

A functional tool to explore the reliability of micro-earthquake focal mechanism solution for seismotectonic purposes

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ABSTRACT

Improving the knowledge of seismogenic faults requires the integration of geological, seismological, and geophysical information. Among several analyses, the definition of earthquake focal mechanisms plays an essential role in providing information about the geometry of individual faults and the stress regime acting in a region. Fault plane solutions can be retrieved by several techniques operating in specific magnitude ranges, both in the time and frequency domain and using different data.

For earthquakes of low magnitude, the limited number of available data and their uncertainties can compromise the stability of fault plane solutions. In this work, we propose a useful methodology to evaluate how well a seismic network, used to monitor natural and/or induced micro-seismicity, estimates focal mechanisms as a function of magnitude, location, and kinematics of seismic source and consequently their reliability in defining seismotectonic models. To study the consistency of focal mechanism solutions, we use a Bayesian approach that jointly inverts the P/S long-period spectral-level ratios and the P polarities to infer the fault-plane solutions. We applied this methodology, by computing synthetic data, to the local seismic network operating in the Campania-Lucania Apennines (Southern Italy) aimed to monitor the complex normal fault system activated during the Ms 6.9, 1980 earthquake. We demonstrate that the method we propose is effective and can be adapted for other case studies with a double purpose. It can be a valid tool to design or to test the performance of local seismic networks and more generally it can be used to assign an absolute uncertainty to focal mechanism solutions fundamental for seismotectonic studies.

INTRODUCTION

Fault plane solutions represent primary information ~~that seismologists can retrieve~~ to describe ~~the earthquake source~~earthquakes. The assessment of earthquake location, magnitude, and focal mechanism are the fundamental operations to characterize the earthquake source ~~through the~~using a point source approximation. ~~After the earthquake location, origin time, and source dimension are identified, the~~The focal mechanism describes the basic geometry and kinematics of a point source in terms of strike, dip, and rake of the fault plane along which the earthquake occurred. So, the focal mechanism is the most important ~~parameter that can be retrieved to recognize the marker of the~~ geometry of the seismogenic faults and their style of faulting. Moreover, ~~the~~ seismicity and focal mechanisms of events, ~~also of small magnitudes~~, are often used to constrain seismotectonic models, individual seismogenic sources, the regional strain, and stress fields, also for small magnitudes. Consequently, an evaluation of their effective reliability becomes a fundamental issue in seismotectonic studies.

Nevertheless, focal mechanisms cannot be calculated and constrained every time an earthquake occurs. Although the calculation of focal mechanisms represents a routine analysis ~~inside the~~for seismological agencies, the solutions are calculated only for a specific range of magnitudes, usually greater than 4. In fact, constraining the solution for earthquakes with small magnitude is still ~~represents~~ a challenge, despite the advancement in the technological process and the use of increasingly performing seismic networks. This is due to several factors that we will analyse in detail. The techniques used to define the focal mechanism of large-to moderate earthquakes are based on the inversion of the moment tensor, ~~that~~which corresponds to a stable and robust procedure, so much that it is the most common method for this type of analysis (Dreger, 2003; Delouis, 2014; Sokos and Zahradnik, 2013; Cesca et al., 2011). This technique requires accurate knowledge of the propagation medium in relation to the range of frequencies used for the modelling ~~of the~~ waveforms recorded during an earthquake. The smaller an earthquake, the higher the frequency range of the signal to be modelled, the more detailed the

67 knowledge and scale of the Earth's interior must be. Several methods have been proposed to achieve
 68 a stable inversion of the moment tensor for earthquakes with a magnitude less than 3. Hybrid ~~approach~~
 69 ~~of the approaches that invert both~~ amplitude and waveform moment tensor ~~inversions, which utilizes use~~
 70 the principal component analysis of seismograms (Vavrycuk et al., 2017) or moment tensor refinement
 71 techniques (Kwiatek et al. 2016; Bentz et al., 2018) ~~to~~ facilitate a robust determination of the source
 72 type and its kinematics. In particular, the retrieved moment tensor is typically decomposed into
 73 volumetric and deviatoric components. Constraining the earthquake as a double-couple source can
 74 erroneously affect the retrieved fault plane solutions, especially in the case of induced seismicity where
 75 the volumetric or non-double couple component must be considered (Kwiatek et al. 2016).

76 Other analytical techniques are based on the recognition of the source radiation pattern ~~that describes~~
 77 ~~the earthquake source.~~ According to the position of seismic stations ~~with respect relative~~ to the source,
 78 seismic waves on seismograms show different amplitudes and polarities. These features ~~are employed~~
 79 ~~in a very simple way by several algorithms to can~~ constrain the geometry of the earthquake faulting
 80 through estimating the angular parameters strike, dip, and rake. The classical method (Raesenberg and
 81 Oppenheimer, 1985;) uses the P-wave polarities, ~~but;~~ more advanced ~~ones use approaches better~~
 82 ~~constrain the focal mechanism of small earthquakes using~~ P- or S- wave amplitudes or amplitude ratios
 83 together with first motions (Snoke, 2003) ~~to better constrain the focal mechanism of small~~
 84 ~~earthquakes.~~ In fact, the use of polarities alone is ~~not convenient inappropriate~~, especially if we consider
 85 micro-seismicity ($M < 3$). The reasons could be the limited number of available data, their uncertainties,
 86 and the difficulty of measuring the P-polarity with a sufficient degree of precision. For these reasons,
 87 different techniques using different types of measurements such as P-wave amplitudes (Julian and
 88 Foulger, 1996; Tarantino et al., 2019), P/S or S/P amplitude ratios measured in the time or the
 89 frequency domain (Kisslinger et al., 1981; Rau et al., 1996; Hardebeck and Shearer, 2003; De Matteis
 90 et al., 2016), or S-wave polarizations (Zollo and Bernard, 1991) have been developed. The joint
 91 inversion of polarities and amplitude ratios led to more stable and robust solutions, allowing to account

92 for geological site effects and to decrease ~~in first approximation~~ the effects produced by the geometric
93 and anelastic attenuations.

94 Two kinds of errors generally influence the goodness of the solution and retrieved model (Michele et
95 al., 2016): the perturbation errors that are related to how the uncertainty on data affects the model,
96 and the resolution errors that are referred to the capability to retrieve a correct model, given a dataset
97 as input or how accurate could be the model that we can recover, even ~~if with~~ error-free data ~~are used~~.
98 The sum of perturbation and resolution errors corresponds to the final errors on the model obtained
99 by solving an inverse problem, as the solution of focal mechanism. In particular, the resolution errors
100 depend on the available data, and so on the initial condition of the inverse problem. In the case of focal
101 mechanism, the number of seismic stations, as well as the seismic network geometry, and the velocity
102 structure of the crust influence the resolution and the reliability of the retrieved model.

103 How will the geometry of a seismic network determine the accuracy of focal mechanism solutions? The
104 answer to this question ~~is not simple and~~ requires a deep knowledge of the geophysical and geological
105 characteristics of the region, often ~~unrealistic~~ unavailable. Moreover, the theoretical relationships that
106 predict the focal mechanism solutions for an earthquake scenario could be very complicated if several
107 factors, such as network configuration, noise level, source magnitude, or source kinematics are taken
108 into account. ~~We want to underline that a~~ network configuration may be optimal for earthquake
109 locations, but not for retrieving fault plane solutions (Hardt and Scherbaum, 1994). In fact, a given
110 geometry may resolve some fault kinematics better than others.

111 A seismic network layout is strictly associated with the goals of the network and the available funds;
112 according to these features, a network operator decides how many stations are required and where
113 they should be located (Havskov et al.; 2011). So, the number of seismic stations, the size, and geometry
114 of the network are defined after a preliminary phase based on the evaluation of the specific
115 seismological target (Trnkoczy et al., 2009; Hardt and Scherbaum 1994; Steinberg et al. 1995; Barta
116 et al. 2000). In the case of small earthquakes, the available recordings come from only a portion of the
117 total network, while the distant stations show a seismic signal buried inside the noise. In order to detect

118 and locate low-magnitude earthquakes, we must increase the number of seismic stations for area units
119 by building a dense seismic network.

120 In this study, we propose a useful tool to evaluate both 1) the reliability of focal mechanism solutions
121 inferred by the inversion of different seismological data and 2) the performance of the seismic network
122 to assess focal mechanism solutions and their errors. We evaluate the network capability to solve focal
123 mechanisms as a function of magnitude, location, and kinematics of seismic source. We consider three
124 synthetic data set: P-wave polarities, P- S-wave amplitude spectral ratios and polarities and amplitude
125 ratios together. Moreover, different levels of noise are considered in order to simulate more realistic
126 conditions.

127 We selected as target the Irpinia Seismic Network (ISNet), a local seismic network that monitors the
128 Irpinia complex normal fault system (Southern Italy), activated during the Ms 6.9 earthquake of 23rd
129 November 1980. Evaluating the specific performance of an existing network for a seismological goal is
130 critical and can be used to decide how to improve its layout.

131

132 **METHODOLOGY**

133 With the main aim to define the reliability of focal mechanisms retrieved by specific seismic networks,
134 we propose a methodology based on an empirical approach that consists of different steps.

135 ***Configuration and Parameter Tuning (Step 1).*** In a preliminary phase, we select for each earthquake
136 simulation the: a) fault plane solution to test, b) seismic observables to be computed (i.e. P-wave
137 polarities or P- S-wave amplitude spectral ratios), c) magnitude, d) the earthquake epicentre and depth;
138 e) the network geometry; f) the noise level. The fault plane solution to test can be derived from
139 instrumental seismicity as one of the strongest earthquakes occurred in the area or a median solution
140 of the available ones or simply a fault plane solution representative of the regional seismotectonic.
141 Once the network geometry and the hypocentre of the earthquake are defined, the seismic stations
142 (number and type) for which the synthetic data are computed must be selected. The number of seismic

stations that record an event depends on earthquake magnitude, source-stations distance, crustal medium properties, and the level of noise. We use an empirical approach, based on the statistical analysis of the local seismicity catalog, that allows us to define, for each magnitude range, a maximum (threshold) epicentral distance for which only the seismic stations within this distance are considered (See data analysis).

Synthetic Data Computation (Step 2). Using a crustal velocity model and the source-receiver relative position, the synthetic data are computed for the theoretical fault plane solution. The seismic observables that can be reproduced are a) P-wave polarities, b) P/S spectral amplitude ratios, and c) polarities and amplitude ratios together. For the P/S spectral level ratios, the Gaussian noise level is added.

Focal Mechanism Inversion (Step 3). We estimated focal mechanism using BISTROP code (De Matteis et al., 2016) that jointly inverts the ratio between the P- and S-wave long-period spectral levels and the P-wave polarities according to a Bayesian approach. BISTROP has the advantage to use different observables for the determination of fault plane solutions, such as the P/S long-period spectral level ratios or P-wave polarities, individually or together. The benefits of the use of spectral level ratios are multiples: 1) they can be measured for a broad range of magnitudes (also for $M < 3$; De Matteis et al., 2016); 2) they can be calculated by automatic procedures without visual inspection; 3) their estimates do not require to identify the first arrival time accurately, but only a time window of signal containing P- or S-phase is mandatory and 4) the spectral amplitude ratios, they can generally be used without the exact knowledge of the geological soil conditions (site effects) and geometric/anelastic attenuation. Moreover, the joint inversion of amplitude spectral ratios and polarities led to constraining fault plane solutions reducing the error associated with the estimates of retrieved parameters. BISTROP solves an inverse problem through a probabilistic formulation leading to a complete representation of uncertainty and correlation of the inferred parameters.

