  
475 spectral change that I[d]/I[g] values begin to drop again (Fig. 4a; Bustin et al., 1995;  
476 Furuichi et al., 2015; Kaneki et al., 2016; Nakamura et al., 2019). A possible  
477 mechanism for such a local temperature elevation could be flash heating due to  
478 friction on the fault plane. Frictional heating is known to occur on fault planes  
479 (Goldsby and Tullis, 2007; Smith et al., 2015), particularly in episodes of rapid  
480 seismic slip (Rabinowitz et al., 2020). However, the magnitude and duration of  
481 elevated temperatures from friction depend on a range of factors such as  
482 permeability, slip duration, and fault thickness (Bustin, 1983; Mase and Smith, 1987;  
483 Fulton and Harris, 2012; Kitamura et al., 2012), and there is therefore uncertainty as

484 to whether this would always be sufficient to alter the Raman spectra. Mase and  
485 Smith (1987) modelled frictional heating on fault planes and found that in porous  
486 rocks, the slip duration would have to be much greater than 100 seconds for thermal  
487 pressurisation to occur. Our results show that  $I[d]/I[g]$  decreases similarly in thrust  
488 faults and in broader shear zones. Transient frictional heating cannot explain the  
489 decrease in  $I[d]/I[g]$  values in the Emaney shear zone, which has undergone mostly  
490 ductile deformation over a more widely distributed area.

491 3. Strain-related spectral changes can also reduce  $I[d]/I[g]$  (Kwiecinska et al., 2010;  
492 Kitamura et al., 2012; Furuichi et al., 2015; Kedar et al., 2020), and this would be  
493 applicable to both fault planes and distributed shear zones. Kedar et al. (2020)  
494 reported a drop of 0.1 to 0.15 in  $I[d]/I[g]$  in the sheared, overturned limb of a  
495 recumbent isoclinal fold, corresponding to an increase in strained microfabrics in  
496 those samples.

497 It is worth noting that although the downward shift in  $I[d]/I[g]$  is very prominent on the Salvadon  
498 thrust plane, there is also a gradual decrease in values as the thrust plane is approached (Figure 6).  
499 In the footwall, which comprises Valanginian marls with a gradual intensification of the strain fabric  
500 approaching the thrust,  $I[d]/I[g]$  values begin to decrease several metres out from the thrust plane.  
501 However, in the hanging wall, which consists of Tithonian limestone,  $I[d]/I[g]$  remains high until  
502 much closer to the thrust. This pattern is also observed in the case of the Tenneverge thrust, which  
503 comprises the same lithologies as the Salvadon thrust. It is possible that such a pattern is indicative  
504 of the respective rheological properties of the hanging wall and footwall lithologies, with the softer  
505 footwall marls forming a deformation shear zone and hence lowering the  $I[d]/I[g]$  value further from  
506 the thrust plane itself, if strain is indeed the mechanism by which the carbon structure is changing.  
507 Meanwhile, the more competent hanging wall limestones do not form such a broad deformation  
508 zone and therefore the potentially strain-related shift in  $I[d]/I[g]$  is reserved for a much narrower  
509 zone just above the thrust plane.

## 510 6.2 RBS

511 Raman band separation (RBS) is reported to increase with increasing temperature (Fig. 4b; Zhou et  
512 al., 2014; Bonoldi et al., 2016; Sauerer et al., 2017), and so should increase with depth towards the  
513 basal thrust in our study. Pressure also affects peak positions (Ross and Bustin, 1990; Bustin, 1995;  
514 Huang et al., 2010). However, the trend is weak in our RBS data. If frictional heating on fault planes  
515 were the primary control on changes in RBS, and we assume an approximate instantaneous slip  
516 magnitude of ~1 m, then it would be expected that temperatures could rise by >400°C (Savage et al.,  
517 2014). This should be enough to produce a shift in RBS which is greater than the general variation we  
518 see in our samples. However, Nakamura et al. (2019) report that in addition to temperature, RBS is  
519 sensitive to lithology and the effects of fluids, which may explain the variable results we see in this  
520 study.