For a double-couple seismic source, the radiation pattern depends on fault kinematics and relative source-station position. In fact, it can be represented as a function of 1) strike, dip and rake angles (φ ,

169 δ, λ) and 2) take-off and azimuth angles (i_h, φ_r). We can define the ratio between P- and S-wave
 170 radiation pattern coefficients as:

$$171 \quad \frac{\mathcal{R}^P(\phi, \delta, \lambda, i_h, \phi_R)}{\mathcal{R}^S(\phi, \delta, \lambda, i_h, \phi_R)} = \left(\frac{\alpha_s^2 \alpha_r}{\beta_s^2 \beta_r} \right) \frac{\Omega_0^P}{\Omega_0^S} \quad (1)$$

172 where Ω_0^P and Ω_0^S are the long-period spectral level of the P- and S-waves, respectively, and $\alpha_s, \alpha_r,$
 173 β_s, β_r , are the P- and S-wave velocities at the source and at the receiver, respectively. Thus, using the
 174 displacement spectra, assuming a given source and attenuation model (Boatwright, 1980), we can derive
 175 from the signal recorded by a seismic station the ratio of radiation pattern coefficients for P- and S-
 176 phases, as well as $\alpha, \beta, i_h, \varphi_r$ are known from the earthquake location and the velocity model used. So,
 177 from a theoretical point of view, the spectral amplitude ratios measured at several seismic stations can
 178 be used to retrieve the ratio of radiation pattern coefficients $\mathcal{R}_{\theta\varphi}^P/\mathcal{R}_{\theta\varphi}^S$ as a function of the source-
 179 receiver azimuth and take-off angles.

180 BISTROP jointly inverts the spectral amplitude ratios with the observed P-wave polarities to infer the
 181 parameters φ, δ, λ of the focal mechanism in a Bayesian framework. A posterior probability density
 182 function (PDF), for the vector of model parameter \mathbf{m} (φ, δ, λ) and the vector of observed data \mathbf{d} , is
 183 defined as:

$$184 \quad q(\mathbf{m}|\mathbf{d}) = \frac{f(\mathbf{d}|\mathbf{m})p(\mathbf{m})}{\int_M f(\mathbf{d}|\mathbf{m}')p(\mathbf{m}') d\mathbf{m}'} \quad (2)$$

185
 186 where $f(\mathbf{d}|\mathbf{m})$ is the conditional probability function that represents the PDF given the data \mathbf{d} and for
 187 parameter vector \mathbf{m} in the model parameter space \mathbf{M} , and $p(\mathbf{m})$ is the a priori PDF. If P-wave polarities
 188 and P/S spectral level ratios are independent datasets, the conditional probability function may be
 189 written as:

190

$$f(\mathbf{d}|\mathbf{m}) = f(\mathbf{d}^L|\mathbf{m})f(\mathbf{d}^P|\mathbf{m}). \quad (3)$$

in which the pdf of the data vector \mathbf{d}^L of N^L measurements of spectral ratios is multiplied for the pdf of data vector \mathbf{d}^P of N^P measurements of P-wave polarities given the model \mathbf{m} .

Assuming that the observables have the same finite variance, for the N^L observations of spectral level ratios the conditional probability function may be defined as:

$$f(\mathbf{d}^L|\mathbf{m}) = \frac{1}{(\sqrt{2\pi}\sigma)^{N^L}} \exp\left(-\frac{\sum_{i=1}^{N^L}\{d_i - [G(\mathbf{m})]_i\}^2}{2\sigma^2}\right) \quad (4)$$

Where $G(\mathbf{m})$ represents a functional relationship between model and data and corresponds to Equation 1 and σ represents the uncertainty on the spectral measure.

For the N^P observations of P-wave polarities, the conditional probability function is (Brillinger et al., 1980):

$$f(\mathbf{d}^P|\mathbf{m}) = \prod_{i=1}^{N^P} \frac{1}{2} [1 + \psi(\mathcal{R}_i^P, \gamma_i, \rho_0) Y_i \text{sign}(\mathcal{R}_i^P)] \quad (5)$$

in which:

$$\psi(\mathcal{R}_i^P, \gamma_i, \rho_0) = (1 - 2\gamma_i) \text{erf}(|\rho_0 \mathcal{R}_i^P(\mathbf{m})|) \quad (6)$$

The quantity reported in square brackets in Equation 5 represents the probability that the observed i_{th} polarity γ_i is consistent with the theoretical one computed from the model \mathbf{m} , whose theoretical P-wave

211 amplitude is \mathcal{R}_i^P and $sign(\mathcal{R}_i^P)$ is its polarity at i_{th} station for a given fault plane solution. The
 212 parameters ρ_s and γ_0 , referring to the errors in ray tracing due to velocity model ambiguity and to the
 213 uncertainty on polarity reading, regulating the shape of the PDF. For more details about the
 214 mathematical formulation, see De Matteis et al. (2016).

215 ***Evaluation of the Results (Step 4).*** Once the best solution is estimated, the focal mechanism
 216 uncertainties and its misfit, respect to the theoretical solution as Kagan angle, are computed. The focal
 217 mechanism parameter (strike, dip and rake) misfit and their uncertainties are also calculated.

218

219 IRPINIA SEISMIC NETWORK

220 As testing case of our methodology, we choose the area of the M 6.9, 1980 Irpinia earthquake
 221 (Southern Italy). Since 2005, ISNet, a local, dense seismic network monitors the seismicity along the
 222 Campania-Lucania Apennines covering an area of about $100 \times 70 \text{ km}^2$ (Figure 1; Weber et al., 2007).
 223 The seismic stations are deployed within an elliptic area whose major axis, parallel to the Apennine
 224 chain, has a NW-SE trend with an average inter-stations distance of 15 km that reaches 10 km in the
 225 inner central zone. Each seismic station ensures a high dynamic range and it is equipped with a strong-
 226 motion accelerometer, Guralp CMG-5T or Kinometrics Episensor, and a short period three-component
 227 seismometer, Geotech S13-J with a natural period of 1 sec. In 6 cases, broadband seismometers are
 228 installed such as the Nanometrics Trillium with a flat response in the range 0.025–50 Hz. ISNet is
 229 operating by INFO (Irpinia Near Fault Observatory) and it provides real-time data at local control centres
 230 for earthquake early warning systems or real-time seismic monitoring (Satriano et al., 2011). Seismic
 231 events are automatically identified and located from continuous recordings by automatic Earth-worm
 232 Binder and data are then manually revised by operators (Festa et al., 2020).

233 The 1980, M 6.9, Irpinia earthquake was one of the most destructive, instrumental earthquakes of the
 234 Southern Apennines, causing about 3000 fatalities and severe damages in the Campania and Basilicata
 235 regions. It activated a NW-SE trending normal fault system with a complex rupture process involving

multiple fault segments according to (at least) three different nucleation episodes delayed each other of 20 s (Bernard and Zollo, 1989; Pantosti and Valensise, 1993; Amoroso et al., 2005). No large earthquakes occurred in the Irpinia region since 1980. A Mw 4.9 earthquake took place in 1996 originating a seismic sequence inside the epicentral area of the 1980 earthquake (Figure 1; Cocco et al., 1999). Recent instrumental seismicity occurs mainly in the first 15 km of the crust showing fault plane solutions with normal and normal-strike slip kinematics, indicating a dominant SW-NE extensional regime (Pasquale et al., 2009; De Matteis et al., 2012; Bello et al., 2021). Low-magnitude seismicity ($M_L < 3.6$) is spread into a large volume related to the activity of major fault segments of the 1980 Irpinia earthquake (Figure 1; Adinolfi et al., 2019; Adinolfi et al., 2020). Seismic sequences or swarms often occurred in the area, extremely clustered in time (from several hours to a few days) and space and seem to be controlled by high pore fluid pressure of saturated Apulian carbonates bounded by normal seismogenic faults (Stabile et al., 2012; Amoroso et al., 2014).

248

249 DATA ANALYSIS

We applied the method we proposed and evaluated the capability of the ISNet local network to resolve fault plane solutions using different observables as input data: a) P-wave polarities, b) P/S spectral amplitude ratios and c) polarities and amplitude ratios together. the analysis is carried out by evaluating the effect of 1) earthquake magnitude, 2) epicentral location, 3) earthquake depth, 4) signal-to-noise ratio, and 5) fault kinematics on retrieved focal solutions as previously described.

Step 1. In order to select focal mechanisms (FMs) to be used for our resolution study (Figure 2a), we carried out statistical analysis to define the most frequent fault plane solutions of instrumental seismicity. We classified, according to the plunge of P- and T-axes, the fault plane solutions reported in De Matteis et al. (2012) choosing only the FMs occurring within the Irpinia area since 2005 to 2011. As shown in Figure 2b, splitting the range of the data into equal-sized bins, we selected the focal mechanism corresponding to the median value of the most populated class. We report it in Figure 2a

as FM2. This corresponds to a normal-strike-slip fault plane solution with strike, dip, and rake equal to 292° , 53° , and -133° , respectively. Then, we decided to test the focal mechanism solution of the 1980 Irpinia earthquake, a pure normal fault (strike, dip, rake: 317° , 59° , -85° ; Westaway and Jackson, 1987; Fig. 2a) here and after FM1. This solution is very similar to the focal mechanism corresponding to: 1) the regional stress field (see Supplementary Material); 2) the M_L 2.9, Laviano earthquake, one of the most energetic earthquakes of the last years (Stabile et al.; 2012), and 3) those of the 2nd, 3rd, 4th most populated bins. Finally, we selected the solution corresponding to the 5th bin reported as FM3 in Figure 2a. This focal mechanism is quite different from the others due to a predominant component along the fault strike (strike, dip, rake: 274° , 71° , -128°)

Step 2. For each of the three selected fault plane kinematics, we calculated synthetic data (P-wave polarities or P- and S-wave spectral amplitudes) at seismic stations varying the earthquake location and by using a local velocity model (Matrullo et al., 2013). We discretize the study area with a square grid ($100 \times 100 \text{ km}^2$), centred on the barycentre of ISNet, with 441 nodes and a sampling step of 5 km. Each node corresponds to a possible earthquake epicentre (Figure 3).

For each grid node and according to the earthquake magnitude to be tested, we have to select the ISNet stations for simulations. The number of seismic stations that record an event depends on earthquake magnitude, source-stations distance, crustal medium properties, and the noise level. Theoretical relationships that link the seismic source to the signal recorded at every single station are quite complicated (Kwiatek et al., 2016; 2020) and are based on the accurate knowledge of crustal volumes in which the seismic waves propagated, such as the three-dimensional wave velocity structure, anelastic attenuation or/and site conditions of a single receiver. To overcome this limitation, we used an empirical approach to define the number and the distance of the seismic stations that record a seismic signal as a function of magnitude, once its epicentral location (grid node) and depth are fixed. Using the bulletin data retrieved by INFO at ISNet during the last two years (January 2019-March 2021; <http://isnet-bulletin.fisica.unina.it/cgi-bin/isnet-events/isnet.cgi>), we selected two earthquake catalog datasets with depths equal to 5 (± 2) km and 10 (± 2) km, respectively, and local magnitude ranging

between 1.0 and 2.5. These choices are motivated by the characteristics of the Irpinia micro-seismicity recorded by ISNet. Then, we divided each dataset into bins of 0.5 magnitudes and for each bin, we retrieved the median number of P-wave polarity readings and the median epicentral distance of the farthest station that recorded the earthquake (Table 1). The bulletin data are manually revised by operators, and we selected only seismic records that provide P- and/or S- wave arrival times. The median value of the distance of the farthest station is then used to select the seismic stations for which synthetic data are calculated. Therefore, for each earthquake simulation of specific magnitude and depth, only the seismic stations with a distance, from the grid node under examination (epicentre), equal or lower than the maximum distance, reported in Table 1, are considered. We run simulations only for earthquakes recorded at least by 6 seismic stations. The synthetic P-wave polarities are simulated only at a number of stations corresponding to the median value previously defined. (Table 1). We pointed out that the number of P-wave polarities empirically assigned is related to the available earthquake catalogue data of the Irpinia region where the seismicity can occur in different portions of the area covered by the network, not always with optimal azimuthal coverage.

Additionally, we simulated the uncertainty on the measure of spectral level ratios or the effect of seismic noise adding a zero mean, Gaussian noise to the synthetic data with a standard deviation equal to two different percentage levels, as 5% and 30%. With this configuration, we simulated:

- Three datasets of seismic observables: P-wave polarities (D1), P/S spectral level ratios (D2) and polarities and P/S spectral level ratios together (D3)
- Two hypocentre depths: 5 km and 10 km
- Three magnitude bins: M_L 1.0 - 1.5 (M1), M_L 1.5 - 2.0 (M2) and M_L 2.0 - 2.5 (M3)
- Three focal mechanism solutions: FM1 (317°, 59°, -85°), FM2 (292°, 53°, -133°) and FM3 (274°, 71°, -128°)

Two level of Gaussian noise: 5% and 30%. When D2 is simulated, in order to solve the verse ambiguity of the slip vector, a P-wave polarity is added to the earthquake data to be inverted for the focal mechanism.

313 **Step 3.** For each earthquake simulation the focal mechanism was estimated by inverting the synthetic
314 data with BISTROP (De Matteis et al.; 2016).

315 **Step 4.** In order to analyse the results, we defined five kinds of map to study how the focal mechanism
316 (FM) resolution and error spatially change in the area where ISNet is installed (Table 2):

317

- 318 • Kagan angle misfit map (KAM)
- 319 • Map of the focal mechanism parameter misfit (FMM)
- 320 • Strike, Dip and Rake error map (FME)
- 321 • Kagan angle average map (KAA)
- 322 • Kagan angle standard deviation map (KAS)

323

324 The Kagan Angle (KA) measures the difference between the orientations of two seismic moment tensors
325 or two double couples. It is the smallest angle needed to rotate the principal axes of one moment tensor
326 to the corresponding principal axes of the other (Kagan et al.; 1991; Tape and Tape; 2012). The smaller
327 the KA between two focal mechanisms, more similar they are. In KAM map, for each node the value of
328 KA between the theoretical and retrieved solution is reported, while in FMM map, the absolute value of
329 the misfit between the strike, dip, and rake angles of the retrieved and theoretical solution is indicated.
330 FME is defined as the error map of strike, dip, and rake in which the uncertainties (standard deviations)
331 are calculated considering all the solutions with probability larger than the 90% (S90) of the maximum
332 probability, corresponding to the best solution retrieved. Additionally, these solutions are used to study
333 how constrained is the FM solution. The KA is calculated between each FM of S90 solutions and the
334 retrieved best solution. The mean and the standard deviation of the resulting KA distribution are plotted
335 in KAA and KAS maps, respectively. The smaller KA mean and std, the more constrained is the obtained
336 fault plane solution (Table 2).

337

338 **DISCUSSION**

We consider the FM1, i.e. the focal mechanism of the 1980 Irpinia earthquake located at 10 km depth, first. Looking at Figures 4 and 5, we see the effect of using the three different datasets. Considering D1, we can calculate the FM only for earthquakes with magnitude 2.0-2.5 for which at least 6 polarities are available. As shown by KAM map in Figure 4a, the retrieved solutions are characterized by high KA ($> 50^\circ$) with limited areas or single nodes with values in the range 40° - 50° . Therefore, D1 ~~is not sufficient to~~ cannot retrieve with acceptable accuracy the FMs for earthquakes with magnitude 2.0-2.5. The same result is obtained for FM2 and FM3 (Figure 4b-c). Comparing the results of the simulations using D2 and D3 (Figure 5), the accuracy of the retrieved solution is improved when P-wave polarities data are added to spectral level ratios. The areas in KAM map with high value of KA ($KA \geq 18^\circ$; red or green areas) disappear or are strongly reduced. Nevertheless, ~~we want to underline that,~~ even with D2 dataset, ~~exception some small areas,~~ the FMs are well retrieved for all magnitudes with the KA misfit mostly lesser than 10° , ~~except in some small areas.~~ The spatial resolution of the network is strongly influenced by the earthquake magnitude. In fact, for both M1 and M2, there are nodes (white areas where we assume the $KA = -1$ as an indeterminate value) for which the FMs cannot be calculated because ~~less than 6 stations (the minimum number of stations (at least 6) are not available~~ (Table 1). At the same time, the areas better resolved correspond to the region inside the network, ~~although with.~~ With D2 and D3 acceptable solutions are calculated for M1 and M2 earthquakes also outside the network, (Figure 5).

Looking at Figure 6, using the D3 dataset, ~~we observe that, among the FM parameters,~~ the dip angle is the best resolved compared with strike and rake angles. ~~ConsideringFor the~~ M2 and M3 focal mechanisms, the misfit of dip is very low ($< 8^\circ$), followed, in ascending order, by rake and strike that show higher values ($10^\circ < \text{misfit} < 22^\circ$). For M1 (Figure 6a-d-g), rake and strike misfits are larger than 50° , with rake worse resolved than strike. The unresolved areas correspond to the regions outside the seismic network.

The KAA and KAS maps (Figures 7 and 8) show how the network constrains the fault plane solution as a function of the epicentral location. Moreover, Figures 7d-e-f and 8d-e-f indicate that the areas with KA mean and ~~std~~ standard deviation greater than 30° and 20° , respectively, are reduced when P-wave

366 polarities and spectral level ratios data are used. On ~~the~~ contrary, only for M1 focal mechanisms, there
 367 is no improvement because the number of P-wave polarities is the same for both D2 and D3 datasets
 368 (Table 1). The worst constrained regions correspond to a belt surrounding the seismic network, with
 369 KA mean $< 30^\circ$ and KA std $< 20^\circ$ for M2 and M3 solutions. For M1, areas with high uncertainty remain
 370 outside and inside the network, specifically in the central and eastern sectors.

371 Looking at the uncertainties of FM parameters, obtained by using the D3 dataset, Figure 9 shows that
 372 the dip is the better-constrained parameter with an error $< 10^\circ$, also for M1 solutions. The rake angle
 373 shows an uncertainty ~~lowesser~~ lower than 20° for M2 and M3, while it ~~overcomes~~ higher than 50° for M1. The
 374 strike angle ~~reveals~~ has the highest uncertainty, with values greater than 50° in the eastern and southern
 375 sectors of the map for ~~all~~ any analysed magnitudes (M1, M2, and M3). Accuracy improves moving from
 376 M1 to M3 earthquakes.

377 ~~As shown in Figure 10, the~~ The accuracy of fault plane solutions, ~~in terms of KA misfit calculated by~~
 378 ~~evaluated~~ using the KA misfit and D3 dataset, is similar for FM1, FM2, and FM3, mostly with values
 379 lesser than 8° for all the magnitudes ~~M1, M2, and M3~~ (Figure 10). FM2 and FM3 show a slightly higher
 380 precision than FM1 in the area inside the seismic network (see FMM, FME, KAA, and KAS maps for FM2
 381 and FM3 in Supplementary Material). In the regions outside the network, where the azimuthal gap
 382 increases, the FMs better constrained in descending order are: FM3, FM2, and FM1. This effect should
 383 be due to the geometric relationship between the spatial distribution of the seismic stations and the
 384 orientation of the principal axes (P, T, B) that characterize the FMs.

385 Considering the effect of hypocentre depth, the results achieved for earthquakes at 5 km depth, by
 386 using the D3 dataset, are overall unchanged (Figure 11). We note that the fault plane solutions are
 387 slightly worse resolved due to a smaller number of P-wave polarities available for M2 and M3. The KA
 388 misfit ~~mainly~~ generally is ~~lesser~~ less than 10° , even though the number and the dimension of areas with
 389 misfits $> 20^\circ$ are greater than those obtained considering earthquakes at 10 km depth. Moreover, the
 390 dip angle shows a misfit lower than strike and rake angles for M1, M2, and M3; the accuracy of the
 391 retrieved FMs parameters is mainly ~~lesser~~ less than 8° , as shown in Figure 11.

Previous analyses are carried out considering ~~by using~~ data affected by 5% Gaussian error. In the last test, we simulated synthetic data adding a 30% Gaussian error. As illustrated in Figure 12, FM solutions show an overall larger misfit, in particular, the KA inside the seismic network is less than 20° . The area best resolved ($KA < 8^\circ$) is ~~considerably~~ reduced to ~~a small~~the central portion of the network. This result indicates that the accuracy of the spectral level ratio estimates is crucial: noisy waveforms with a low signal-to-noise ratio can critically affect the result of the focal mechanism inversion. So, seismic noise as well as the number of available stations, variable due to the operational conditions, strongly influence the capability of the seismic network to retrieve a fault plane solution. Using the results of our simulations, we classified the focal mechanism provided by De Matteis et al. (2016) according to a quality code based on the resolution of fault kinematics (Table 3). In fact, we assigned to focal mechanisms of the Irpinia instrumental seismicity a quality ies A, B and C for the solutions that fall into the bins relative to FM3, FM2 and FM1 kinematics, respectively. The quality A, B and C correspond to the average value of KA misfit ($FM3=2.4^\circ$, $FM2=3.1^\circ$, $FM1=4.5^\circ$) calculated for M1, M2 and M3 magnitudes using D3 dataset and considering earthquakes at 10 km depth with 5% Gaussian errors. As last analysis, we carried out a test in a more general framework, without a fixed network configuration. We explored the reliability of focal mechanism estimation as a function of the uniformity of the focal sphere coverage, defined by the number of recording seismic stations and azimuthal gap. We simulated 10400 earthquakes fixing the fault plane solution and varying: 1) the number of seismic stations (6-30), 2) the take-off angle and 3) the azimuth of each single station. For each possible number of seismic stations, we run about 400 simulations, and we randomly sampled the focal sphere varying the azimuth and take-off of the stations ~~and, so, thus~~ changing the geometrical configuration of our virtual network of each simulation. We computed the KA between the theoretical and retrieved focal mechanism (best) solutions, using only P-polarities, for each simulation. We show the results in ~~the~~ Figures 13 and S7, as 3-D histograms and 3-D scatter plot, respectively. In ~~the~~ Figures 13a, as expected, ~~we note that, as~~ the number of stations increases, while the KA and its range of variation decrease. If the number of stations is less than nine, only few solutions have $KA < 40^\circ$. Figure 13b shows that ~~the~~ most value of KA less than 30° are obtained for azimuthal gap less than about 80° . ~~Moreover,~~

~~these evidences are shown in the~~ Figure S7, ~~where~~ the relation among the KA, azimuthal gap and number of stations is ~~clearified~~ by the three-dimensional spatial point patterns as well by the projections of the data on the three coordinate planes.

CONCLUSIONS

We studied the focal mechanism reliability retrieved by the inversion of data recorded by ISNet, a local dense seismic network that monitors the Irpinia Fault System in Southern Italy. Three different datasets of seismological observables are used as input data for focal mechanism determination: a) P-wave polarities, b) P/S long-period spectral amplitude ratios, and c) joint polarities and amplitude ratios. Starting from empirical observations, we computed synthetic data for a regular grid of epicentre locations at two depths (5 and 10 km), for earthquake magnitude in the range 1.0-2.5, and for three focal mechanism solutions. Two different levels of Gaussian error (5% and 30%) are added to the data.