## 521 6.3 FWHM[d]

522 D-peak width (FWHM[d]) exhibits a reverse trend to that of  $I[d]/I[g]$ , decreasing slightly with depth  
523 towards the basal thrust. This supports experiments by Zeng and Wu (2007), who observed a  
524 decrease in FWHM[d] with increasing temperature, although their experiments were on samples at  
525 300°C and above. Zhou et al. (2014) also observed a decrease in FWHM[d] with increasing  $I[d]/I[g]$  in  
526 solid bitumen, similar to the trend seen in this study. Studies on coal approaching a magmatic  
527 contact by Chen et al. (2017) indicate that both FWHM[d] and FWHM[g] decrease with increasing  
528 temperature, but in this study we only observe a decrease in FWHM[d], whilst FWHM[g] changes

529 very little. However, the starting material in this study is amorphous carbon rather than coal, which  
530 has a different crystalline structure. This may account for differences between our study and that of  
531 Chen et al. (2017).

532 Unlike  $I[d]/I[g]$ , there is little change in FWHM[d] across fault planes or shear zones, which suggests  
533 that the two parameters are not directly related. For samples affected by polishing during sample  
534 preparation, it has been noted that Raman spectral peak widths are less influenced than peak  
535 intensities (Ammar et al., 2011; Hu et al., 2015), suggesting that FWHM[d] does not change  
536 significantly due to shearing. In light of our results, it may be possible to extend this suggestion to  
537 shearing on fault planes and in shear zones. This is supported by the results of fine-scale transects  
538 across strained localities (Figure 6), where the error and general variation in FWHM[d] seems to  
539 outweigh any significant shift on fault planes or in shear zones.

#### 540 **6.4 R2**

541 R2 (the area ratio) shows a similar trend to that of  $I[d]/I[g]$ , only not as pronounced. Note that the  
542 range in R2 values is lower than that of  $I[d]/I[g]$  due to the normalised denominator used to  
543 calculate R2 (Equation 1), so a weaker trend than  $I[d]/I[g]$  is expected. There is a drop in R2 on thrust  
544 planes and within the Emaney shear zone, this drop is 6-20% (a difference of 0.035; see Fig. 5(c)),  
545 compared to 15-40% for  $I[d]/I[g]$ . However, the percent change in R2 on fault planes is comparable  
546 to the total change in R2 between the upper and lower sections of the stratigraphy (though this is  
547 subject to a high degree of variation).

548 Since peak area is a product of peak intensity and peak width, it follows that R2 is dependent on  
549  $I[d]/I[g]$  and FWHM[d]. These two parameters have opposing trends with depth through the  
550 sequence, resulting in a general dampening of any R2 trend. However, in strained samples,  $I[d]/I[g]$   
551 tends to drop, whilst FWHM[d] remains unchanged. This means that R2 also drops, and makes this  
552 parameter more sensitive to strain-related spectral changes than to burial trends.

#### 553 ~~7.5 Schito and Corrado (2018) calculated temperatures~~

554 ~~The Schito and Corrado (2018) equation uses  $I[d]/I[g]$ , RBS, FWHM[d] and [g], and individual peak~~  
555 ~~areas to calculate %Ro, which can be subsequently used to estimate temperature. In this study,~~  
556 ~~temperatures estimated using the Schito and Corrado (2018) equation sit between 70 and 110°C.~~  
557 ~~This range is lower than expected for rocks that have been buried to 6-9km, as suggested by~~  
558 ~~previous studies (Pfiffner, 1993; Kirschner et al., 1999; Austin et al., 2008). This could be due to a~~  
559 ~~weak geothermal gradient, or that the equation is not applicable in this instance due to the Raman~~  
560 ~~parameters used.~~

561 ~~The temperature trend calculated using this equation indicates a low geothermal gradient of~~  
562 ~~~15°C/km. If this thermal gradient were to be extrapolated upwards through the previous overlying~~  
563 ~~stratigraphy, the range of 70-110°C at 6km depth would almost be appropriate. It is possible that~~  
564 ~~through thrust stacking, the regional scale geothermal gradient could be flattened; if the majority of~~  
565 ~~thrust emplacement occurred during exhumation, then peak temperatures could have remained~~  
566 ~~relatively low.~~

567 ~~The second term in the equation is RBS, which does not exhibit a trend in our data. As a result, its~~  
568 ~~presence in the equation may subdue the range of estimated temperatures. Therefore, at lower~~  
569 ~~maturities, the RBS term in the Schito and Corrado (2018) equation may become irrelevant and~~  
570 ~~make it less effective for temperature determination.~~



571 Our data shows a small shift (of the order of 10°C) in temperature on fault planes and in shear zones,  
572 but the direction and exact magnitude is inconsistent (as noted previously by Muirhead et al., in  
573 review). Since the most significant term in the equation is  $I[d]/I[g]$ , and our data shows that  $I[d]/I[g]$   
574 is strongly affected by strain-related spectral changes, it follows that the equation should be  
575 sensitive to strain. Although a strain-induced error in apparent temperature of  $\pm 10^\circ\text{C}$  will not  
576 significantly impact the performance of the Raman geothermometer, it highlights the importance of  
577 context when estimating temperatures using this method.