~~Main conclusions can be summarized as follows.~~Our results show that:

- The joint inversion of P-wave polarities and P/S spectral amplitude ratios allows retrieving accurate FM (KA misfit $< 8^\circ$) also for earthquakes with magnitude ranging between 1.0 and 2.5, at depths of 5 km and 10 km. Due to the low-energy magnitude, the number of P-wave polarities ~~is not adequate to~~cannot constrain fault plane solutions.
- The spatial resolution analysis of ISNet shows that the most accurate FM solutions are obtained for earthquakes located inside the network with —strike, dip and rake misfit $< 8^\circ$. Nevertheless, outside the network or at its borders, acceptable solutions can be calculated even if the azimuthal coverage is ~~not adequate in~~adequate (especially for M2 and M3 events). This ~~peculiarity~~ is due to the geometrical relationship between the ~~recording~~ seismic stations and the orientation of the principal axes (P, T, B).
- The geometry of the network allows to ~~well~~ resolve well fault plane solutions varying between normal and normal-strike focal mechanism with ~~mainly~~ strike, dip and rake misfit ~~lesser~~generally

less than 10° and for the magnitude range 1.5-2.5. The network resolves ~~a~~ slightly better normal-strike fault plane solution than ~~a~~ pure normal focal mechanism.

- Among the FM parameters, the dip angle shows the lowest uncertainty. Strike and rake angles ~~reveal~~have higher errors especially for M 1-1.5 earthquakes in the region outside the seismic network.
- ~~Dataset affected by Adding a~~ 30% Gaussian error ~~provide a worsening in~~worsens the accuracy of the retrieved FMs. ~~Although~~Despite the ~~high error level, the area of well resolved~~higher uncertainty, fault plane solutions (KA misfit $< 20^\circ$) ~~persists~~are still resolved in the central part of the network, especially for M2 and M3.

The methodology described in this work can be a valid tool to design ~~or to~~and test the performance of local seismic networks, aimed at monitoring natural or induced seismicity. Moreover, given a network configuration, it can be used to evaluate the reliability of FMs or to classify ~~the retrieved~~ fault plane solutions that represent a fundamental information in seismotectonic studies. Although it is a theoretical study, many earthquake scenarios with several magnitude, locations and noise conditions can be simulated to mimic the real seismicity.

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TABLES

Depth 5 km	Max Distance (km)	No. P-polarities
<i>M_L 1.0 -1.5</i>	30	1
<i>M_L 1.5 - 2.0</i>	49	1
<i>M_L 2.0 - 2.5</i>	57	4
Depth 10 km	Max Distance (km)	No. P-polarities
<i>M_L 1.0 -1.5</i>	33	1
<i>M_L 1.5 - 2.0</i>	40	5
<i>M_L 2.0 - 2.5</i>	66	6

Table 1 Maximum distance of the farthest triggered seismic station and number of P-wave polarities as function of earthquake magnitude and depth. The values, empirically derived from the ISNet bulletin, are used for the earthquake simulations.

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Figure No.	Map	Focal Mechanism Solution	Magnitude Bin	Depth	Noise Level	Dataset
4	Kagan angle misfit	FM1, FM2, FM3	M3	10 km	5%	D1
5	Kagan angle misfit	FM1	M1, M2, M3	10 km	5%	D2, D3
6	focal mechanism parameter misfit	FM1	M1, M2, M3	10 km	5%	D3
7	Kagan angle average	FM1	M1, M2, M3	10 km	5%	D2, D3
8	Kagan angle standard deviation	FM1	M1, M2, M3	10 km	5%	D2, D3
9	focal mechanism error	FM1	M1, M2, M3	10 km	5%	D3
10	Kagan angle misfit	FM1, FM2, FM3	M1, M2, M3	10 km	5%	D3
11	focal mechanism parameter misfit	FM1	M1, M2, M3	5 km	5%	D3
12	Kagan angle misfit	FM1	M1, M2, M3	10 km	30%	D3

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Table 2. Summary of the Figures 4-12 with parameters used for earthquake simulations whose results are represented as a specific map.

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P-plunge (°)	P-trend (°)	T-plunge (°)	T-trend (°)	Strike (°)	Dip (°)	Rake (°)	Quality
55	344	31	196	325	20	-40	A
51	334	36	181	320	15	-30	A
55	14	31	226	355	20	-40	A
53	205	34	49	180	15	-40	A
55	72	33	272	35	15	-50	A
51	177	32	37	290	80	-110	A
54	292	34	91	10	80	-80	A
77	146	9	7	270	55	-100	B
80	235	10	55	325	55	-90	B
76	103	2	6	110	45	-70	B
76	117	2	214	290	45	-110	B
76	82	7	199	275	40	-110	B
75	190	15	10	280	60	-90	B
75	205	15	25	295	60	-90	B
85	230	5	50	140	40	-90	B
83	146	0	53	150	45	-80	B
80	240	10	60	330	55	-90	B
81	233	5	353	270	50	-80	B
81	347	5	227	130	50	-100	B
55	93	10	198	255	45	-140	C
55	133	10	238	295	45	-140	C
48	130	2	38	275	60	-140	C
48	305	2	37	340	60	-40	C
55	202	7	102	345	60	-130	C
58	121	2	27	270	55	-130	C
58	131	2	37	280	55	-130	C
55	342	7	242	125	60	-130	C
47	138	11	36	165	50	-30	C
49	182	14	289	340	45	-150	C
58	151	2	57	300	55	-130	C
49	168	14	61	190	45	-30	C
59	308	15	64	355	65	-60	C
57	306	14	59	115	40	-140	C
57	76	14	189	245	40	-140	C
45	85	6	348	225	65	-140	C
55	22	7	282	165	60	-130	C
57	241	14	354	50	40	-140	C
55	98	7	198	135	60	-50	C
51	115	2	22	145	55	-40	C
55	147	7	47	290	60	-130	C

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Table 3. Fault plane solutions of instrumental seismicity occurred in Irpinia region in 2005-2008 and calculated by De Matteis et al., (2012). The solutions are classified according to a quality code based on the resolution of fault plane kinematics as derived in this study. The result of our simulations suggests a quality as follows: FM1=C, FM2=B, FM3=A.

FIGURES

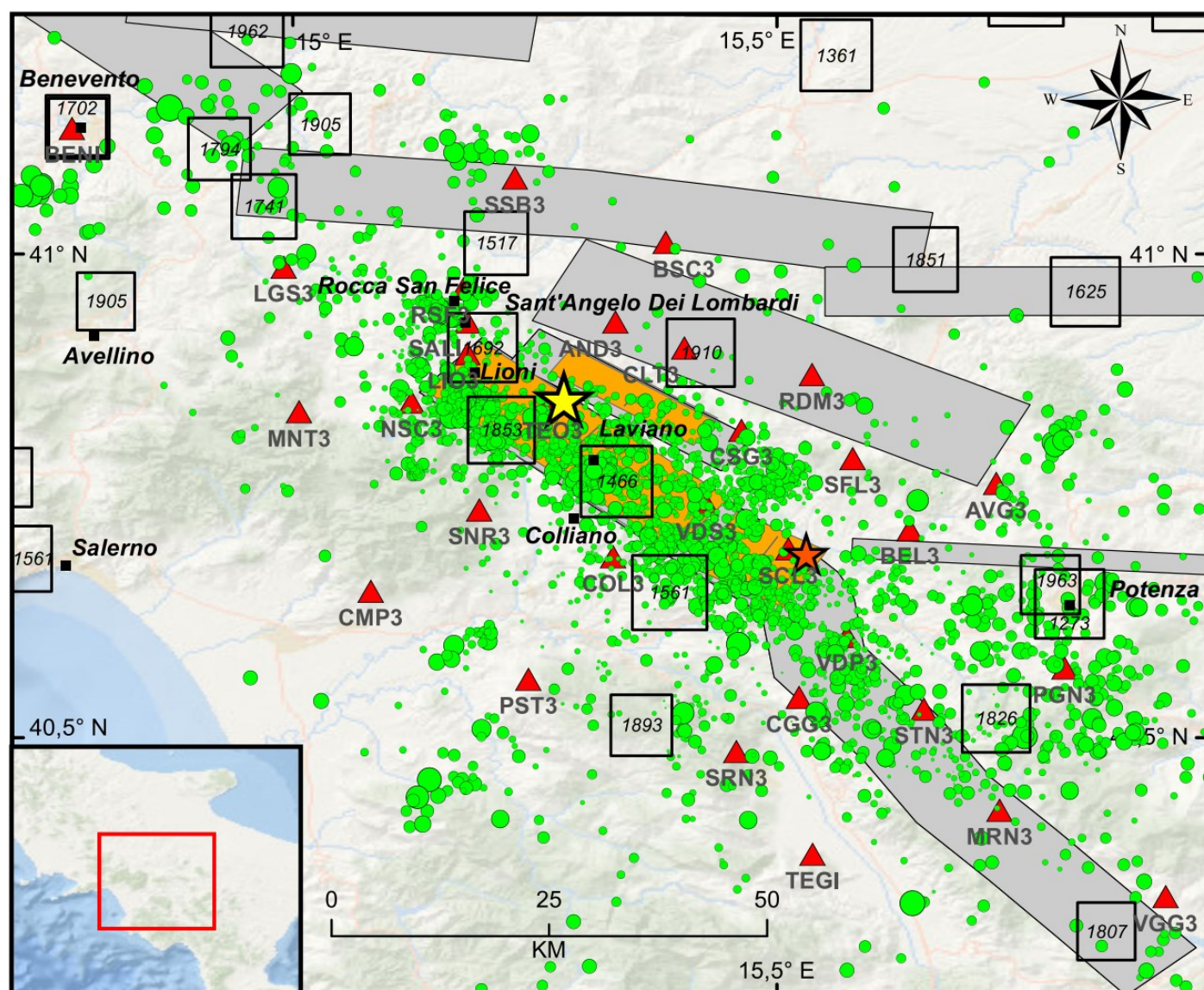


Figure 1. Epicentral map of the earthquakes (green circles) recorded by Irpinia Seismic Network (ISNet, red triangles) from 2008 to 2020 (<http://isnet-bulletin.fisica.unina.it/cgi-bin/isnet-events/isnet.cgi>). The yellow and orange stars refer to the epicentral location of the 1980, M 6.9, and of the 1996, M 4.9 earthquakes, respectively. Historical seismicity is shown with black squares ($IO \geq 6-7$ MCS). Seismogenic sources related to the Irpinia fault system are indicated by orange rectangles; potential sources for earthquakes larger than M 5.5 in surrounding areas are indicated in grey (Database of Individual Seismogenic Sources, DISS, Version 3.2.1)