578 A transient temperature rise during frictional heating on a fault plane may be too short-lived to  
579 promote spectral changes (Bustin, 1983; Fulton and Harris, 2012; Kitamura et al., 2012; Furuichi et  
580 al., 2015). This may be one explanation for the lack of a consistent temperature increase on the  
581 thrust planes in this study. Other explanations include differing thicknesses of active slip (Raboniwitz  
582 et al., 2020), or a low slip magnitude in a single event (Polissar et al., 2011; Savage et al., 2014;  
583 Savage et al., 2018; Raboniwitz et al., 2020). However, these do not explain the distinct drop in  
584 apparent temperature on the Salvadon thrust plane, or the elevated temperature values within the  
585 Emaney shear zone, where frictional heating should not play a role.

#### 586 7.6 Kouketsu et al. (2014) calculated temperatures

587 The Kouketsu et al. (2014) equation for calculating temperature uses just the FWHM[d] parameter,  
588 and gives temperatures of 200–380°C. The lower end of this temperature range overlaps with that  
589 expected for a burial depth of 6–9 km, but we would not expect to see temperatures above 250°C.  
590 Our data suggests that strain should have a limited effect on temperatures derived from this  
591 equation, since FWHM[d] is reportedly insensitive to strain (Ammar et al., 2011; Hu et al., 2015).

592 In the burial trend, there is significant variation on a sub-km scale. In their paper, Kouketsu et al.  
593 (2014) highlight an error of  $\pm 30^\circ\text{C}$  associated with the equation. This magnitude of error,  
594 necessitates temperature trends to be identified over km-scale distances or greater for a normal  
595 geothermal gradient.

#### 596 7.7 Lahfid et al. (2010) calculated temperatures

597 Using the Lahfid et al. (2010) equation, our data yields apparent temperatures (138–312°C) that are  
598 closer to the expected range for a burial depth of 6–9 km (150–250°C). However, the trend shown by  
599 these temperatures indicates a high geothermal gradient (70–80°C/km) that is not easily explained  
600 through thrust tectonics.

601 Since the Lahfid et al. (2010) equation is entirely dependent on  $R_2$ , it follows that strain will  
602 significantly affect the results. This fits with a consistent drop of 40–50°C observed on fault planes  
603 and in shear zones within the study area.

#### 604 7.8 Summary

605 From our observations we suggest that the Schito and Corrado (2018) equation is less effected by  
606 strained environments than the Lahfid et al. (2010) equation. The Kouketsu et al. (2014) equation is  
607 also more suited to strained environments. However, the Schito and Corrado (2018) equation  
608 produces temperature estimates and a geothermal gradient lower than expected for the region and  
609 the Kouketsu et al. (2014) equation, in our case study, shows variation in temperature predictions on  
610 a sub-km scale, making it less suitable for general use. Unlike the Schito and Corrado (2018)  
611 equation, the Lahfid et al. (2010) equation demonstrates a more consistent error in the most  
612 strained rocks, and the predicted temperatures are more in line with those predicted for the area.  
613 The consistency of the shift on thrust planes and in shear zones with the Lahfid et al. (2010)

614 temperature calculation suggests that it might be possible to correct for this with contextual sample  
615 knowledge, or by comparison with other equations.

## 616 **6.5 Implications for Raman geothermometry**

617 There are numerous geothermometric equations which use Raman spectral parameters to calculate  
618 predicted maximum temperatures. These have been developed for different geological settings,  
619 each applicable to a particular carbon type, predicted temperature range, and methodology (such as  
620 laser wavelength and deconvolution method). Such examples include but are not limited to Lahfid et  
621 al. (2010), Kouketsu et al. (2014), Wilkins et al. (2018), Schito and Corrado (2018), and Muirhead et  
622 al. (2019). However, our work demonstrates that strain has an effect on certain Raman spectral  
623 parameters, with some more affected than others. This could have significant implications for the  
624 results of Raman geothermometers. Here we consider one geothermometric equation – that of  
625 Schito & Corrado (2018) – as an example in order to demonstrate how strained rocks might impact  
626 results.