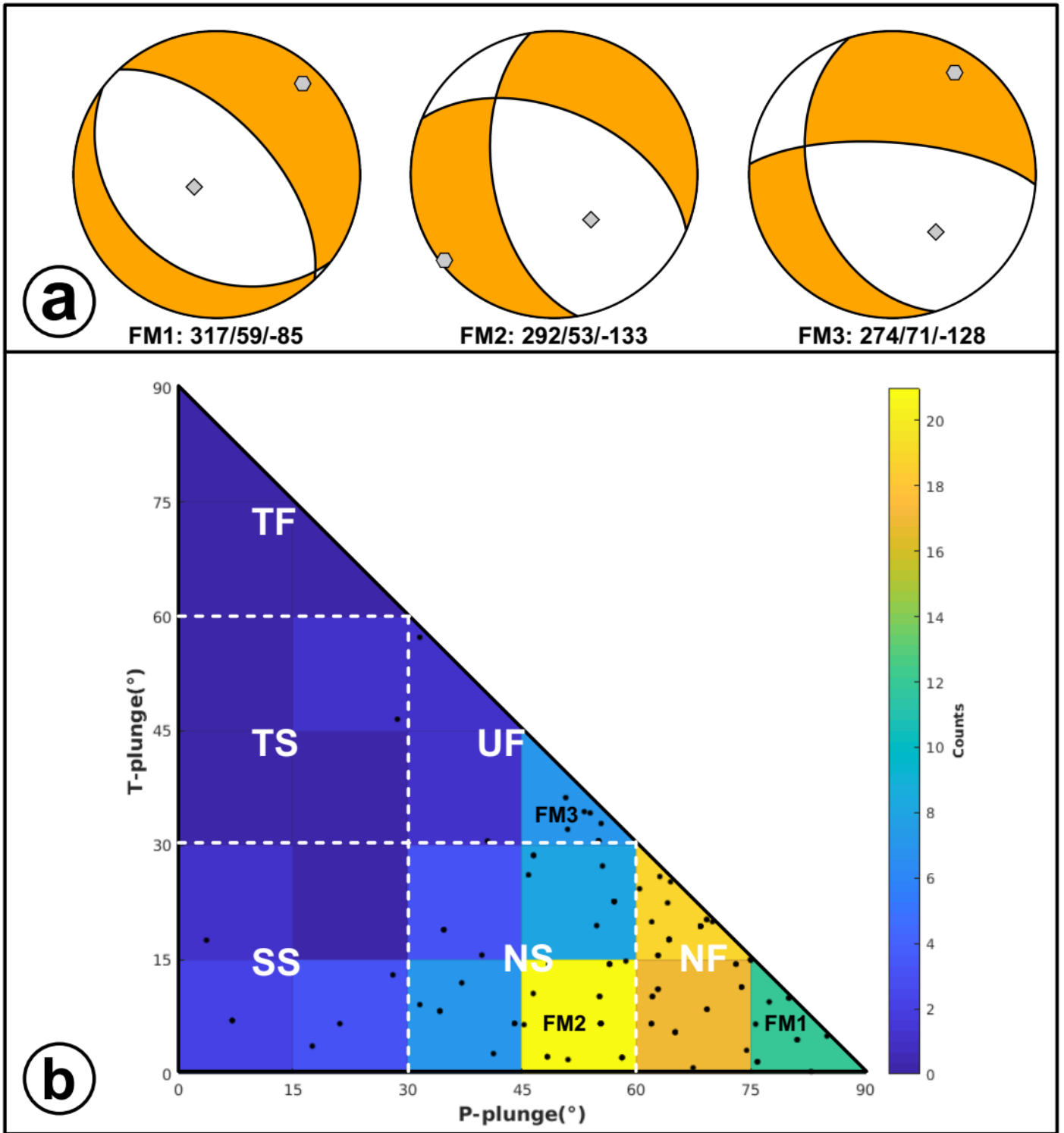


Figure 2. Fault plane solutions used for earthquake simulations. a) From left to right: 1) Ms 6.9, 23rd November 1980 (FM1; Westaway) 2) and 3) Median focal mechanism found from solutions of the 1st (FM2) and 5th (FM3) most populated bin of histogram of panel b. b) Fault plane solutions (black dots) are classified according to the plunge of P- and T-axes with the specific tectonic regimes (Legend: NF, normal fault; NS, normal-strike; SS, strike-slip; TF, thrust ; TS, thrust-strike; UF, unknown fault). The number of earthquakes (colour bar) is counted in bins of 15° × 15°.

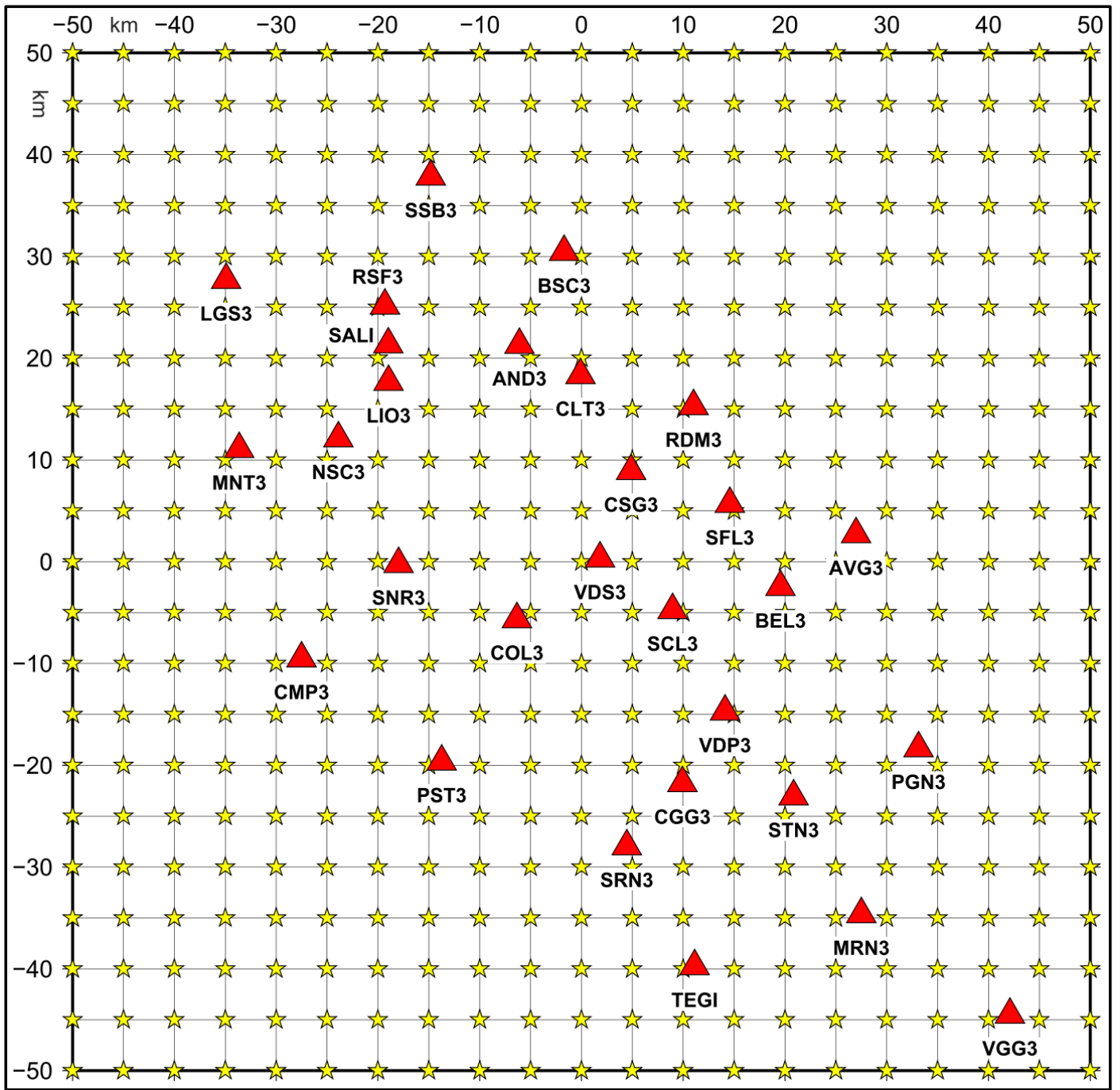


Figure 3. Regular grid of epicentres (yellow stars) used for simulating earthquakes. The area is 100x100 km² with 5 km of spacing along both horizontal coordinates. Irpinia Seismic Network (ISNet) is reported with red triangles.

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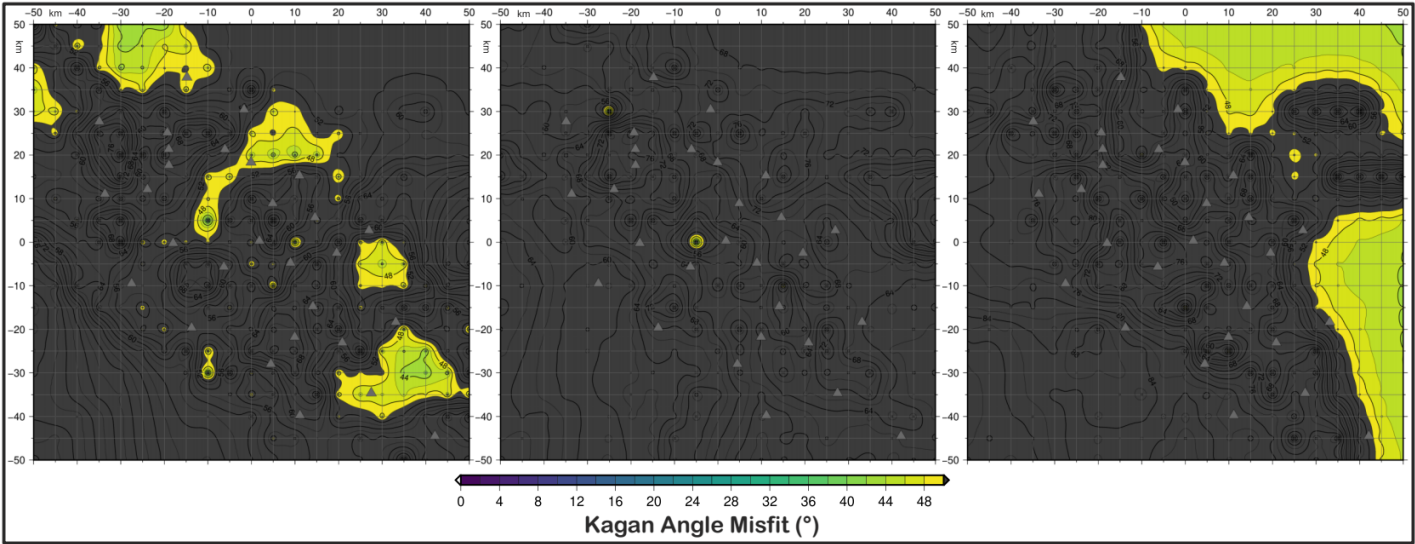


Figure 4. KAM (Kagan angle misfit) map for retrieved focal mechanisms with D1 dataset as input data and simulating earthquakes with M3 magnitude and FM1 (a), FM2 (b) and FM3 (c) theoretical fault plane solution at 10 km depth.

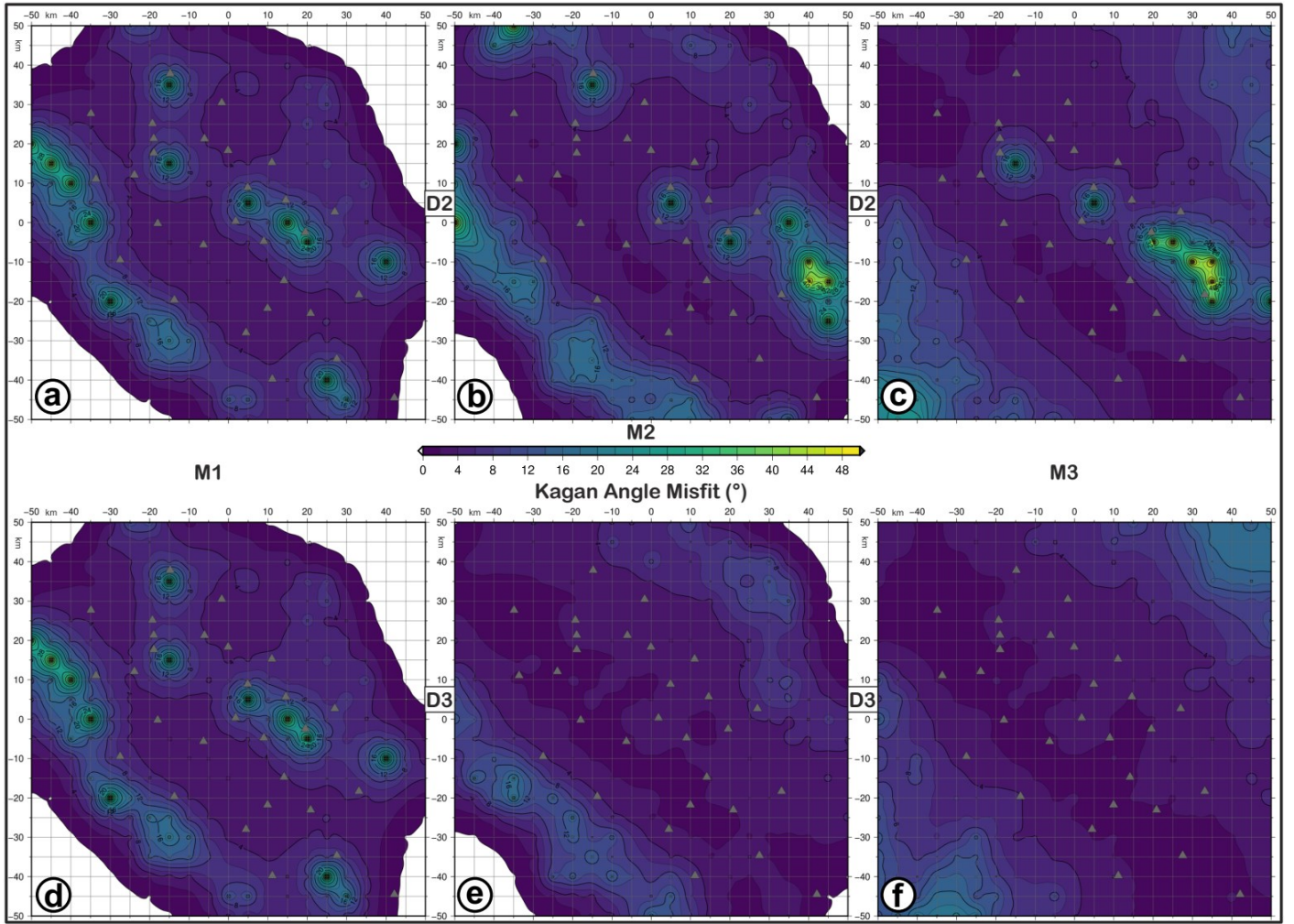


Figure 5. KAM (Kagan angle misfit) map for retrieved focal mechanisms with D2 (a, b, c) and D3 (d, e, f) datasets as input data and simulating earthquakes with M1 (a, d), M2 (b, e) and M3 (c, f) magnitudes and FM1 theoretical fault plane solution at 10 km depth. The level of Gaussian noise is set to 5%.

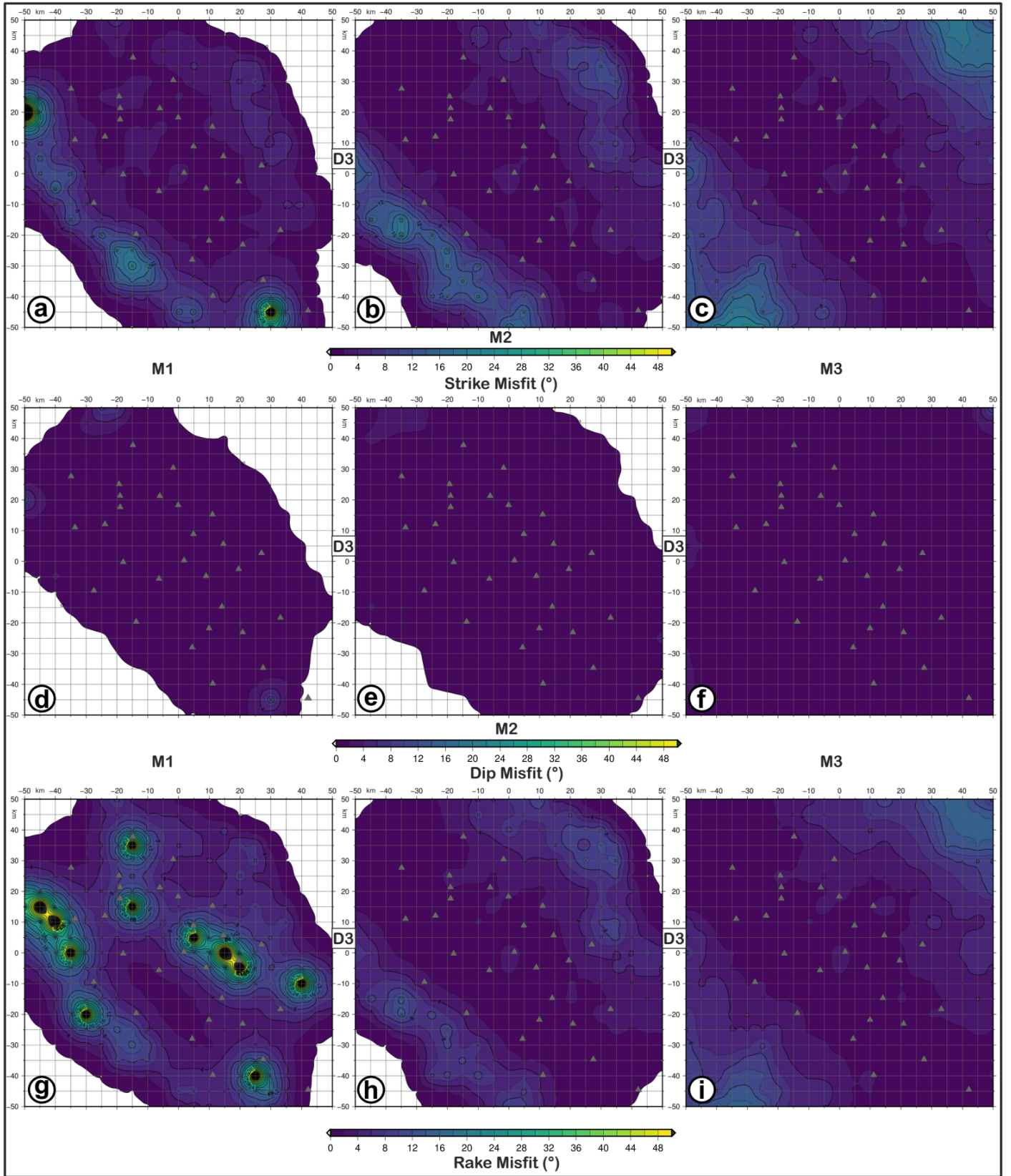


Figure 6. FMM (focal mechanism parameter misfit) maps for retrieved focal mechanisms with D3 datasets as input data and simulating earthquakes with M1 (a, d, g), M2 (b, e, h) and M3 (c, f, i) magnitudes and FM1 theoretical fault plane solution at 10 km depth. a, b, c refer to strike misfit; d, e, f refer to dip misfit; g, h, i refer to rake. The level of Gaussian noise is set to 5%.

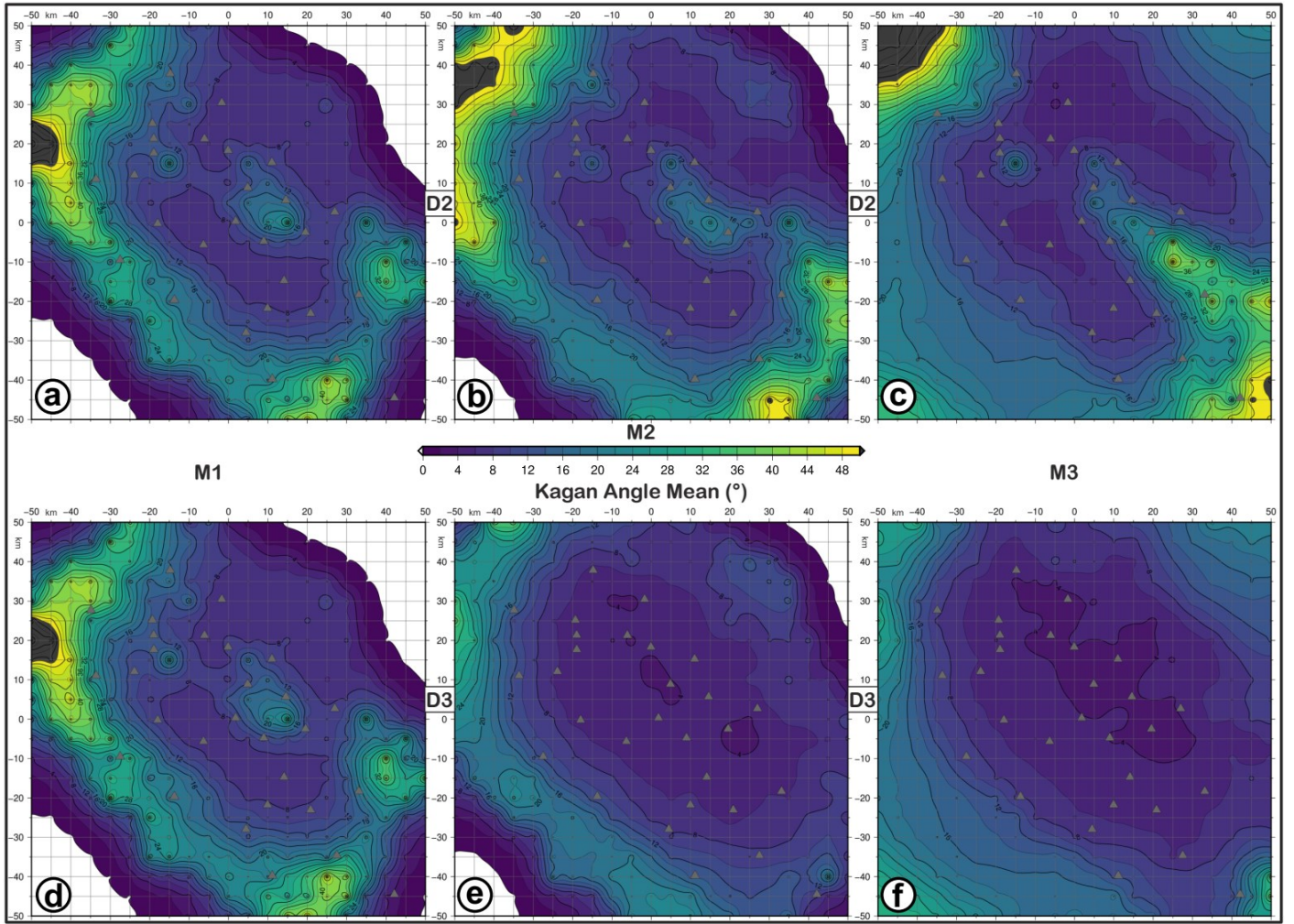


Figure 7. KAA (Kagan angle average) maps for retrieved focal mechanisms with D2 (a, b, c) and D3 (d, e, f) datasets as input data and simulating earthquakes with M1 (a, d), M2 (b, e) and M3 (c, f) magnitudes and FM1 theoretical fault plane solution at 10 km depth. The level of Gaussian noise is set to 5%.