627 This equation uses (in order of decreasing significance) the following Raman spectral parameters:  
628 I[d]/I[g] (intensity ratio), RBS (Raman Band Separation), FWHM[d] (D-peak width), FWHM[g] (G-peak  
629 width), A[d] (D-peak area) and A[g] (G-peak area) to give a %Ro equivalent between 0.3 and 1.0:

$$630 \quad Ro = -3.0211 + 0.33633(I_D/I_G) + 0.01251(RBS) + 0.0024823(FWHM_D) - 0.000000376(A_D) \\ 631 \quad \quad \quad - 0.0033158(FWHM_G) + 0.000000595(A_G)$$

632 The result is then converted to an approximate temperature value using the equation proposed by  
633 Barker and Pawlewicz (1986):

$$634 \quad T_1 = \frac{\ln(\%Ro_{eq}) + 1.68}{0.0124}$$

635 Applying the this to our results gives a calculated temperature range across all the samples of 79°C  
636 to 104°C, with a 10°C increase in local temperature on the Salvador thrust plane and a 5-10°C  
637 decrease in the Emaney shear zone; otherwise, little change is noted in the strained samples.

638 An actual temperature increase due to frictional heating would explain any increase in temperature  
639 on a fault plane but may not account for observations in a ductile shear zone. Further to this,  
640 frictional heating could easily raise temperatures to values outside the calibration range of the  
641 Schito & Corrado (2018) equation, giving erroneous results using this equation and necessitating the  
642 use of a different geothermometer in these samples. However, a transient temperature rise during  
643 frictional heating on a fault plane may be too short-lived to promote spectral changes (Bustin, 1983;  
644 Fulton and Harris, 2012; Kitamura et al., 2012; Furuichi et al., 2015). Inconsistencies in the  
645 magnitude of temperature change can potentially be explained by differing thicknesses of active slip  
646 (Raboniwitz et al., 2020), or a low slip magnitude in a single event (Polissar et al., 2011; Savage et al.,  
647 2014; Savage et al., 2018; Raboniwitz et al., 2020). However, these do not explain the distinct drop in  
648 apparent temperature on the Salvador thrust plane, or the elevated temperature values within the  
649 Emaney shear zone, where frictional heating should not play a role.

650 The most significant term in the equation is I[d]/I[g], and our data shows that I[d]/I[g] is strongly  
651 affected by strain-related spectral changes. It therefore follows that the equation should be sensitive  
652 to strain, but the fact that not all strained samples produce calculated temperature shifts of the  
653 same direction or magnitude suggests that the process is more complex than simply strain or  
654 temperature having an effect. Regardless of cause, however, an error in calculated temperature of  
655 ±10°C in a stratigraphic sequence with an overall temperature range of only 25°C highlights the



656 importance of context when estimating temperatures using this method. For example, if using this  
657 temperature data to reconstruct a burial history, then a strained sample might be 'out' by over a  
658 kilometre, or it might give the correct value. It is therefore important that more work is done to  
659 calibrate Raman geothermometers in rocks which have undergone strain in natural environments.

### 661 **8.1 Calculated temperature trends**

662 Spectral data from the samples in this study were applied to three different temperature equations  
663 developed by Schito and Corrado (2018), Kouketsu et al. (2014) and Lahfid et al. (2010) respectively.

664 Between thrusts, the Schito and Corrado (2018) equation produced a broad trend of weakly  
665 increasing temperatures with depth. Overall, these temperatures were lower than expected. The  
666 Lahfid et al. (2010) equation produced temperatures which were within the expected range for 6 km  
667 of burial, but also indicated a very high geothermal gradient. The Kouketsu et al. (2014) equation  
668 gave temperatures which were higher than expected for the proposed burial depth, with only a  
669 slight trend towards higher values at depth, and a high degree of variation. This variation suggests  
670 that the Kouketsu et al. (2014) equation is unsuitable for establishing temperature gradients over  
671 distances of less than a few km. In fold-thrust systems such as the one investigated here, in which  
672 calculated temperatures may be influenced by strain, an equation is required that can resolve  
673 temperature changes over hundreds of metres at the least. We conclude that although various  
674 geothermometric equations carry applicable temperature ranges, choosing the most appropriate  
675 equation is complex and dependent on multiple factors.