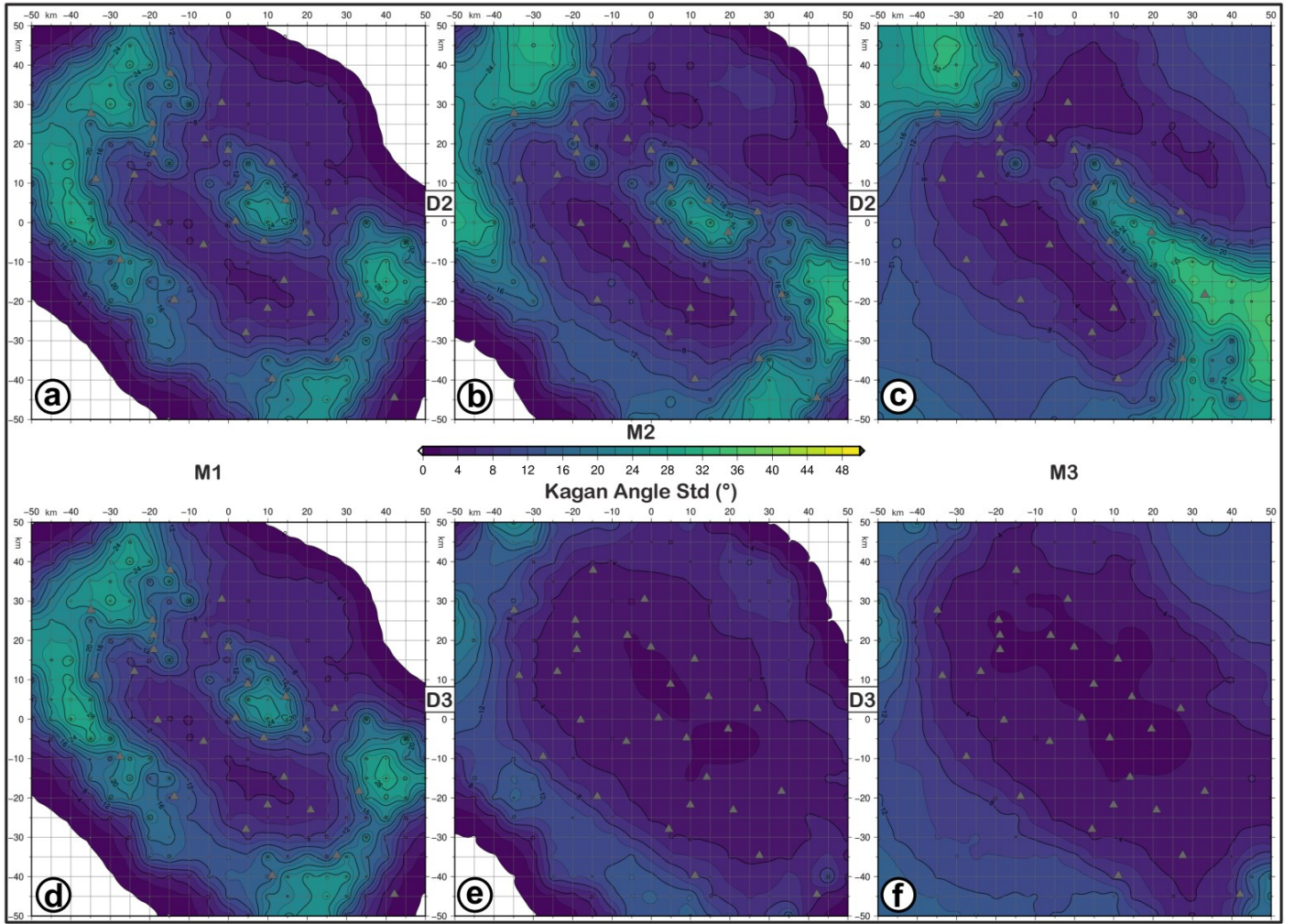


Figure 8. KAS (Kagan angle standard deviation) maps for retrieved focal mechanisms with D2 (a, b, c) and D3 (d, e, f) datasets as input data and simulating earthquakes with M1 (a, d), M2 (b, e) and M3 (c, f) magnitudes and FM1 theoretical fault plane solution at 10 km depth. The level of Gaussian noise is set to 5%.

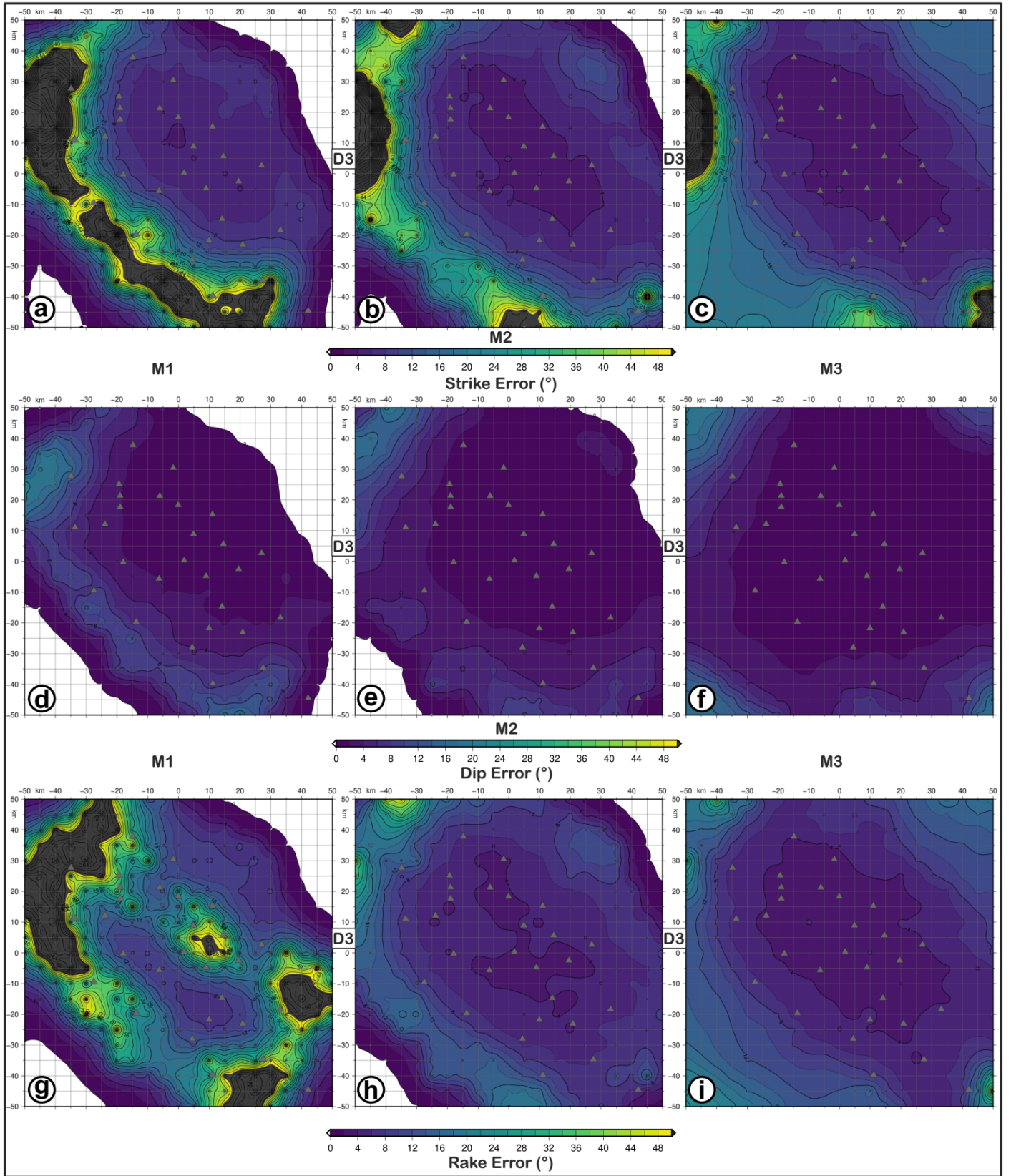


Figure 9. FME (strike, dip and rake error) maps for retrieved focal mechanisms with D3 datasets as input data and simulating earthquakes with M1 (a, d, g), M2 (b, e, h) and M3 (c, f, i) magnitudes and FM1 theoretical fault plane solution at 10 km depth. a, b, c refer to strike error; d, e, f refer to dip error; g, h, i refer to rake error. The level of Gaussian noise is set to 5%.

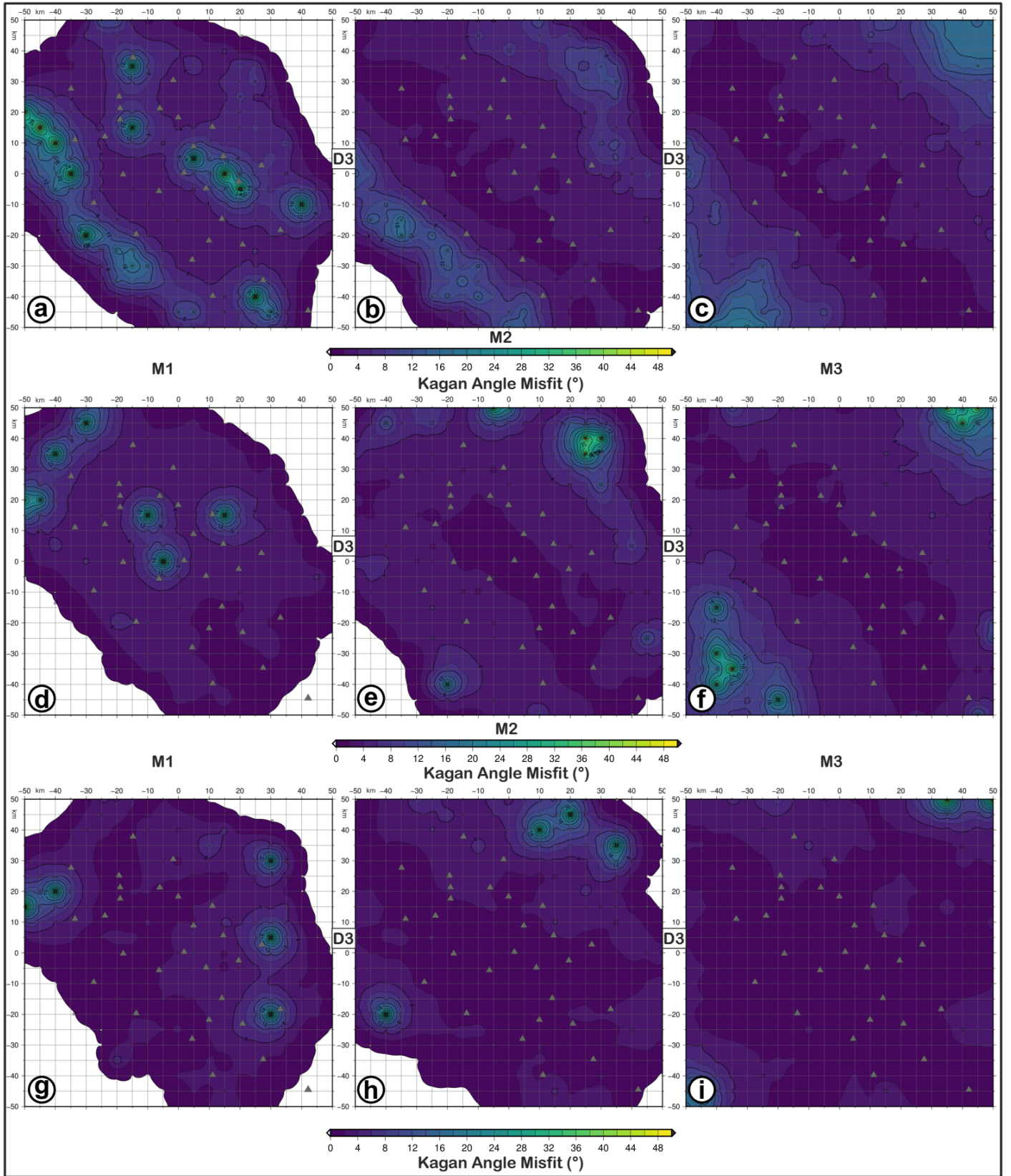


Figure 10. KAM (Kagan angle misfit) maps for retrieved focal mechanisms with D3 datasets as input data and simulating earthquakes with M1 (a, d, g), M2 (b, e, h) and M3 (c, f, i) magnitudes and FM1 (a, b, c), FM2 (d, e, f) and FM3 (g, h, i) theoretical fault plane solution at 10 km depth. The level of Gaussian noise is set to 5%.

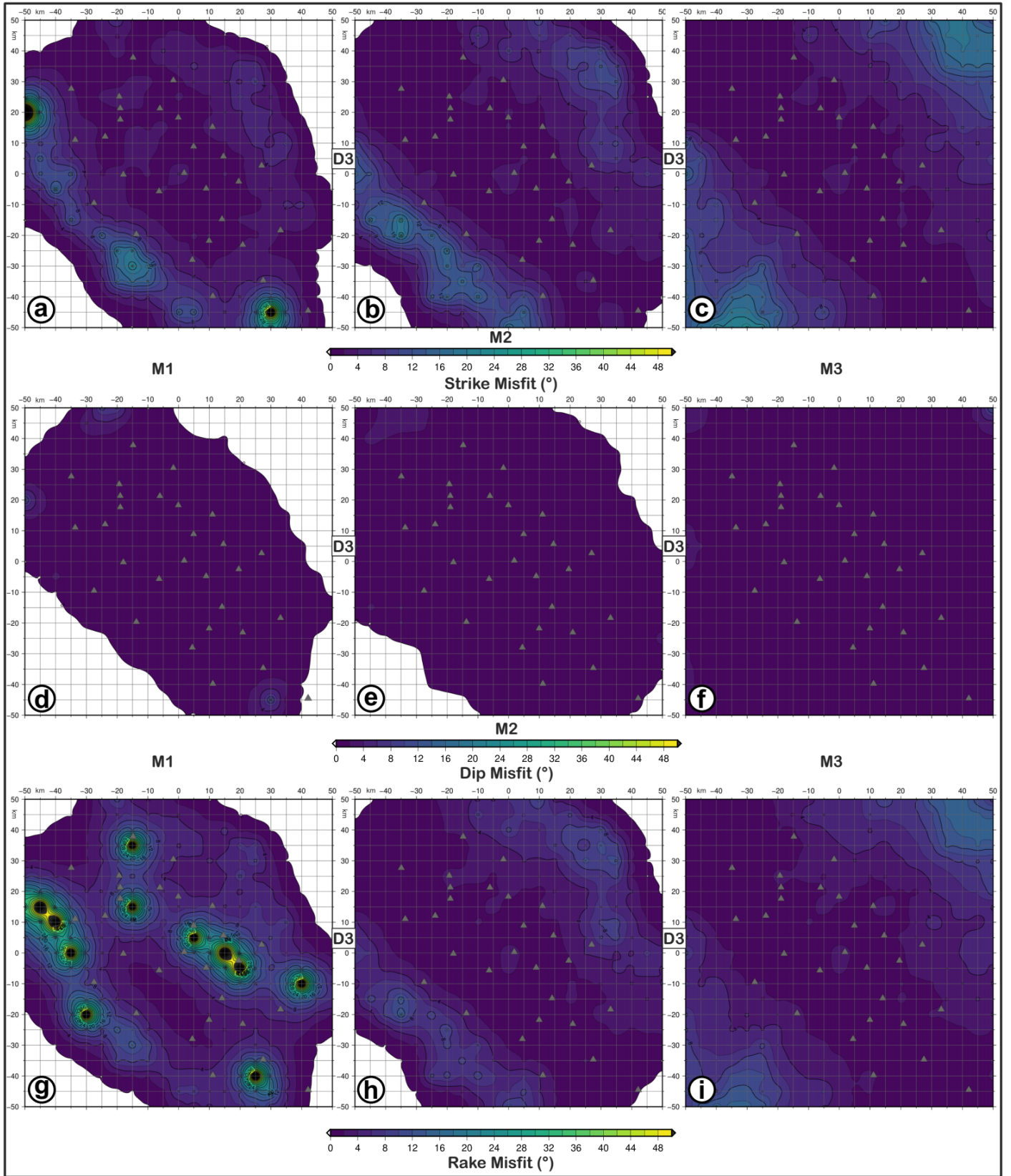
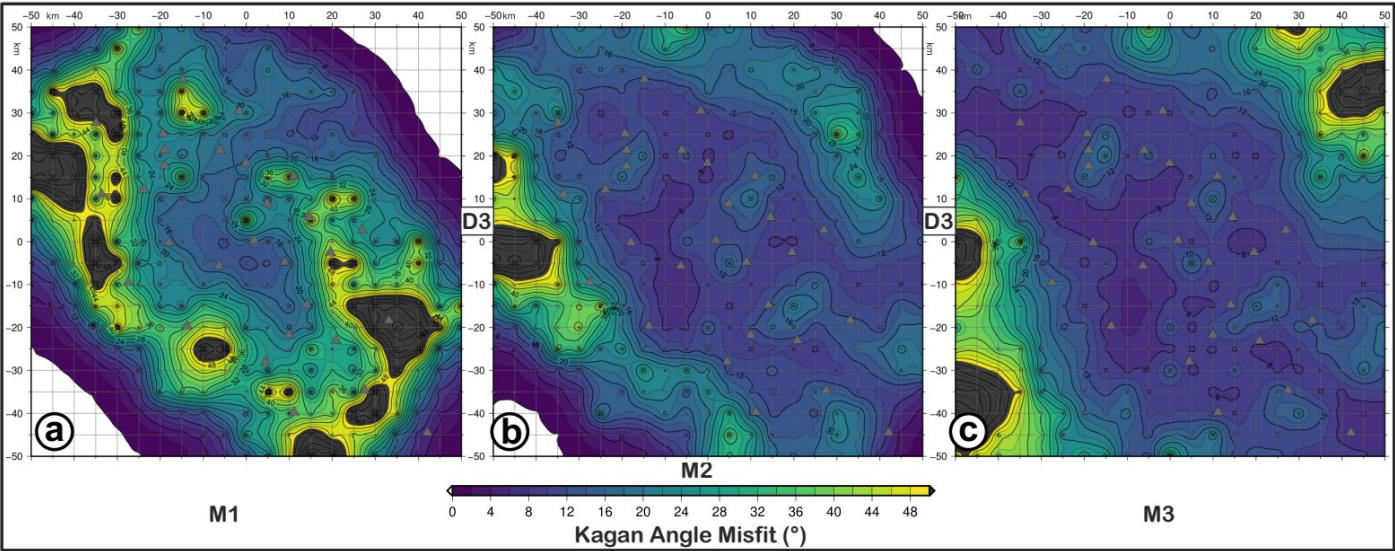


Figure 11. FMM (focal mechanism parameter misfit) maps for retrieved focal mechanisms with D3 datasets as input data and simulating earthquakes with M1 (a, d, g), M2 (b, e, h) and M3 (c, f, i) magnitudes and FM1 theoretical fault plane solution at 5 km depth. a, b, c refer to strike misfit; d, e, f refer to dip misfit; g, h, i refer to rake. The level of Gaussian noise is set to 5%.

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Figure 12. KAM (Kagan angle misfit) map for retrieved focal mechanisms with D3 (a, b, c) datasets as input data and simulating earthquakes with M1 (a), M2 (b) and M3 (c) magnitudes and FM1 theoretical fault plane solution at 10 km depth. The level of Gaussian noise is set to 30%.

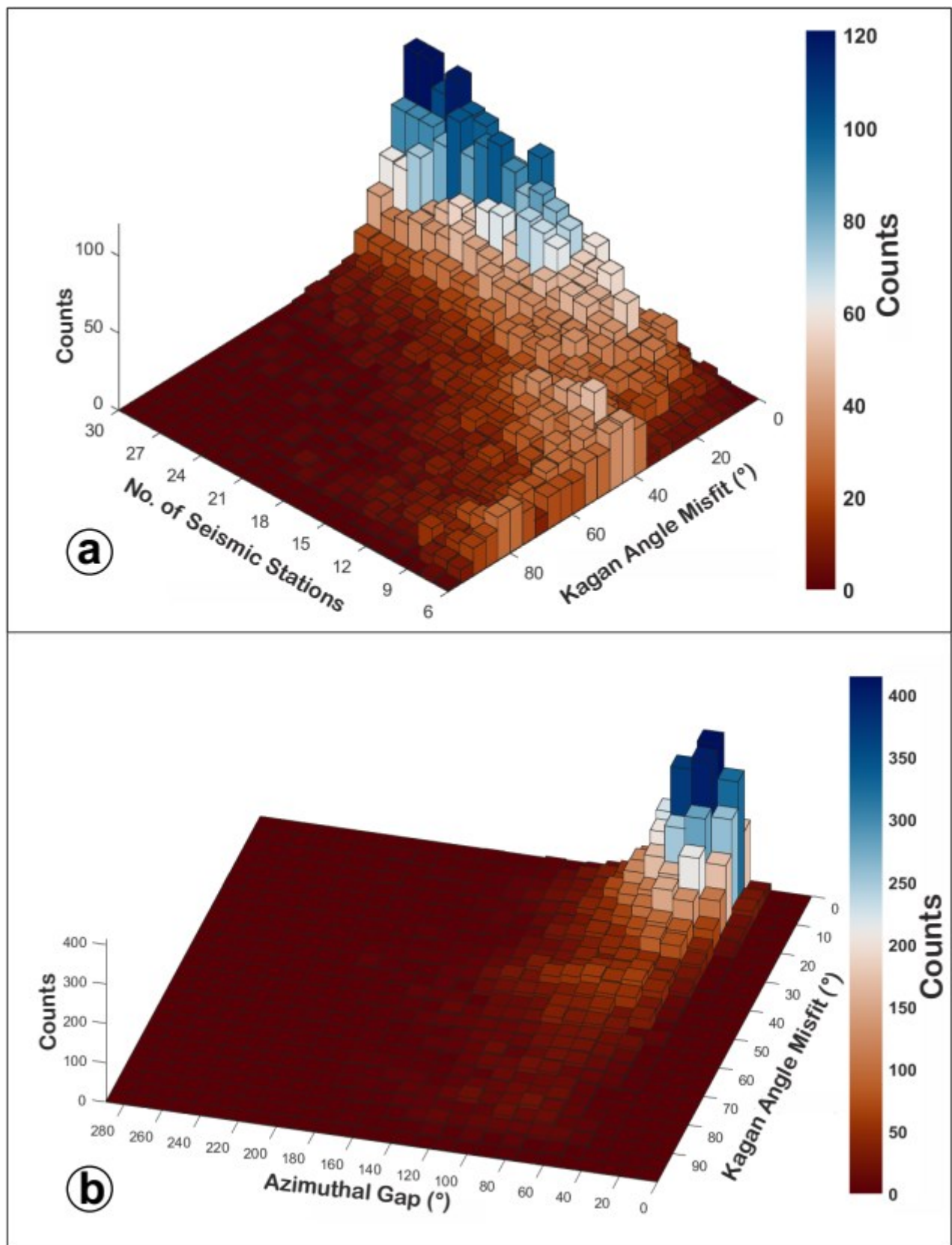


Figure 13. 3D-histograms of the test results in terms of number of stations (a), azimuthal gap (b) and KA misfit. The simulations were carried out with a free network configuration.