### 676 **8.2 Impact of strain on calculated temperatures**

677 Choice of equation is particularly important in the context of fold-thrust systems, where strain  
678 intensity can vary on multiple scales. The use of multiple parameters in the Schito and Corrado  
679 (2018) equation suggests that the equation should be relatively insensitive to strain. However, on  
680 thrust planes, the Schito and Corrado (2018) calculated temperatures dropped by 0-10°C, while  
681 increasing by the same magnitude in a distributed shear zone. This suggests that the equation is  
682 indeed sensitive to strain-related spectral changes, likely due to the fact that  $I[d]/I[g]$  is the  
683 dominant term in the equation. The use of multiple terms in the equation may help to produce more  
684 reliable results (as the influence of different parameters counteract), it is important to consider  
685 which parameters may have the most influence in different areas and indeed different samples.

686 The results of the Kouketsu et al. (2014) equation showed considerable variations in temperature,  
687 with no consistency when moving from background to locally strained samples. Therefore, despite  
688 being based on a parameter which is, in theory, relatively unaffected by strain, temperatures  
689 produced by this equation appear unreliable in this case. The Lahfid et al. (2010) equation recorded  
690 a consistent drop of 40-50°C in strained samples, suggesting that this temperature equation is  
691 significantly affected by strain-related spectral changes. The consistency of this calculated  
692 temperature drop suggests that it might be feasible to compare these results with those of another  
693 equation, and/or the geothermal gradient to distinguish the effect of strain-related spectral changes  
694 from those induced by temperature. This finding is important not only for improving the robustness  
695 of Raman spectroscopy as a geothermometer in fold-thrust systems, but also for the potential to  
696 develop a Raman-based strain tracker.

## 697 **8 Conclusions**

698 Analysis of samples from an Alpine carbonate fold-thrust system has revealed trends and anomalies  
699 in Raman spectral data. We chose four key parameters which are frequently used to assess thermal  
700 maturity of organic carbon in rock samples and plotted the values at the corresponding sample sites  
701 on a cross-section. By separating samples that had been affected by locally high strain (such as on  
702 fault planes or in shear zones) from those that had only been subjected to the background regional  
703 strain, and by plotting metre-scale transects across these strained sites, we were able to apply  
704 context to the data and hence discern regional thermal trends from localised strain-related  
705 anomalies.

706 Parameters showed varying sensitivities to strain and temperature. In background samples,  $I[d]/I[g]$   
707 increased with depth towards the basal thrust, suggesting an expected ‘burial trend’. FWHM[d]  
708 decreased with depth, whilst R2 – a product of  $I[d]/I[g]$  and, to some extent, FWHM[d] – increased  
709 slightly. RBS showed no discernible trend. In strained samples,  $I[d]/I[g]$  dropped by 0.1 to 0.15 (up to  
710 40% depending on location in the stratigraphy), and R2 showed a small decrease. There was little  
711 change, if any, in FWHM[d] or RBS in strained samples. In the fine-scale transects across the  
712 Salvador and Tenneverge thrusts,  $I[d]/I[g]$  and R2 both showed a gradual decrease towards the  
713 thrust in the footwall marls where a shear zone was present, whilst the decrease appeared to be  
714 more abrupt and closer to the thrust plane in the hanging wall limestones. We suggest that this may  
715 be due to rheological differences in the two lithologies and may also be linked to the differing levels  
716 of strain experienced by the hanging wall and footwall of a thrust fault. In both cases, the driving  
717 factor for the rate of decrease of  $I[d]/I[g]$  and R2 approaching the thrust would be the extent of  
718 strain partitioning in the rocks immediately above and below the thrust surface. A lack of change –  
719 or at least a lack of consistency and magnitude – in FWHM[d] to match these  $I[d]/I[g]$  changes  
720 suggests that strain rather than temperature is the main driver. This is supported by the similarity in  
721 the results plotted across the Emaney shear zone. However, since the influence of frictional heating  
722 cannot be ignored in the case of a thrust surface, there is scope for future work to attempt to further  
723 separate these two signals in naturally deformed rocks such as these.

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725

## 726 **CRedit Author Statement**

727 Kedar: fieldwork, Raman spectroscopy analysis and interpretation, original draft preparation, figure  
728 preparation; Bond: original conceptualisation, input into rewriting and framing original draft,  
729 fieldwork (support); Muirhead: Raman spectroscopy interpretation, input into writing and re-  
730 drafting of original draft.

731 All authors have contributed to the writing and framing of the manuscript and discussion of all  
732 concepts.

## 733 **Declaration**

734 The authors declare that they have no conflict of interest.

## 735 **Acknowledgements**

736 This study was carried out as part of a University of Aberdeen PhD, supported by the UKRI Centre for  
737 Doctoral Training in Oil & Gas [grant number NE/R01051X/1].

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